Astronomy

Allison Barta Chad Morris

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Day 1: Astronomy The Original Science

- Identify the units of a calendar
- Describe 2 early ideas about the structure of the universe
- Describe the contributions of Brahe, Kepler, Galileo, Newton & Hubble to modern astronomy
- Astronomy the Original Science Worksheet

Day 2: Telescopes

- Compare refracting telescopes with reflecting telescopes
- Explain how the atmosphere limits astronomical observations And explain how astronomers over come these limitations.
- Telescopes worksheet
- Astronomy Word Search

Day 3: Constellations

- Define Constellations and Galaxies
- Discuss the origins of Constellations
- Organize the class into four groups for sky study And a Constellation Tour Project
- Spectacular Night Sky Work sheet

Day 4: Mapping the Sky

- Explain how constellations are used to organize the night sky
- Describe how the altitude of a star is measured.
- Explain how the celestial sphere is used to describe the location of objects in the sky.
- All students create an astrolabe
- Mapping the sky worksheet

Day 5: Mapping the Sky

- Compare results of Home sky Astrolabe lab
- Compare size and scale in the universe.
- Explain how red shift indicates that the universe is expanding.

Day 6: Review

• Complete Review - Reinforcement worksheet & key topics

Day 7: Test

Standards For Earth And Space Science Grade 8

Strand III Earth and Space Science

Sub-Strand C. The Universe

Standard The student will compare objects in the solar system and explain their interactions with the Earth.

Benchmarks

- 1. The student will recognize that the sun is the principal energy source for the solar system and that this energy is transferred in the form of radiation.
- 2. The student will explain how the combination of the Earth's tilted axis and revolution around the sun causes the progression of the seasons and weather patterns.
- 3. The student will compare and contrast the planets, taking into Account their compositions, mass and distance from the sun and Recognize the conditions that allowed life to flourish on Earth.
- 4. The student will use the predictability of the motions of the Earth and sun to explain the length of day, length of year, phases of the moon, eclipses, tides and shadows.

Strand III Earth and Space Science

Sub-Strand C. The Universe

Standard The student will describe the composition and structure of the Universe

Benchmarks

- 1. The student will recognize that the universe consists of many of billions of galaxies, each containing many billions of stars and that there are vast distances that separate these galaxies and stars from one another.
- 2 The student will recognize that the sun is a medium-sized star and is the closest star to Earth. It is the central and largest body in the solar system and is one of billions of stars in the Milky Way Galaxy.

21 BRIGHTEST STARS IN THE SKY

STAR		CONSTELLATION	MAGNITUDE	
1.	Siruis	Canis Major	-1.46	
2.	Canopus	Carina	-0.72	
3.	Rigel Kentauras	Centaurus	-0.27 (dbl)	
4.	Arcturus	Bootes	- 0.04	
5.	Vega	Lyra the Harp	+ 0.03	
6.	Capella	Auriga	+0.08	
7.	Rigel	Orion	+0.12 (dbl)	
8.	Procyon	Canis Major	+ 0.38	
9.	Achernar	Eridanus	+ 0.46	
10.	Betelgeuse	Orion	+ 0.50	
11.	Hadar	Centaurus	+ 0.61	
12.	Altair	Auquila the Eagle	+ 0.77	
13.	Aldebaran	Taurus the Bull	+ 0.85 (var)	
14.	Acrux	Crux	+ 0.87	
15.	Antares	Scorpius	+ 0.96	
16.	Spica	Virgo	+0.98	
17.	Pollux	Gemini Twins	+ 1.14	
18.	Fomalhant	Piscis	+ 1.16	
19.	Deneb	Cygnus the Swan	+ 1.25	
20.	Mimosa	Crux	+ 1.25	
21.	Regulus	Leo the Lion	+ 1.35	

(dbl) = Double star; combined magnitude of components is given

(var) = Variable star; Brightest magnitude is given

Ptolemy was a Greek astronomer who lived between 85-165 A.D. He put together his own ideas with those of Aristotle and Hipparchus and formed the geocentric theory. This theory states that the Earth was at the center of the universe and all other heavenly bodies circled it, a model which held for 1400 years until the time of Copernicus.

Ptolemy is also famous for his work in geography. He was the first person to use longitude and latitude lines to identify places on the face of the Earth.

Nicholas Copernicus was a Polish astronomer who lived between 1473-1543. Before his time, people believed in the Ptolemaic model of the solar system, which maintained that the Earth was the center of the universe.

Copernicus changed this belief when he introduced the heliocentric model, centered around the sun. He claimed that all the planets, including Earth, moved in orbits around the sun, and showed how this new system could accurately calculate the positions of the planets.

Tycho Brahe was a Danish astronomer who lived between 1546-1601. For over twenty years, he made very accurate observations of the night sky, all without the aid of a telescope, which had not yet been invented. Tycho also built the world's first observatory and kept a star catalogue with over 1000 stars.

Tycho's records were used by Johan Kepler to describe the orbits of planets around the sun and disprove the Ptolemaic theory.

Johan Kepler was a German astronomer who lived between 1571-1630. He introduced three important laws of planetary motion and helped the Copernican model of the solar system gain general acceptance.

Kepler inherited Tycho Brahe's observational data on Mars following Brahe's death and showed, mathematically, that Mars followed an elliptical orbit. This new revelation contradicted the age old belief that heavenly bodies all moved in perfect circles. During his life, Kepler also cast horoscopes and wrote science fiction novels.

Galileo Galilei was an Italian astronomer and physicist who lived between 1564-1642. He challenged Aristotle's proposition that heavenly bodies were divine and therefore perfect and blemish-free.

Galileo was the first person to use a telescope to look at the heavens. He discovered sunspots, and craters and peaks in the moon. Galileo's work offended the Roman Catholic Church and he was sentenced to house arrest for the later years of his life. Today, he is remembered as a martyr for scientific truth.

Isaac Newton was an English scientist and mathematician who lived between 1642-1727. He had one of the most brilliant minds the world has ever known. Legend has it that seeing an apple fall gave Newton the idea that gravity, the force which keeps us bound to the Earth, also controls the motion of planets and stars.

Newton's contributions to science include the universal law of gravitation, the development of a whole new field in mathematics called calculus, and his famous three laws of motion

Albert Einstein was a German physicist who lived between 1879-1955. Probably the most well-known scientist of the twentieth century, Einstein came up with many original theories and invented modern physics.

He is most famous for his theory of relativity, which makes bold statements about the nature of light and also shows the relationship between mass and energy. Einstein's accurate predictions on the link between gravity with space and time also made him a celebrity.

Edwin Hubble was an American astronomer who lived between 1889-1953. His observations of galaxies helped him develop the idea of an expanding universe, which forms the basis of modern cosmology, the study of the origin of the universe. He also discovered a relationship between a galaxy's speed and its distance.

Hubble's studies were interrupted by service in both World Wars. The Hubble space telescope, currently on an observation project in space, bears his name.

History of the Telescope

Phoenicians cooking on sand discovered glass around 3500 BCE, but it took about 5,000 years more for glass to be shaped into a lens for the first telescope. A spectacle maker probably assembled the first telescope. Hans Lippershey (c1570-c1619) of Holland is often credited with the invention, but he almost certainly was not the first to make one. Lippershey was, however, the first to make the new device widely known.

The telescope was introduced to astronomy in 1609 by the great Italian scientist Galileo Galilei, who became the first man to see the craters of the moon, and who went on to discover sunspots, the four large moons of Jupiter, and the rings of Saturn. Galileo's telescope was similar to a pair of opera glasses in that it used an arrangement of glass lenses to magnify objects. This arrangement provided limited magnification--up to 30 times for Galileo--and a narrow field of view; Galileo could see no more than a quarter of the moon's face without repositioning his telescope. In 1704, Sir Isaac Newton announced a new concept in telescope design whereby instead of glass lenses, a curved mirror was used to gather in light and reflect it back to a point of focus. This reflecting mirror acts like a light-collecting bucket: the bigger the bucket, the more light it can collect. The reflector telescope that Newton designed opened the door to magnifying objects millions of times--far beyond what could ever be obtained with a lens.

Newton's fundamental principle of using a single curved mirror to gather in light remained the same. The major change that took place was the growth in the size of the reflecting mirror, from the 6-inch mirror used by Newton to the 6-meter (236 inches in diameter) mirror of the Special Astrophysical Observatory in Russia, which opened in 1974.

The idea of a segmented mirror dated back to the 19th century, but experiments with it had been few and small, and many astronomers doubted its viability. It remained for the Keck Telescope to push the technology forward and bring into reality this innovative design.

A binocular is a optical instrument for providing a magnified view of distant objects, consisting of two similar telescopes, one for each eye, mounted on a single frame. The first binocular telescope was invented by J. P. Lemiere in 1825.

Early telescopes, like Galileo's, were designed on the principles of refraction, or redirecting light by bending its rays as it passes from one medium, like air, into and out of another medium, like glass. A convex objective lens, located at one end of the telescope's tubular body, gathered as much light as possible from a distant object. The larger the lens, the more light it could gather. As light rays passed through the objective lens, its curvature directed the rays to converge and form an image of the viewed object near the other end of the tube, at a place called the focal plane. A concave eyepiece then magnified the tiny image for the viewer to see.

The lenses in Galileo's telescope had a greenish hue from iron contained in the glass and were filled with tiny bubbles that distorted the image. In addition, as explained in more detail below, a colored halo surrounded the images seen in the small field of view. Yet another shortcoming of Galileo's telescope was that magnification could only be improved by increasing the focal length, which meant focusing the light farther behind the objective

lens. Tubes were made longer and lenses larger, but there were practical limits to this design: a glass lens large enough to capture the sparse rays from more distant stellar objects would sag under its own weight.

Isaac Newton introduced a new concept in telescope design in which mirrors replaced glass lenses. In a reflecting telescope, a curved concave mirror at the base of the tube gathered light and reflected it to a point of focus situated about halfway back along the tube. There, a second mirror -- flat and angled -- reflected light to an eyepiece located on the side of the tube.

Newton's design held some distinct advantages over Galileo's. For example, lenses naturally cause different wavelengths of light to separate as they pass through them. Light of different colors focuses at different points, causing distortions in the color of the objects under observation, especially around their edges. Mirrors, by contrast, do not separate the colors in this way.

Telescope

Sir Isaac Newton once wrote, in what must be one of the most oft-quoted lines in the history of science, "If I have seen further it is by standing upon the shoulders of giants." Well, any of us who has ever looked through a telescope could say the same, for the two principal types in use today were developed by giants: the refractor by Galileo and the reflector by Newton. Those two scientific geniuses, as Sir Isaac himself acknowledged, built upon the work of others, and how they came to design their inventions offers a compelling glance at the workings both of the telescope and of the scientific process itself. Making the far near

The invention that set the stage for the telescope was the eyeglass, which appeared in the mid-13th century. The first spectacles, designed to correct farsightedness, bore glass magnifiers that were biconvex, meaning they curved outward on both sides. (Each resembled a lentil, or *lens* in Latin.) Mostly used to help older people read, these lenses focused on objects between 12 and 20 inches away from the eye.

Nearsightedness, a more common affliction, proved more difficult to correct. It required biconcave lenses—those curving inward on each surface—that had to bring objects into focus at the specific distance at which one's eyesight failed. The poorer one's vision, the greater the distance the lenses needed to provide focus.

In 1608, someone in Europe—it's not clear who—figured out that if you placed a lens for the farsighted about 12 to 14 inches away from a lens for the nearsighted, and then peered through the latter lens, distant objects would miraculously appear as if close by. (Oh, to have seen that pioneer's expression upon first realizing this!) Place those lenses in a tube and voilà, you have a spyglass.

Within months, Galileo had not only learned of the new device but was well on his way to improving its design. In his workshop in Padua, Italy, he discovered that plano-convex and plano-concave lenses worked best—that is, lenses with a plane on one side and curved surfaces on the other. Then, drawing on his skills as a professor of mathematics at the University of Padua, he determined the mathematical relationship that governed the instrument's ability to magnify. A spyglass with a plano-convex lens that focuses at 12 inches and a plano-concave lens that focuses at four inches, he found, magnifies images three times (12 divided by four). Galileo played with this formula until, by the late fall of 1609, he'd made a spyglass that could magnify what is seen by 20 times. No other spyglass maker could match that.

That fall, Galileo also did what apparently no one else had ever done with a spyglass before: train the instrument on the heavens. In short order he began making astronomical findings that would shake our understanding of our place in the universe to its foundations. Among them was the discovery of four moons orbiting Jupiter. To Galileo, the moons proved that not everything in space circled the Earth, and therefore our planet was not the absolute center of the universe, as the Church maintained the Bible had it.

Building a better refractor

The spyglass-turned-telescope had limitations, some of which Galileo was able to design around. To reduce distortions such as elongations and blurriness caused by the curvature of the "objective" lens—the convex lens at the far end of the telescope—Galileo ground a lens larger than he needed, for example. He then placed cardboard around the edges of the lens so that light entered that portion of the lens where curvature-related distortions were

least apparent.

Other problems Galileo would have to leave to others. One concerned magnification. In striving to make images he saw through his telescope ever larger, Galileo found that his field of view became ever smaller. He reached a point of diminishing returns beyond which enlarging the image made what was seen through the telescope too small to be of practical use. That point, he found, was achieved when he succeeded in magnifying the image 20 or 30 times.

Galileo's contemporary Johannes Kepler, a German mathematician, discovered a way to get beyond the magnification ceiling. Instead of a concave lens near the eye, Kepler used a convex lens. The result was that the image magnified by the convex objective lens was further magnified by the now-convex eyepiece lens. The only problem was that the resulting image was upside down. Some astronomers added a third convex lens to right the image, but this added unwanted distortions. Eventually, most astronomers simply accepted the inverted view. After all, did it really matter that they were seeing planets and stars upside down?

Kepler's advance created a new challenge, however. Because his telescope offered a much greater field of view, astronomers had plenty of room to improve magnification. One way to do that, they knew, was to make the focal length of the eyepiece shorter. This, however, increased the curvature of the lens, which exacerbated the distracting fringes of color that surrounded images. The other way was to reduce the curvature of the objective lens. This results in fewer color problems, because the less a lens is curved the less it acts like a prism. But this made the focal length of the objective lens—and thus the instrument itself—longer. As Richard Panek writes in his book Seeing and Believing: How the Telescope Opened Our Eyes and Minds to the Heavens, things got a little out of hand: Six to eight feet—that was the length of a good astronomical telescope in 1645. Five years later it was 10 to 15 feet. Ten years after that, 25 feet. Ten years after that, 40 to 50 feet. By 1673, [Polish astronomer] Johannes Hevelius had constructed a telescope 150 feet long on the shores of the Baltic Sea....One astronomer cheered the coming day when aerial telescopes would have a focus of 1,000 feet and human spectators could marvel at the antics of the animals on the Moon.

Alas, extremely long telescopes proved to be impracticable. Most useful were those no longer than 30 or 40 feet—as Panek says, "long but not *that* long...."

Goodbye to color

Designers of shorter telescopes still faced the problem of chromatic aberrations, though—that is, until Isaac Newton solved it. In 1672, Newton published a seminal paper on light and color, in which he showed that white light is a mixture of all colors of the spectrum. When white light passes through a curved lens it breaks down into its various component colors, each of which comes into focus at a different point on the optical axis—hence the color fringes seen in refractors.

Other scientists had speculated that using mirrors rather than lenses might correct this problem, but Newton was the first to put such thoughts into practice. He cast a two-inch metal mirror and ground it so that it had a spherical curvature. He placed the mirror at one end of a tube. Light coming in the other end reflected off the mirror back toward the opening. The reflected light struck a secondary mirror Newton had affixed inside the tube at a 45-degree angle. That secondary mirror, in turn, bounced the light (and thus the image

seen) into a convex eyepiece lens built into the side of the tube.

This was the first working reflecting telescope. Other designers proved unable to grind mirrors with a regular curvature, so the reflector remained largely a curiosity until the mid-18th century, when it finally began to come into its own.

Standing on shoulders

Both types of telescopes—Galileo's refractor and Newton's reflector—underwent improvements in the coming centuries. Designers of refracting telescopes figured out, for one, that by using two different kinds of glass together in the objective lens, they could cause the red and blue ends of the spectrum to converge towards a single focal point, thereby solving the color-fringe problem. Despite myriad embellishments, however, most optical telescopes in use in the 21st century derive from the two types developed in the 17th century by Galileo and Newton, on whose shoulders all astronomers, both amateur and professional, stand today

Making an Astrolabe

In this activity you will make an astronomical device called an astrolabe.

Ancient astronomers use astrolabes to measure the location of stars in the sky and to help them navigate. For around two thousand years now, people have been using astrolabes for navigation. The Greeks are credited with inventing the instrument. Moslems were using astrolabes by the eighth century, and in 1381 the English author Geoffrey Chaucer wrote a Treatise on the Astrolabe. But the astrolabe's heyday was in the navigational boom of the 1400 and 1500s. Now the Global Positioning System has eliminated the need for navigation by the stars, except perhaps in emergency situations, but it is still an interesting skill to know.

You will use the astrolabe to measure the angle or altitude of an object.

Procedure:

- 1. Tie one end of a piece of string that is 15cm long to the center of the straight edge of a protractor. Attach a paper clip or washer to the other end of the string.
- 2. Tape a straw lengthwise along the straight edge of the protractor.

Your astrolabe is complete!

- 3. Look through the straw at a distant object. The curve of the astrolabe should point toward the ground.
- 4. Hold the astrolabe still and carefully pinch the string to the protractor. Count the number of degrees between the string and 90 degrees. Or subtract the number you have from 90 and that is the altitude.

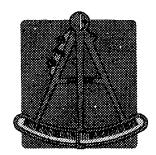
If outside at night you can:

Locate Polaris, then use your astrolabe to find its altitude. Sight the star through the straw and note what degree the string lines up at on the protractor. (Read the inner set of numbers, from 0-90 degrees.) This number is the zenith angle. To find the altitude angle, subtract the zenith angle from 90 degrees. This number will be the same as, or very close to, the latitude at your sighting location.

Analysis:

- 1. What is the altitude of the object? How would the altitude change if you moved closer to the object?
- 2. Explain how you would use an astrolabe to find the altitude of a star. What are the advantage and disadvantages of this method of measurement?

Astrolabe Activity	
LP	Date



Objectives:

- To learn how to use an astrolabe
- To define the term altitude
- To understand the relationship between angle and altitude
- To understand the relationship between the altitude of Polaris and latitude

Materials:

- 1. Class set of astrolabes made by teacher
- 2. Or, each student can make and keep their astrolabe prior to this activity.

Procedure:

- 1. Using your Astrolabe, find the angle in degrees (altitude) of the following objects in our classroom. Then, chose 4 objects of your choice.
- 2. When making your measurements, stand as far away from the object as you can.
- 3. Record your measurements in the table provided.

Object being measured	Angle in degrees (Altitude)
Middle of Clock	
Pulley of the planetarium	
Fire alarm	
Kermit's left eye	
Top of front wall	
Top of the T.V.	
Your choice	



•	is and Results: ine the word altitude:			
2. Whe	en measuring inside the classro	oom, how does d	istance affect the altitude?	
	en measuring celestial bodies f re standing in your backyard, s	-	-	
	e latitude of Englewood, NJ is a at is the altitude of Polaris at		tude of Polaris is 40 degrees	i .
	City	Latitude	Altitude of Polaris	
	London, England	51°N		
	Panama City, Panama	9°N		
	Hammerfest, Norway	70°N		
	Cairo, Egypt	30°N		
	San Juan, Puerto Rico	18°N		
	North Pole	90°N		
	Sydney, Australia	33°S		



V	Class	Data
Name	Class	Date
Skills Worksheet		
Directed Reading A		

Dire

Section: Astronomy: The Original Science

- 1. In what way did people in ancient cultures mark the passage of time?
- 2. What science did the study of the night sky eventually become?

OUR MODERN CALENDAR

Match the correct definition with the correct term. Write the letter in the space provided.

- 3. roughly the amount of time required for a. day the moon to orbit once around the Earth **b.** month
- c. year **4.** the time required for the Earth to orbit once around the sun
- 5. the time required for the Earth to rotate once on its axis

WHO'S WHO OF EARLY ASTRONOMY

- 6. Most early astronomers thought that the universe consisted of
 - a. the sun and the Earth.
 - **b.** the sun and the planets.
 - c. the sun, the moon, and the planets.
 - **d.** the sun, the moon, and the Earth.
- 7. What was Ptolemy's theory of the universe?
 - a. The Earth was at the center of the universe, and the sun, moon, and other planets revolved around it.
 - b. The sun was at the center of the universe, and the Earth and the other planets revolved around it.
 - c. The sun and the moon revolved around the Earth, but the other planets revolved around the sun.
 - **d.** The planets revolved around the sun in elliptical orbits.



Name	Class Date
Directe	ed Reading A continued
1	 B. How long did Ptolemy's Earth-centered theory remain the popular theory for the structure of the universe? a. about 100 years b. about 500 years c. over 1,500 years d. over 5,000 years
!	9. Why was Ptolemy's theory of the universe helpful even though it was incorrect?
	 a. It helped revolutionize astronomy. b. It predicted the motions of the planets better than any other theory at the time. c. It promoted new research in the study of the universe. d. It explained the sun's role in the universe.
1	 O. What was Copernicus's theory of the universe? a. The Earth was at the center of the universe, and the sun, moon, and other planets revolved around it. b. The sun was at the center of the universe, and the Earth and the other planets revolved around it. c. The sun and moon revolved around Earth, but the other planets revolved around the sun. d. The planets revolved around the sun in elliptical orbits.
1	1. Which astronomer's theory led to major changes in science and society?a. Hubbleb. Brahe
	c. Ptolemy d. Copernicus
1	 What was Brahe's theory of the universe? The Earth was at the center of the universe, and the sun, moon, and other planets revolved around it. The sun was at the center of the universe, and the Earth and the other planets revolved around it. The sun and moon revolved around Earth, but the other planets revolved around the sun. The planets revolved around the sun in elliptical orbits.
1	 3. Why was Brahe's work helpful even though his theory of the universe was incorrect? a. He accurately described the planets' orbits. b. He made detailed measurements of the sun. c. He explained the sun's role in the universe.

d. He made very precise observations of the planets and stars.

Name	Class	Date
Directed Reading A continued		
14. What did Kepler beli	eve about the universe	?
		rse, and the sun, moon, and
other planets reve		1.1
		e, and the Earth and the
other planets rev		h, but the other planets
revolved around		ii, out the other planets
	ved around the sun in e	elliptical orbits.
15. What laws did Keple	r state that are still in u	ıse today?
a. laws of planetary		•
b. laws of planetary		
c. laws of solar mot	ion	
d. laws of gravity		
16. Who was one of the	first scientists to use a	telescope?
a. Galileo		
b. Kepler		
c. Newton d. Hubble		
17. What four discoveries did 0 bodies like the Earth and n		ed planets are physical
bodies like the Earth and h	or waridering stars:	
18. What did Newton prove ab	out gravity?	
what did rewish prove as	out gravity.	
19. Newton's laws of motion a	nd gravity helped to ext	olain many other scientists'
observations. For example, Newton's laws?		_
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Name Date	
Directed Reading A continued	
MODERN ASTRONOMY	
20. What were two milestones in the development of modern astronomy?	
21. What did many astronomers believe about galaxies prior to the 1920s?	
22. What did Edwin Hubble prove in 1924?	
27 What haliaf of ather actual and all Edwin Halble's discourse of a	
23. What belief of other astronomers did Edwin Hubble's discovery confirm?	
24. What tools do astronomers use to study space?	
25. How are computers used to study space?	

Name	Class	Date
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Skills Worksheet



Directed Reading A

Section: Telescopes
1. An instrument that gathers electromagnetic radiation and concentrates it for
better observation is a(n)
2. The most common type of telescope is thetelescope.
OPTICAL TELESCOPES
 3. An optical telescope is an instrument that a. collects and focuses visible light for closer observation. b. collects and focuses invisible light for closer observation. c. collects visible light and breaks it apart. d. collects invisible light and breaks it apart.
 4. The simplest optical telescope contains a. an objective lens and a mirror. b. two objective lenses. c. an obtuse lens and a lens in the eyepiece of the telescope. d. an objective lens and a lens in the eyepiece of the telescope. 5. What is the function of the objective lens?

6. What is the function of the lens in the eyepiece?



Name	Class	Date
Directed Reading A continued		
Two types of optical telescopes a In the space provided, write FR if FL if the phrase describes a refle	the phrase describes	
7. uses lenses to gather	and focus light	
8. uses mirrors to gather	and focus light	
9. used by most professi	onal astronomers	
10. cannot focus images p	perfectly	
11. focuses all colors of l	ght to the same focal	point
12. distorts images if lens	s is too large	
13. can use large mirrors	to gather light	
14. flaws in the glass don 15. What are two disadvantages		_
16. What are three advantages o	f reflecting telescopes	s?
17. What do very large reflecting in one spot?	telescopes use to ga	ther more light and focus it



e on Earth to put	a telescope?
e on Earth to put	a telescope?
e on Earth to put	a telescope?
escope? Explain v	why
 	
	
-	visible light?
ic elements.	
•	l.
	pectrum nas a
electromagnetic r	adiation.
=	
irom	
olet light.	
tion in order from	shortest wavelandt
gamma rays, radio	_
	well prove about diation. cic elements. magnetic spectrum y the human eye. electromagnetic setic energy. electromagnetic ra rays. waves. from olet light.

ame	Class	D	ate
Directed Reading A cont	tinued		
ach of the following wave tmosphere. In the space (ne wavelength is unblock	provided, write B if		
25. infrared light			
26. gamma rays			
27. X rays			
28. visible light			
29. microwaves			
NONOPTICAL TELESCOP	ES		
30. Why do astronomers s	study the entire elec	tromagnetic spectr	um?
			
			
31. Radio telescopes are a			
because radio wavele		_	
52. The			nore flawed than
the lenses and mirrors 33. Why can chicken wire	-	_	scope?
with the control of t	o de lasea las une sun.		жо р о.
	· · · · · · · · · · · · · · · · · · ·	 	
34. Why have scientists p in space?	out ultraviolet, infrar	red, gamma-ray, and	X ray telescopes
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Studving Space

Holt Science and Technology

WordSearch 09/27/06, Earth Science

Name	<u> </u>	. <u> </u>	Date
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Instructions: Complete the puzzle. Use the clues to help you find and circle the terms hidden in the puzzle.

ZMUKGRTXPEDLCONSTELLAT IONAG AHRMBEZBHSXQGA JP CMZU ł Р J ٧ Z Κ 0 DPT U BGDMX NUGY В K Q Е Κ R X W Υ В 0 0 L-Q R M E G 0 $P \cdot J$ G Е T F 1 UY Ζ K J В B S C S F В Κ В Α UW Н Z н в Х Р C Ζ DNUO Ρ Ε S ΝU Υ N G Т Т E ΑZ 0 N B Н Α Ν V Υ Q ٧ Υ В U Ν J U G O С C T F J Ε Υ С Ε X W U U NBV F F М J B Ľ Т HWKВ K AYEQ Ε D W 0 D 0 Q L CYOU С L J ZA 1 T O S C S U G D 1 ZGPQHA S 1 ٧ S V Ε Υ S D JNR GF BWSK M M S Α Ζ V Е 1 ח N NCKS R Ρ PHHCOB ΗZ J D - 1 K Q G S Q Ζ MZEREF R Α С T 1 Ν GΤ Ε L Ε S С 0 Ε Κ Z GXDQA WPMQ T S 0 H NΑ L T 1 U D ΗТ Т Ρ AQE CZJ S В RHO0 Α Т Υ Н D Κ Υ NQPSO HXUR KNOΖ Н F Υ J 1 Н Ζ E O W S RQE R B Υ CKE GMOMD U Ν W ΚX Q OТ В Υ 0 Υ ΥZ BWZDΕ C Α S В H D Χ 0 S R V HKDF Α Υ M O N0 R T F Ε Α Ν U S Α S Υ X B M OAWP GKDZ Υ HV Ε В В Α R G 0 D 0 J T H F Р Ζ LQDEWL L U T J UL NPO R UMFQ T E M E ADMGNF Ε RG 0 ΝP Τ EUP CQ Ζ NMGHΥ 1 BOXΑ Υ W C ΧG В ECT ROMAGNE T C S Ρ E C ZQVEL 1 Т RUMH BUVQAZCOPF UZMXJ JP S 1 V J NQXE GXOKKQCWTZEMBNAMJUPXMUHNJ

Clues:

- 1. telescope
- 2. year
- 3. horizon
- 4. day
- 5. month
- 6. refracting telescope
- 7. astronomy
- 8. altitude
- 9. zenith
- 10. light-year
- 11. constellation
- 12. reflecting telescope
- 13. electromagnetic spectrum

Scramble 09/27/06	, Earth Science		•
Name	Date,		
Instructions: Complete	the puzzle. Use the clues to hel	lp you unscramble	e the terms
1. D Y A			T
2. IERAL-HYTO	g D		
3. RYEA			
4. NCITLTEMC	AREGEO TRPSUMCE		
5. OEILALOCT	SNTN		
6. MYTOORASN	·	ı	
7. HZNTEI			
8. NHRIZOO			
9. AEDLITTU			:
10. SLPOCETEE	1		
11. CFEARGTIN	R EOEPTCESL		, ,
12. TNCEFERLI	G ELSCPETEO		
13. MOTNH			·
Clues: 1. the time required to	for Earth to rotate once on its axis	s	
2. the distance that I	ight travels in one year; about 9.4	6 trillion kilometers	
3. the time required	for the Earth to orbit once around	the sun	
4. all of the frequence	cies or wavelengths of electromagne	etic radiation	

5. a region of the sky that contains a recognizable star pattern and that is used to describe the location of objects in space

page 1

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Info For Chapter 18 test

Altitude Zenith

Horizon vernal Equinox

Light-Year Celestial Sphere

Kepler Newton

Hubble Galileo

Telescope Electromagnetic Spectrum

Reflecting Telescope Refracting Telescope

Tycho Brahe Copernicus

Ptolemy day

Year month

Celestial sphere declination

Ecliptic constellation

Right ancension Doppler Effect

Red Shift Blue Shift

Circumpolar Know the Electromag. Spectrum

How does earths atmosphere affect starlight

- 6. the study of the universe
- 7. the point in the sky directly above an observer on Earth
- 8. the line where the sky and the Earth appear to meet
- 9. the angle between an object in the sky and the horizon
- 10 an instrument that collects electromagnetic radiation from the sky and concentrates it for better observation
- 11. a telescope that uses a set of lenses to gather and focus light from distant objects
 - 12. a telescope that uses a curved mirror to gather and focus light from distant objects
 - 13. a division of the year that is based on the orbit of the moon around the Earth

Name	 _ Class	Date .	

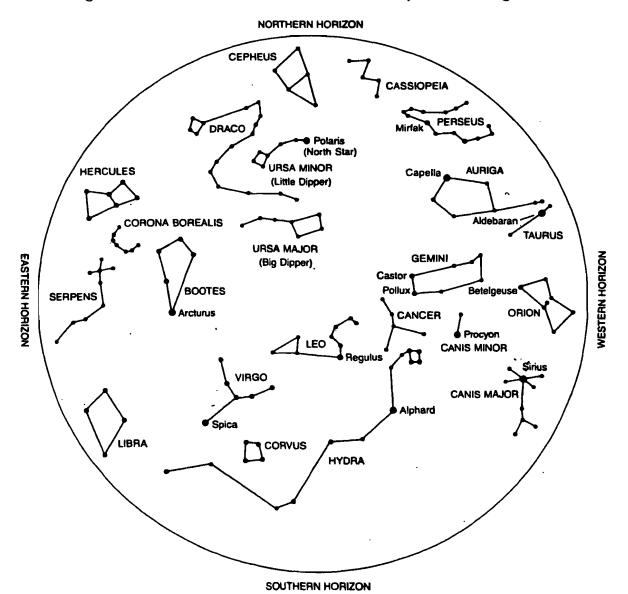
ACTIVITY Stars and Galaxies

CHAPTE

Spectacular Night Sky

Interest in the night sky led primitive people to study the patterns of stars they saw and to give these patterns names. The patterns, or constellations, were named in honor of various mythological characters. In all, some 88 constellations have been identified. Some are easy to locate in the night sky while others are not. Can you think of some explanations for this?

The diagram below is a star chart of the spring sky in a particular location. The outer edge of the circle represents the horizon as the observer faces north, south, east, or west. To use the star chart, you would place the paper horizontally above your head and align the correct edge of the circle with the direction in which you are facing.



1.	In which direction should you face to see the constellation Ursa Minor?
2.	What well-known star is the tip of the handle of Ursa Minor?
3.	Which constellation is directly overhead in spring?
4.	What is another name for the Big Dipper?
5.	How can you use the Big Dipper to find the Little Dipper?
6.	In which constellation is the star Sirius located?
7.	Why is Sirius easy to identify?
8.	In which direction would an observer have to look to see the constellation Taurus?
9.	Why do the stars in the night sky appear to twinkle?
10.	How many stars can be seen in the night sky by any one observer at any one time?
11	How many constellations can you identify in the night sky?

Name	Class	Date
14/10		Duit

ACTIVITY Stars and Galaxies

Constellation

2

CHAPTE

Naming Constellations

Ask your friends what sign they were born under and they can probably tell you immediately. These zodiac signs are actually the names ancient peoples gave to various constellations, or star groups. You will find a list of constellations below. However, in each case, either the constellation's name or its English equivalent has been left out. Fill in the constellation's name or its English equivalent in the appropriate space.

English Name

1.	Aquarius	
2.		Ram
3.	Cancer	
4.	Capricornus	
5.		Princess
6.	Aquila	
7.		Swan
8.		Dragon
9.	Gemini	
10.	Hercules	
11.		Lion
12.	Libra	
13.		Wolf
14.	Hydra	
15.	Orion	
		U
17.		Scorpion
18.	Taurus	

19.	Cassiopeia	
20.		Fish
21.	Sagittarius	
22.		Great bear
23.	Ursa Minor	

Name	Class	Date		

Skills Worksheet



L	Directed Reading A
	ection: Mapping the Stars How did ancient cultures group the stars in the sky?
2.	What are two things that people have a better understanding of as a result of
	advances in astronomy?
	TTERNS IN THE SKY 3. What are constellations? a. regions of the sky that contain recognizable star patterns b. stars c. star patterns d. galaxies
	 4. How did people in ancient cultures use the locations and movements of constellations? a. to create land boundaries b. to make roads c. to measure the universe d. to navigate and keep track of time
	 5. The ancient Greek constellation Orion was the same as a. the Japanese constellation of a hunter. b. the Japanese constellation of a drum. c. the Great Bear. d. the Ursa Major.
	 6. Which of the following is true of constellations? a. All cultures interpret the sky in the same way. b. Every star or galaxy belongs to a constellation. c. All ancient civilizations had the same names for the same constellations.



constellations.

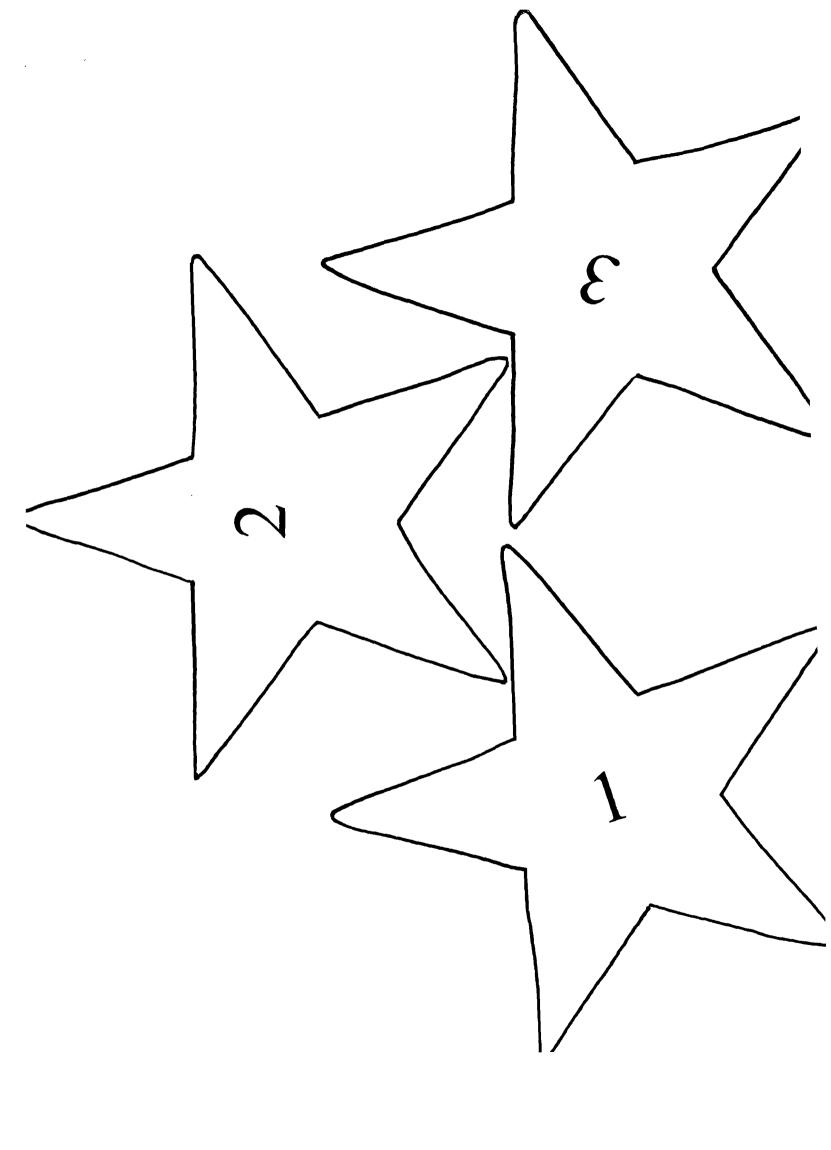
 ${f d.}$ Astronomers disagree on the names and locations of the

lame	Class	Date
Directed Reading A continue	ed	
7. How many constell a. 88 b. 128 c. over 1,000 d. over 10,000	llations are there?	
8. The apparent local locations a. every day. b. from season to c. every year. d. every other year.	season.	n the night sky change their
a. because the Eab. because the Eac. because the sta	ions seem to change loo rth tilts on its axis rth revolves around the ars rotate around the Ea astronomical optical illu	rth
 a. People in all pa b. People in Chile States. c. People in the None people in the Solution d. People in the None People in the N	Jorthern Hemisphere sec outhern Hemisphere.	
FINDING STARS IN THE NIC		planet's location is a(n)
12. What are three reference relation to a person's pos	_	e a star or planet's position in
	·	

Name	Class	Date
Directed Reading A continued		
Match the correct definition with the co provided.	rrect term. W	rite the letter in the space
13. an imaginary point directly a	above an	a. zenith
observer's head		b. celestial equator
14. the line where the sky and the	he Earth	c. horizon
appear to meet		d. altitude
15. the angle between an object and the horizon	in the sky	e. celestial sphere
16. an imaginary sphere that sur the Earth	rrounds	
17. an imaginary extension of the equator into space	ne Earth's	
18. The location of the sun on the first	day of spring	is the
19. Astronomers measure		_ in hours by how far east ar
object is from the vernal equinox.		v
20. Astronomers measure		_ in degrees by how far nort
or south an object is from the celes		•
21. Some stars located near Earth's polnight. What are these stars called?	les can be see	en year-round, at all times of
THE SIZE AND SCALE OF THE UNIVER		ight travala in
22. A light-year is equal to the d a. 1 month.	istance mai i	ight travels in
b. 1 year.		
c. 9.46 years. d. 9.46 trillion years.		
•		
23. One light-year is about 9.46	trillion	
a. yards. b. meters.		
c. kilometers.		
d. miles.		
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	Name		Class	Date
 a. about 1 billion light-years b. 9.46 billion light-years c. more than 10 billion light-years d. about 100 billion light-years THE DOPPLER EFFECT 25. What is the name of the effect that describes how the pitch of a sour seems higher as it gets closer and lower as it gets farther away? a. sound effect b. wavelength effect c. drowser effect d. doppler effect 26. When a star or galaxy moves quickly away from an observer, the light it emits a. appears bluer than it usually would. b. appears redder than it usually would. d. appears lighter than it usually would. 27. When a star or galaxy moves quickly toward an observer, the light it emits a. appears bluer than it usually would. b. appears redder than it usually would. c. appears darker than it usually would. d. appears lighter than it usually would. 28. An effect in which a star or galaxy appears to move quickly toward an observer is called 29. An effect in which a star or galaxy appears to move quickly toward an observer is called 30. Edwin Hubble discovered that the light from all galaxies except the Milky Way's close neighbors is affected by 	Directed	Reading A continued		
	24.	a. about 1 billion light-yearsb. 9.46 billion light-yearsc. more than 10 billion light-years	vears	can see?
seems higher as it gets closer and lower as it gets farther away? a. sound effect b. wavelength effect c. drowser effect d. doppler effect 26. When a star or galaxy moves quickly away from an observer, the light it emits a. appears bluer than it usually would. b. appears redder than it usually would. c. appears darker than it usually would. d. appears lighter than it usually would. 27. When a star or galaxy moves quickly toward an observer, the light it emits a. appears bluer than it usually would. b. appears redder than it usually would. c. appears darker than it usually would. d. appears lighter than it usually would. 28. An effect in which a star or galaxy appears to move quickly away from an observer is called	THE DOPI	PLER EFFECT		
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observer is called 29. An effect in which a star or galaxy appears to move quickly toward an observer is called 30. Edwin Hubble discovered that the light from all galaxies except the Milky Way's close neighbors is affected by	27.	emitsa. appears bluer than it usuab. appears redder than it usuc. appears darker than it usu	lly would. ally would. ally would.	rd an observer, the light it
29. An effect in which a star or galaxy appears to move quickly toward an observer is called30. Edwin Hubble discovered that the light from all galaxies except the Milky Way's close neighbors is affected by	28. An eff	ect in which a star or galaxy a	ppears to mo	ve quickly away from an
observer is called 30. Edwin Hubble discovered that the light from all galaxies except the Milky Way's close neighbors is affected by	observ	ver is called		
30. Edwin Hubble discovered that the light from all galaxies except the Milky Way's close neighbors is affected by	29. An eff	ect in which a star or galaxy a	ppears to mo	ve quickly toward an
Way's close neighbors is affected by	observ	ver is called	·	
	30. Edwin	n Hubble discovered that the li	ght from all g	alaxies except the Milky
Jimow and Bawin Habble determine that the draverse mast be expanding.				



Name	Class	
Skills Worksheet		
Reinforcement		

Complete this worksheet after you finish reading the section "Astronomy: The Original Science."

Stella Star, a reporter for the *Back in Times*, interviewed scientists for an article she is writing on the history of astronomy. She interviewed Ptolemy, Kepler, Copernicus, Newton, Galileo, Brahe, and Hubble. When Stella looked over her notes, she discovered that she forgot to indicate which scientist said what. Help Stella organize her notes by identifying the scientist most likely to have said each statement. Write your answer in the space provided.

Ptolemy Kepler Copernicus Newton Galileo Brahe Hubble 1. "I've finally worked out an explanation as t planets orbit the sun and moons orbit planet gravity that keeps an object in orbit!" 2. "The Earth is the center of the universe, an the planets and stars orbit our planet. My the predicts the motion of the planets better the					
 "I've finally worked out an explanation as t planets orbit the sun and moons orbit plane gravity that keeps an object in orbit!" "The Earth is the center of the universe, an the planets and stars orbit our planet. My the predicts the motion of the planets better the 					
planets orbit the sun and moons orbit plane gravity that keeps an object in orbit!" 2. "The Earth is the center of the universe, an the planets and stars orbit our planet. My the predicts the motion of the planets better the					
the planets and stars orbit our planet. My t predicts the motion of the planets better th					
other theory of my day."	heory				
many of you think our galaxy is the only on	3. "It appears the Milky Way has neighbors. I know many of you think our galaxy is the only one, but it looks like that's just not so. This means the universe is a lot bigger than we thought."				
4. "I prefer the theory of the sun-centered unit over other theories. The telescope has help to discover that the planets are not just dot light, but physical bodies like the Earth."	ed me				
I conclude that our universe is Earth-center sun and the moon revolve around the Earth the other planets revolve around the sun."	red. The				
Earth is at the universe's center, but I'm cer that the sun is at its center. I am certain the planets orbit the sun."	rtain				
7. "I do not agree with my mentor's theory that sun revolves around the Earth. I have used precise data to propose another theory in wall of the planets revolve around the sun in elliptical orbits."	his vhich				

Name:	- 11	_ _	Class:		Date:			•	ID:
Chapt	er 1	8		1		ţ		•	
, Multip l Identify			npletes the stateme	nt or answe	rs the auestion.		,	•	
	,				y				
· · ;	1 .	Which of the folloa. X rays b. microwaves	owing types of elect	romagnetic c. d.	radiation is blo infrared light visible light	cked by the	Earth	's atmos	sphere
· .	2.	Why do astronoma. to reduce air jb. to get closer t	to objects in space ference from the Ea	-			•		
	3.		ows that the universemoving closer toget moving apart.	ther. c.				ing.	
	4.	What is a constell a. a region of the b. a group of sta	ie sky	. c. d.	a star pattern a galaxy			,	
 ,	5.	What does a teles a. magnetic part b. electromagne		pace and foo c. d.	cus for closer of lenses wavelengths	oservation?		•	
	6.	What is the imagi a. a zenith b. a celestial spl	nary sphere, created	d by scientis c. d.		ds the Earth	1?		
	7.	a. Galaxies werb. Galaxies werc. Galaxies wer	tell the universe wa e moving away from the moving toward ear the getting bigger. of galaxies was inco	n each other ich other.			•		
	8.	a. a growing ceb. an increase in	n the number of cor of other galaxies						
	9.	Where do scientis a. in deserts b. in valleys	sts put telescopes to	avoid inter c. d.		arth's atmos	sphere'	?	
	10.	What is an imagina. celestial spheb. zenith	nary point directly a ere	above an obe c. d.		ı			

Name:				ID: A
				· · · · · · · · · · · · · · · · · · ·
	11.	Which of the following is NOT a type of electronspectrum? a. microwave c. b. visible light d.	omagnetic radiation found on the electron radio wave ocean wave	nagnetic
	12.	How does the Earth's atmosphere affect starlight. a. It blocks it. b. It stretches it. d.		
	13.	Which of these would be shorter if Earth rotated a. years c. b. months d.		. u
	14.	The vernal equinox is used to establish a star's a. zenith. c. b. distance from the Earth. d.	declination. right ascension.	ı
	15.	Copernicus's theory was not accepted when he a. the center of the universe. c. b. an average star. d.	a source of energy.	ne sun wa
	16.	 An X-ray telescope is NOT used on Earth beca a. blocked by the Earth's atmosphere. b. destroyed by the Earth's magnetic field. c. very dangerous to humans. d. distorted by the Earth's winds. 	use X rays are	, ·
	17.	Which of the following scientists thought the E a. Ptolemy c. b. Hubble d.	Newton	
	18.	•	iation? starlight sunlight	
	19.	What did redshift tell Hubble about the univers a. The universe is getting smaller. c. b. The universe is getting larger. d.	The universe is getting colder.	(
	20.	How long does Earth take to orbit once around a. day c. b. week d.	month	
	21.	How long does Earth take to rotate once on its a. day c. b. week d.	month	
	22.	About how long does the moon take to orbit the a. day c. b. week d.	month	

Name: _			ID: A
	23.	Which of the following is NOT found on the electromagnetic spectrum? a. gamma ray c. ultraviolet light b. X ray d. black hole	1 ,
Complet Complet		1 ech statement.	
	24.	A phenomenon in which sound seems to increase or decrease in relation to the direction it is the	is movi
	25.	An effect in which a star or galaxy appears to move quickly away from an observer is calle	. d
	26.	When a star or galaxy appears to move quickly toward an observer, an effect called occurs.	
	27.	A nonoptical telescope that detects X-rays is the	
		Stars that can be seen in all seasons and that never set are	,
		Use the terms from the following list to complete the sentences below.	and the property of
estatus personali oligaji makan personali se		altitude declination ecliptic constellation month reflecting telescope right ascension year refracting telescope	
	29.	Ursa Minor is an example of a(n)	
	30.	An instrument that uses a mirror to gather and focus light is a(n)	'
	31.	The angular distance between a star and the horizon is the star's	
And the state of t	32.	The apparent path of the sun across the celestial sphere as seen from Earth is called the	
	33.	One can tell how far north or south an object is from the celestial equator by the object's	
	34.	The amount of time the moon takes to orbit the Earth is roughly a(n)	

Name:					ID
				;	
	Use the terms from th	e following list to complete t	he sentences below.	,	
· !	zenith	horizon	ı	. 1	
	constellation	celestial sphere	• •		•
			49		
35	An imaginary sphere	that surrounds the Earth is th	ne	·	•
36	The line where the sky	v and Earth appear to meet is	s the		

Short Answer

- 39. What advantage did Galileo have over Ptolemy in understanding the structure of the universe?
- 40. What would it mean about the size of the universe if the galaxies were moving toward each other instead of apart from each other?

Matching

Match each item with the correct statement below.

37. An imaginary point directly above an observer's head is a

38. A region of the sky with distinct star patterns is a _

a. Galileo Gallilei

f. Edwin Hubble

b. Tycho Brahe

g. Ptolemy

c. Johannes Kepler

h. day

d. Sir Isaac Newton

i. year

e. Copernicus

- j. month
- 41. showed that planets and moons stay in orbit due to gravity
- 42. was one of the first persons to use a telescope to observe celestial bodies
- 43. developed a theory of a sun-centered universe
- 44. stated that planets move in elliptical orbits around the sun
- 45. used a mural quadrant to measure the positions of planets and stars
- 46. developed a theory of an Earth-centered universe in 140 CE
- 47. proved the existence of galaxies other than the Milky Way
- 48. the time required for the Earth to orbit once around the sun
- 49. roughly the amount of time required for the moon to orbit once around the Earth
- 50. the time required for the Earth to rotate once on its axis

	Name:		me:		,	1	ID: A
,					1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1		
	,	,	Match each item with the corre a. electromagnetic spectrum b. telescope	c.	w. reflecting telescope refracting telescope	•	•
		52. 53.	uses mirrors to gather and focus all of the frequencies or wavele uses lenses to gather and focus collects electromagnetic radiati	ngths of electrom light		observation	
			Match each item with the corre a. altitude b. horizon c. light-year	ect statement belo d. e. f.	w. zenith vernal equinox celestial sphere		. (
		56. 57. 58. 59.	the sun's location on the first of an imaginary sphere that surrou the distance that light travels in the line where the Earth and sky the angle between an object in an imaginary point in the sky d	unds the Earth one year y appear to meet the sky and the ho			
,			Match each item with the corre a. Kepler b. Hubble showed that gravity keeps plant showed that planets have ellipti	c. d. ets in orbit	w. Newton Galileo		

63. proved there were many galaxies

64. one of the first scientists to use a telescope