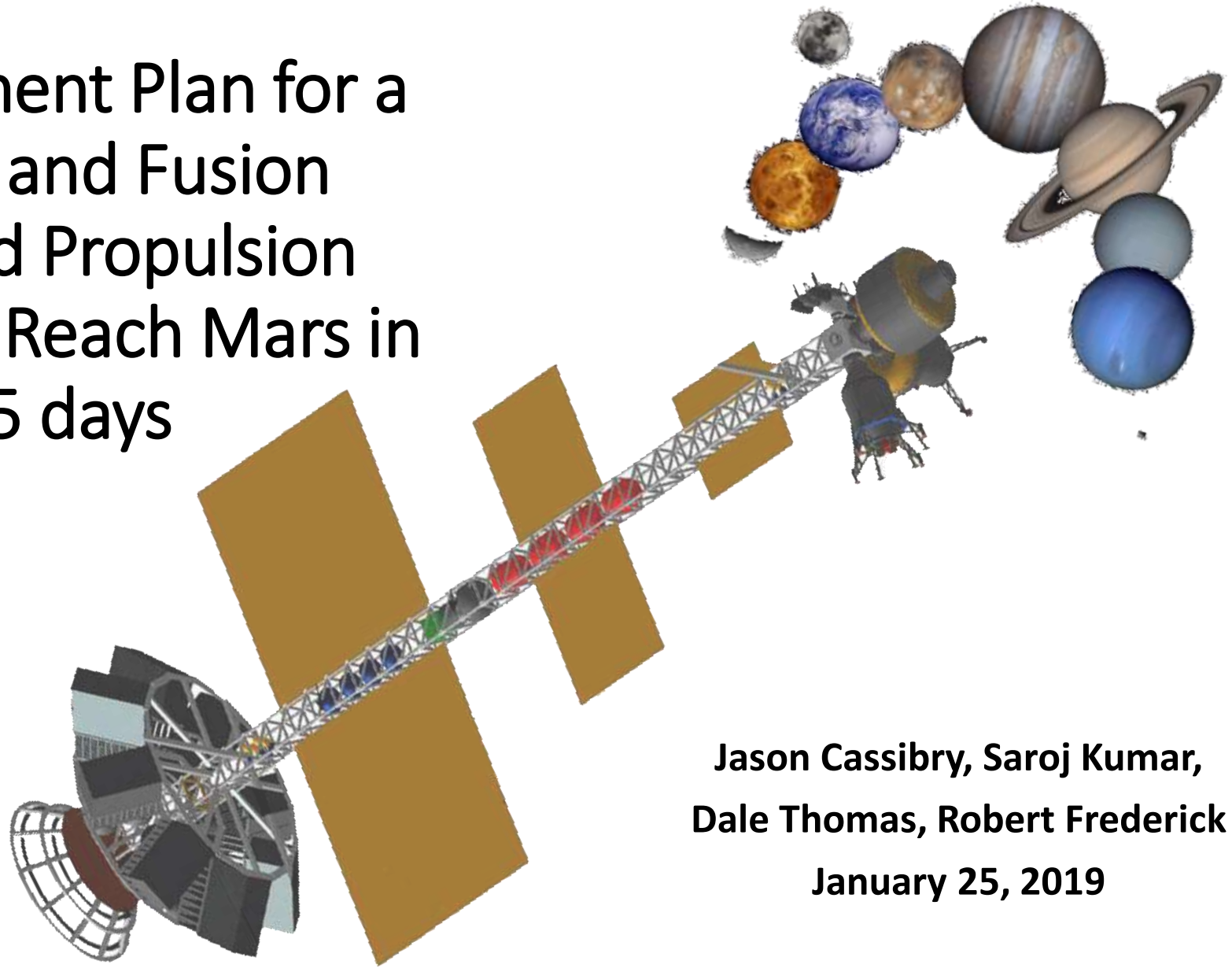


Development Plan for a Fission and Fusion Powered Propulsion System to Reach Mars in 45 days

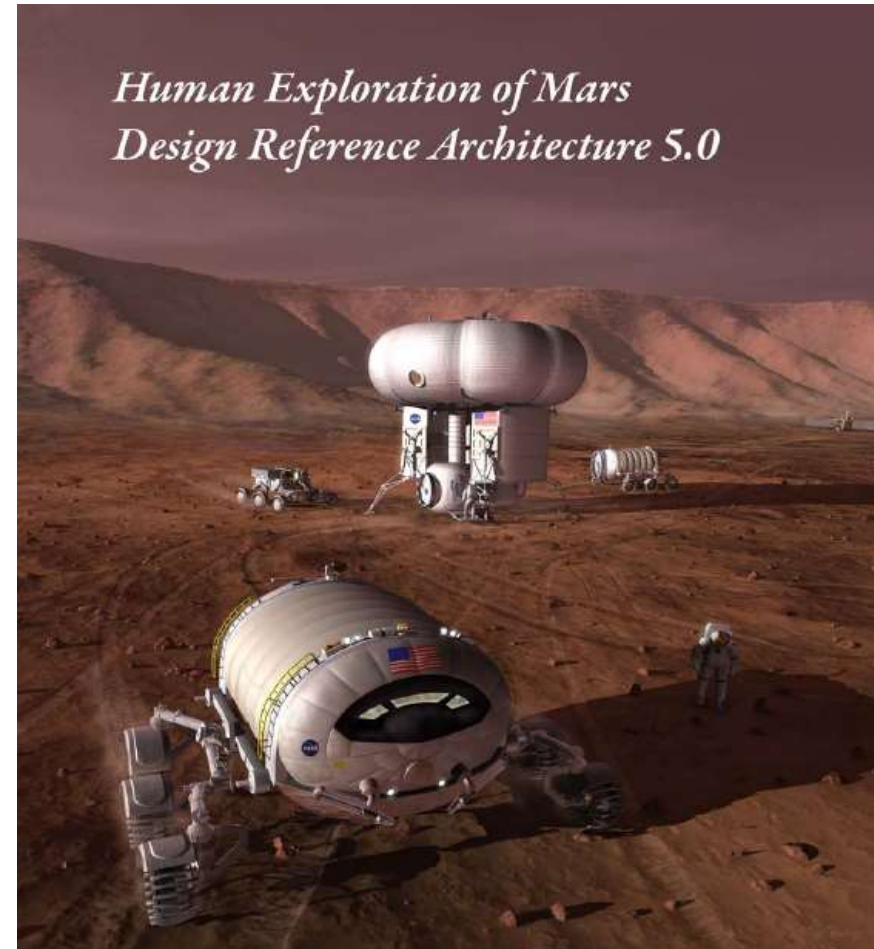


**Jason Cassibry, Saroj Kumar,
Dale Thomas, Robert Frederick
January 25, 2019**



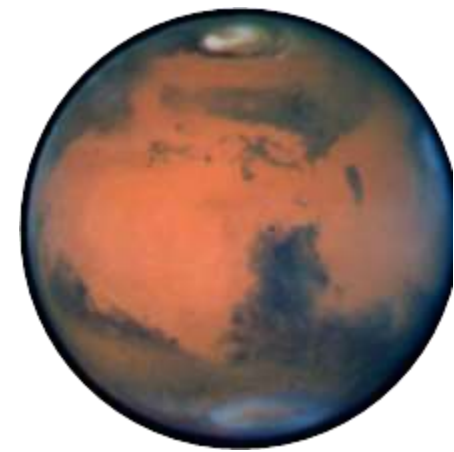
Why Send Humans to Explore Other Planets?

- Faster speed and higher efficiency to optimize field work
 - Agility and dexterity to go places that are difficult for robotic access
 - Innate intelligence, ingenuity, and adaptability to evaluate in real time and improvise to overcome surprises
- Overcome communication problems
 - time lag with mission control, e.g. 6- to 20-minute communications transit time for Mars
 - small number of daily uplink and downlink communications passes



Benefits of Nuclear Technology in Space

Routine Human Piloted Mars Missions



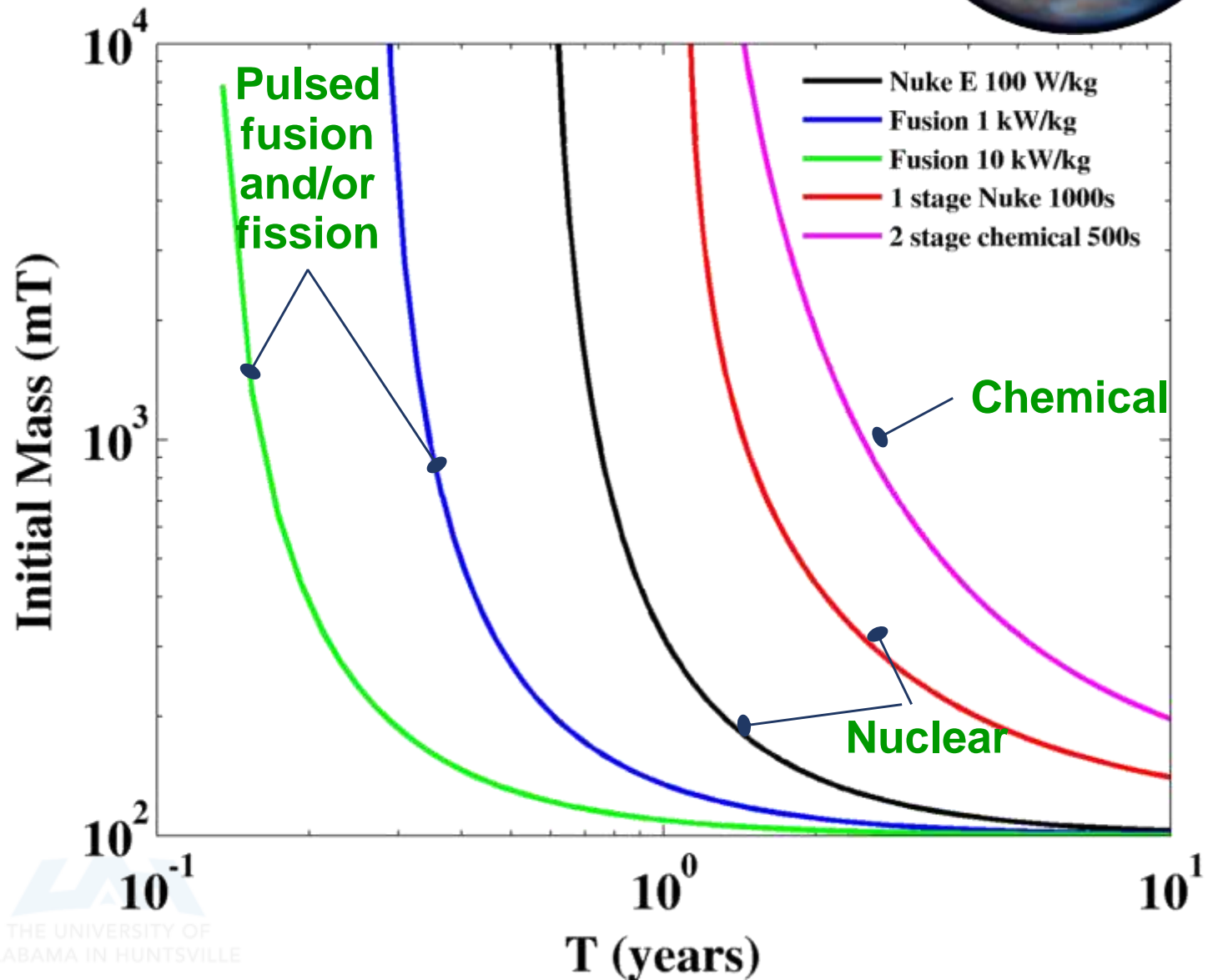
Roundtrip to Mars Initial Mass in LEO for 100 mT payload

Performance

- Larger Payloads
- KW to MW of 'House' Power

Safety

- Faster Trip Times
- Mission Abort Capability

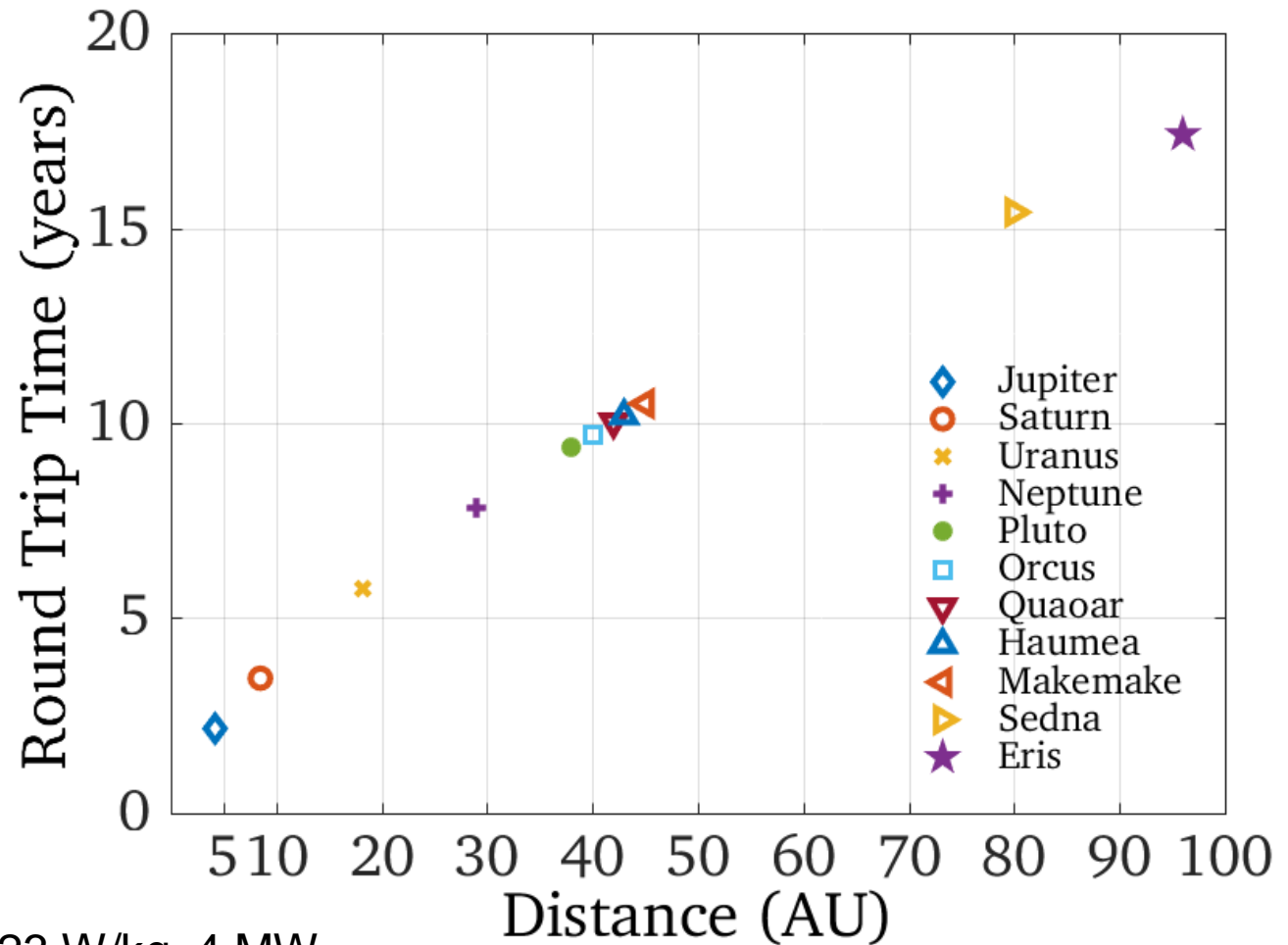


Benefits of Nuclear Technology in Space

Rapid Sample Return Missions from Deep Space

Roundtrip sample return trip times for a 100 metric ton (IMLEO) vehicle

- $\alpha=1$ kW/kg
- Assumed distance is the perihelion of the celestial body orbit



Comparable specific powers

- Diesel engine on freight train 23 W/kg, 4 MW
- NERVA thermal nuclear rocket 41 kW/kg, 1.4 GW
- KRUSTY nuclear electric reactor, 6 W/kg, 6 kW

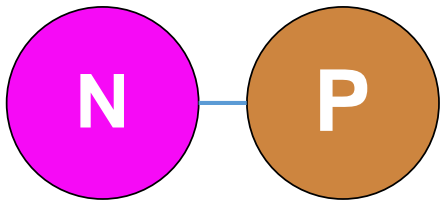


What is fusion?

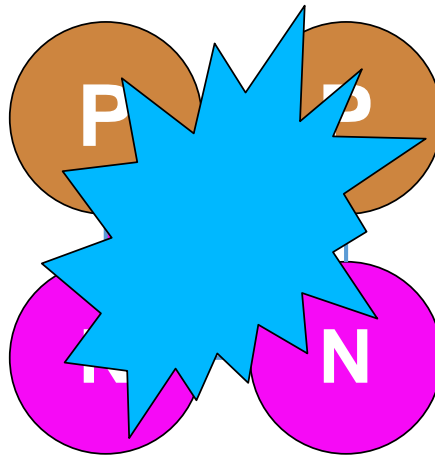
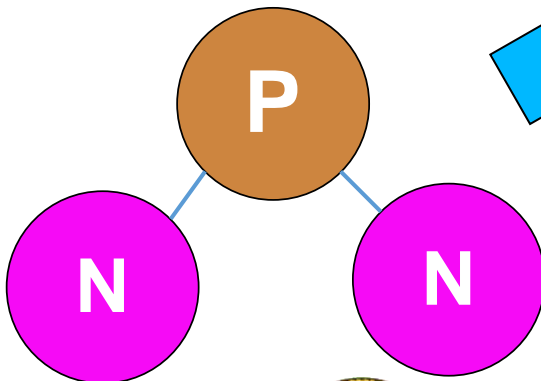
- The process by which multiple like-charged atomic nuclei join together to form a heavier nucleus.

Neutron + a lot of energy

Deuterium



Tritium

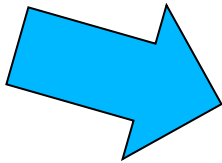


Helium + a lot of energy

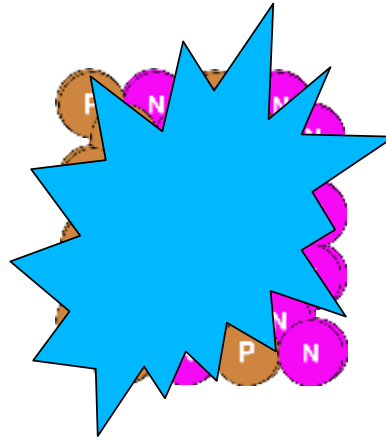
What is fission?

- The process by which a neutron strikes a heavy nucleus, causing the nucleus to split into two smaller fragments.

Neutron



Unsuspecting
Heavy Nucleus



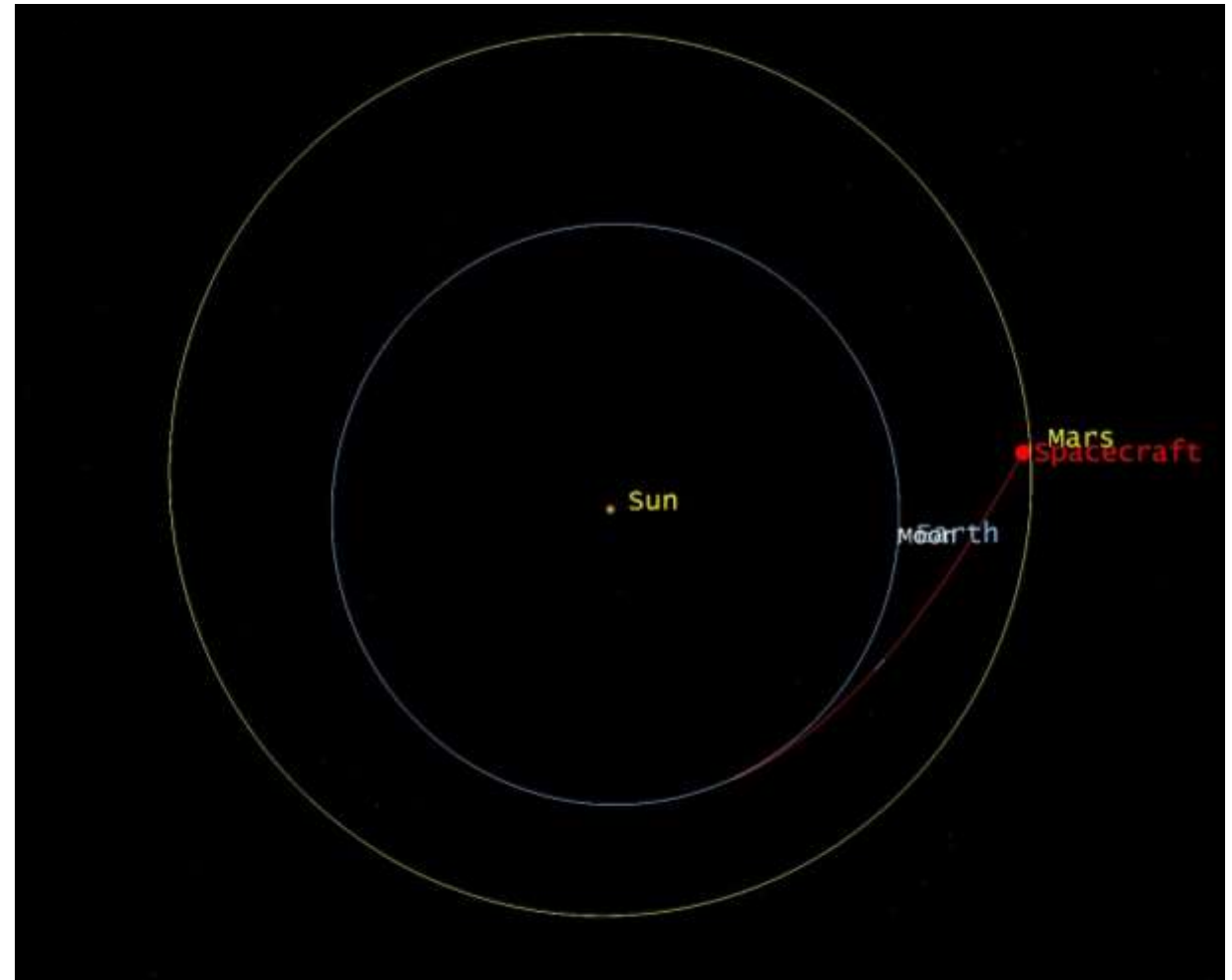
2 or 3 neutrons

2 fission fragments
and a lot of energy

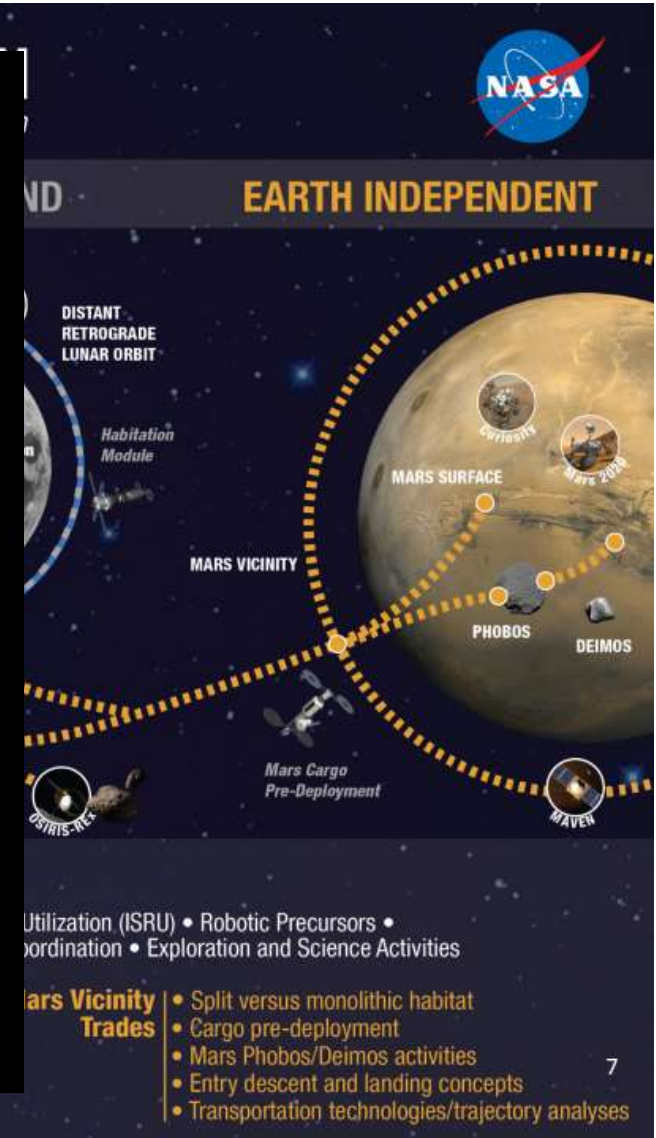
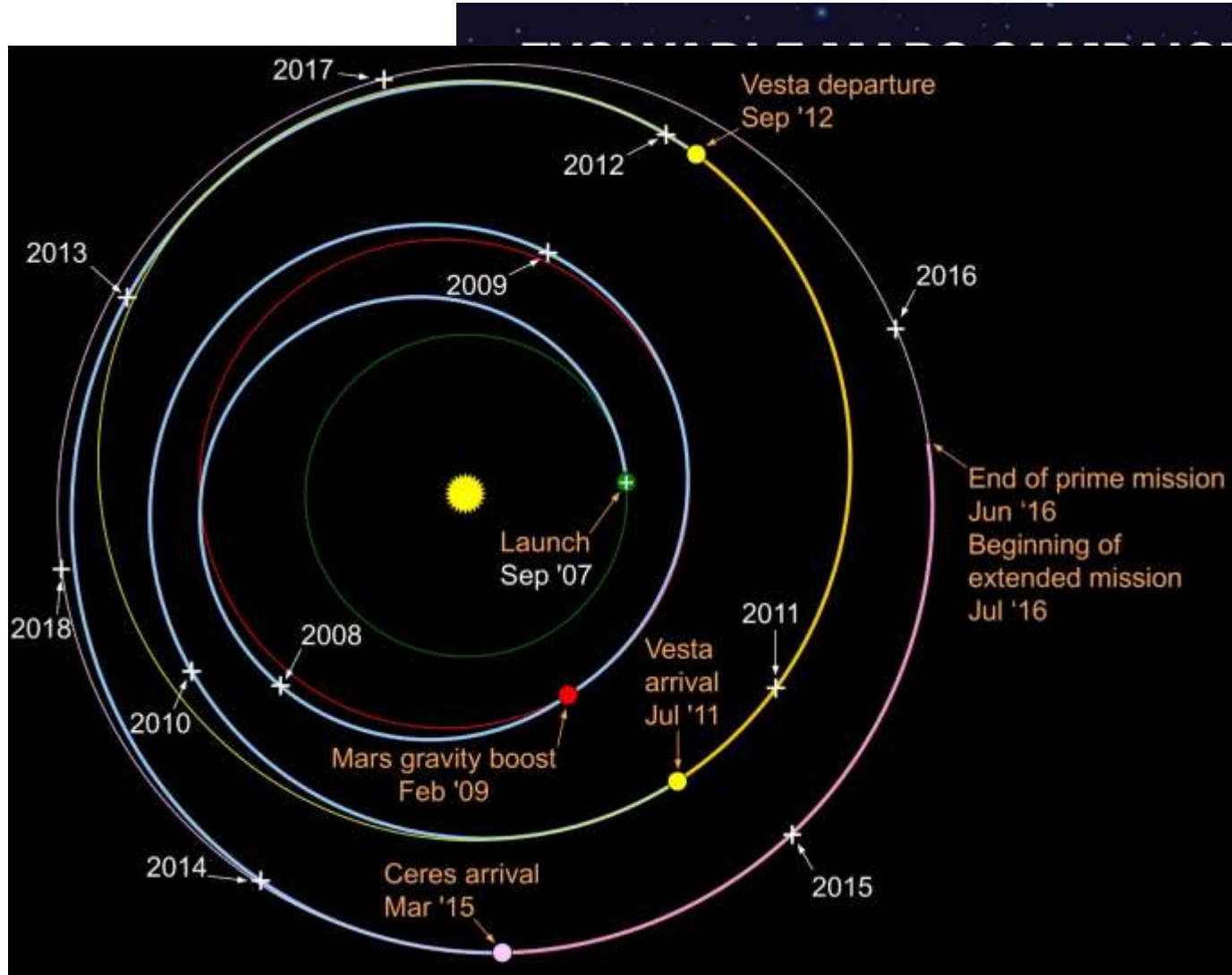
Preliminary medium fidelity analysis based on Mars mission with thrust of 250N and Isp of 5000 sec. The analysis was performed for heliocentric phase trajectory i.e, spacecraft departure from the Earth's SOI to Mars (flyby). The total trip time is 62 days with 2 days of coast phase.

Key challenges

- Departure from Earth sphere of influence
- Capture at Mars orbit
- The minor challenge that controlled thermonuclear fusion has not been accomplished in its 60 year history ...



The Problem for Advanced Propulsion Departure

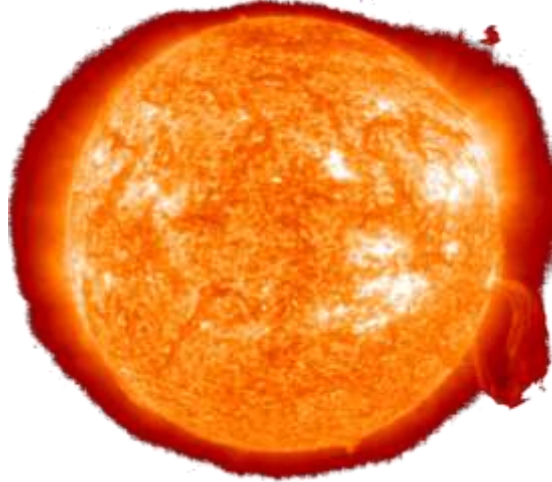


point and trajectory analyses



The Problems for Fusion

Temperature > 100,000,000 °K

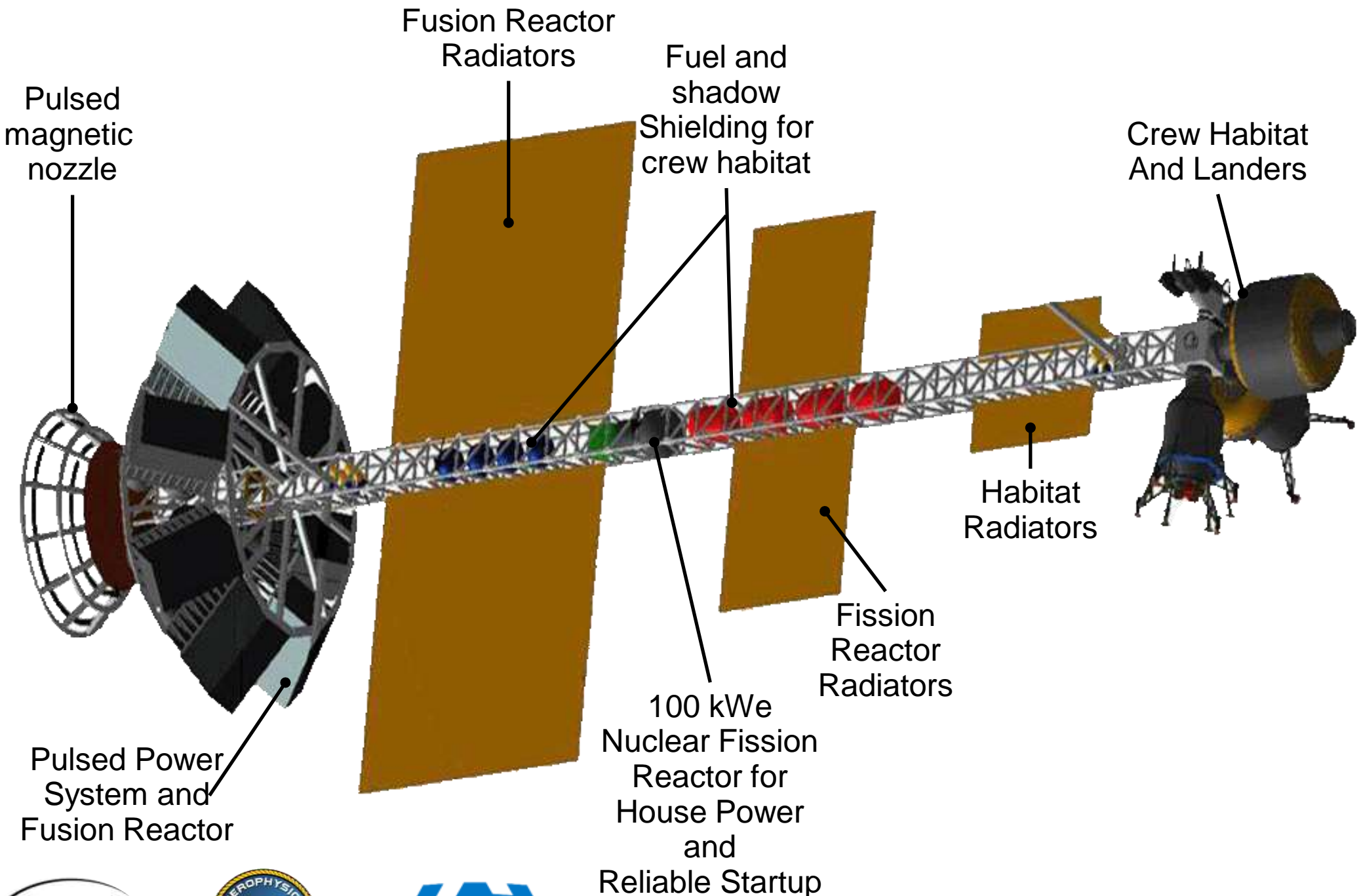


- Reactor > Yankee Stadium

- Costly Fuels

Costs per Kilogram

Tritium	\$30,000,000
^3He	\$1,231,000
$^6\text{Lithium}$	\$6,000
Deuterium	\$4,000



NASA's Vision is to reach for new heights and reveal the unknown, so that what we do and learn will benefit all humankind.

NASA's Mission is to drive advances in science, technology, aeronautics, and space exploration to enhance knowledge, education, innovation, economic vitality, and stewardship of Earth.

The NASA Space Technology Mission Directorate (STMD):

- Advances broadly applicable, transformational technology to infuse solutions into applications for which there are multiple customers
- Competitively selects technology development efforts based on technical merit
- Leverages the technology investments of other Government agency, academic, industry, and our international partners
- Coordinates with internal and external stakeholders, including academia, industry and other Government agencies
- Results in new inventions, new capabilities and the creation of a pipeline of innovators aimed at serving future national needs
- Grows the Nation's innovation economy and creates new high-tech jobs

Our vision: To utilize a complementary and multidisciplinary team to research and advance a bimodal fission and fusion hybrid propulsion system and associated technologies to TRL 3 or higher to help fulfill the NASA STRI goal of rapid interplanetary space exploration and interstellar precursor missions.

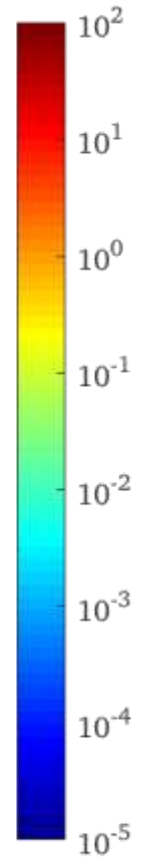
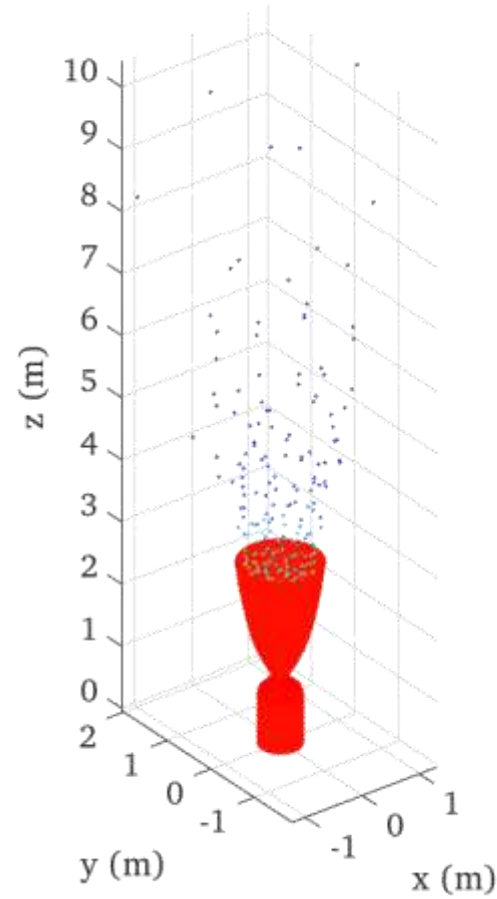
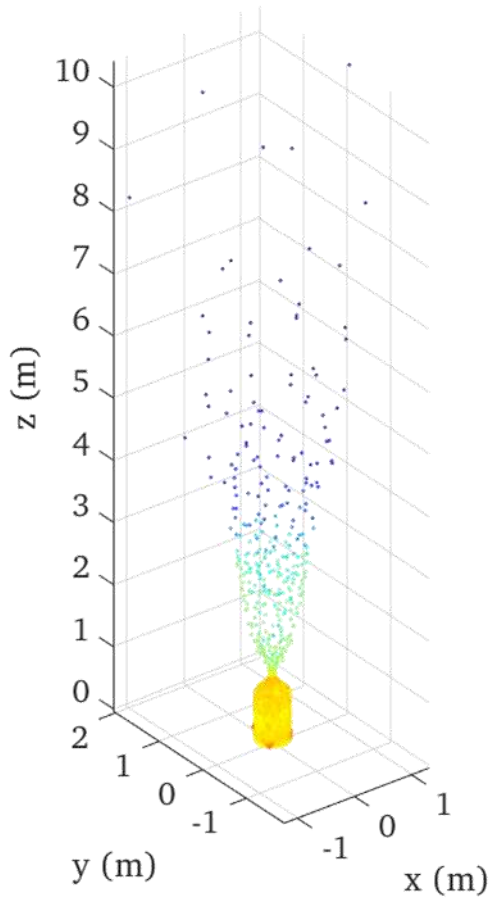


Example of a bimodal nuclear thermal rocket



“Nuclear Thermal Propulsion (NTP): A Proven Growth Technology for Human NEO/Mars Exploration Missions,” Borowski, Stanley K., McCurdy, David R., Packard, Thomas W., 2012 IEEE Aerospace Conference; 3-10 Mar. 2012; Big Sky, MT.

Pilot simulation of a 3D NERVA nozzle





Charger 1 Fusion Propulsion Facility





UAH Propulsion Research Center
 Prof. R. Frederick, PhD, Director
 Prof. D. Thomas, PhD, Deputy Director
<https://www.uah.edu/prc>



A. Edmondson, MBA
 Research Prog. Administrator/Budget Analyst
 M. Kitts, PRA; M. Nematl, OCE; J. Prince SS3

Eminent Scholar In Propulsion
 Prof. P. Ligrani, PhD

Aerospace Propulsion and Systems Integration Strategic Planning
 Prof. D. Thomas, PhD, Eminent Scholar in Systems Engineering
 Prof. P. Collopy, PhD, Chair of ISEEM Department; Prof. S. Mahalingam, PhD, Dean of the College of Engineering
 G. Karr, Principal Research Scientist VI; M. Miller, Research Scientist V
 A. Aseuron, GTA

Propulsion Laboratory and Safety
 Dr. D. Lineberry, PhD, Laboratory Operations
 T. Hall, Test Engineer

Energy & Power Systems
 Prof. P. Ligrani, PhD

Fusion Propulsion & Power
 Prof. J. Cassibry, PhD

Plasmas & Combustion
 Prof. G. Xu, PhD

Computational Modeling
 Prof. S. Rani, PhD

Aerospace Materials & Structures
 Prof. G. Nelson, PhD

Propellants & Energetics
 Prof. R. Frederick, PhD

Propulsion System Technology Test-bed
 D. Lineberry, PhD

J. Blackmon, PhD
 P. Collopy, PhD
 K. Frendi, PhD
 C. Kang, PhD
 S. Mahalingam, PhD
 G. Nelson, PhD
 S. Rani, PhD

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 P. McInturff, GRA; C. Ren, GRA;
 Z. Ren, GRA; M. Su, GRA; M. Suzuki,
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 M. Anderson, UGA; A. Clck, UGA;
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R. Cortez (RI)
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 T. Englestad, PTGS
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 B. Winterfing, GRA

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 K. Frendi, PhD

 R. Dextre, GRA
 R. Gott, GTA
 P. Salvador, GRA

 N. Latham, UGA
 N. Mann, UGA
 W. Manneschildt, UGA
 R. Nakano, UGA
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F. Ewere, PhD (Lecturer)
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 C. Kang, PhD
 S. Mahalingam, PhD
 S. Ravindran, PhD
 B. Shotorban, PhD

 V. Rani, GTA
 Wilson, GTA
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 SS5

K. Hazeli, PhD
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 J. Buckley, GRA; R. Chow, GRA;
 J. Indeck, GRA; D. Koda, UGA;
 P. Morrison, SS5

 E. Andrew, UGA
 G. Andrew, UGA
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J. Baird, PhD
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 G. Xu, PhD
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 C. Blankenship, PTGS;
 S. Bluestone, PTGS; C. Freeman,
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 Ranade, GTA

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 P. Ligrani, PhD
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 A. Bower, GRA; V. Braswell,
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 J. Agnew, UGA; D. Battle, SS4;
 J. Bgalow, MAE 490; B. Chubb,
 MAE 490; D. Corey, UGA; B. Ha,
 MAE 490; D. Hajian, MAE 490;
 J. Hobbs, MAE 490; J. Hunter,
 MAE 490; N. Long, MAE 490;
 F. McKee, MAE 490; W. Pico,
 MAE 490; E. Powers, MAE 490;
 E. Zimovan, MAE 490

GRA=Graduate Research Assistant; GTA=Graduate Teaching Assistant; MAE 490=Student in MAE 490; OCE=On-Call Employee; Post-Doctoral Fellow;
 PTGS=Part-Time Graduate Student; PRA=Propulsion Research Assistant; RI=Research Institute; SS3=Student Specialist III; SS4=Student Specialist IV; SS5=Student Specialist V;
 UGA=Undergraduate Assistant
 09/13/2017



UAH Fusion Propulsion Consortium



Recent milestones for making Charger 1 operational include the oil and water deionization systems

Oil system

Water deionization system



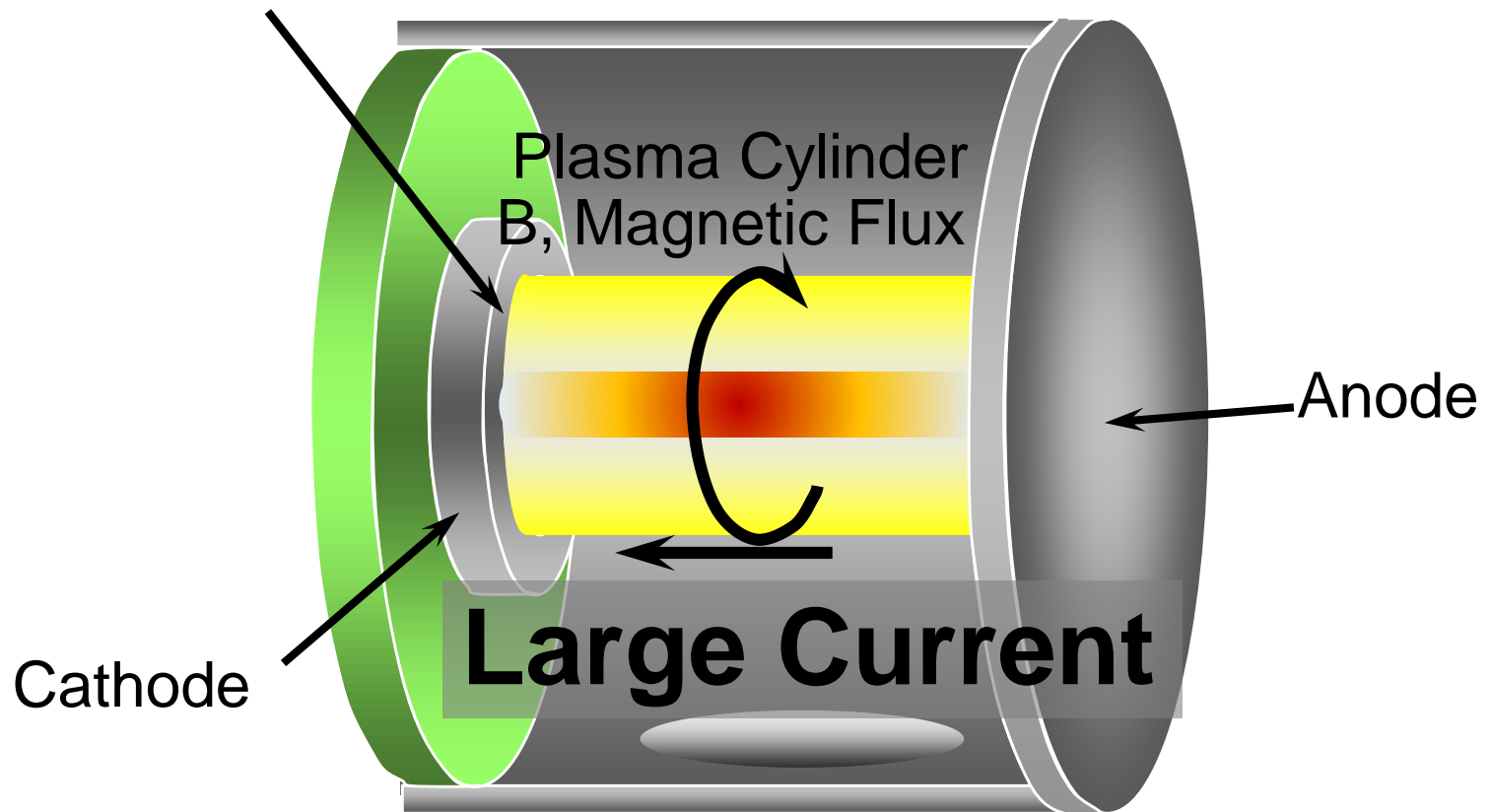
Trigger system
(not shown, approximate placement)

Spark gap switch system

Notional Z-Pinch Target

Vaporized Wire Array

Evacuated Chamber



Pulsed Fission Fusion Hybrid (PUFF)

- Fuel

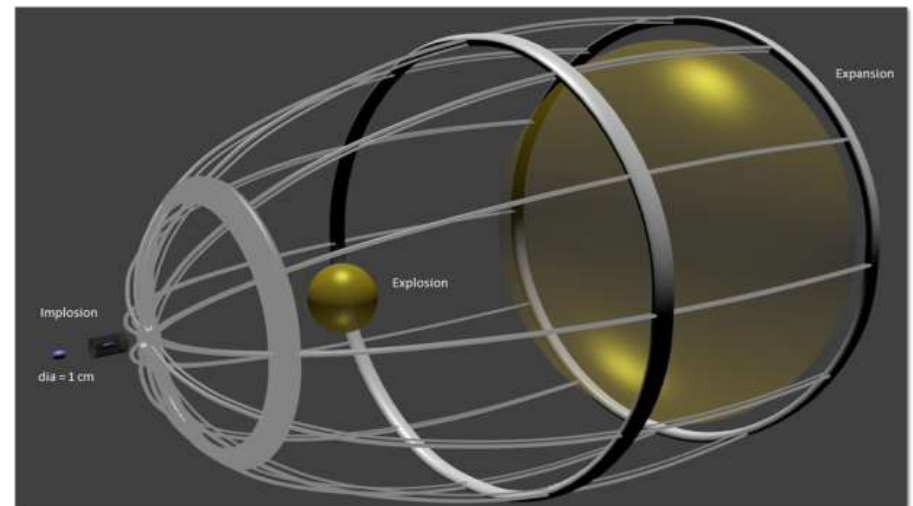
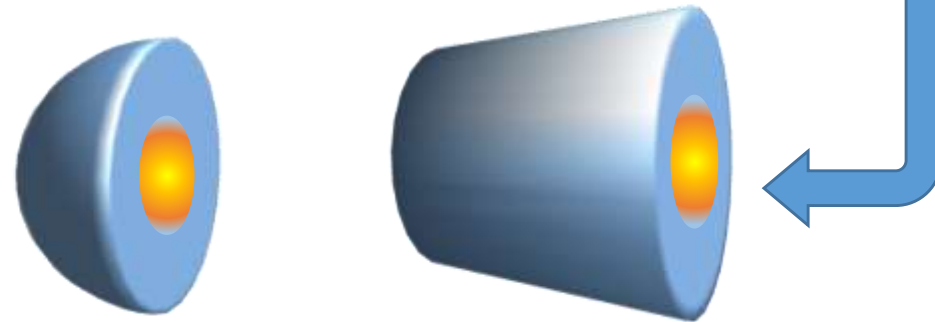
- Fission liner (^{238}U ^{232}Th)
- ^6Li D or D-T center
- Initial neutron source
 - Spontaneous (AmBe)
 - Fusion (DT, DD)

- Geometry

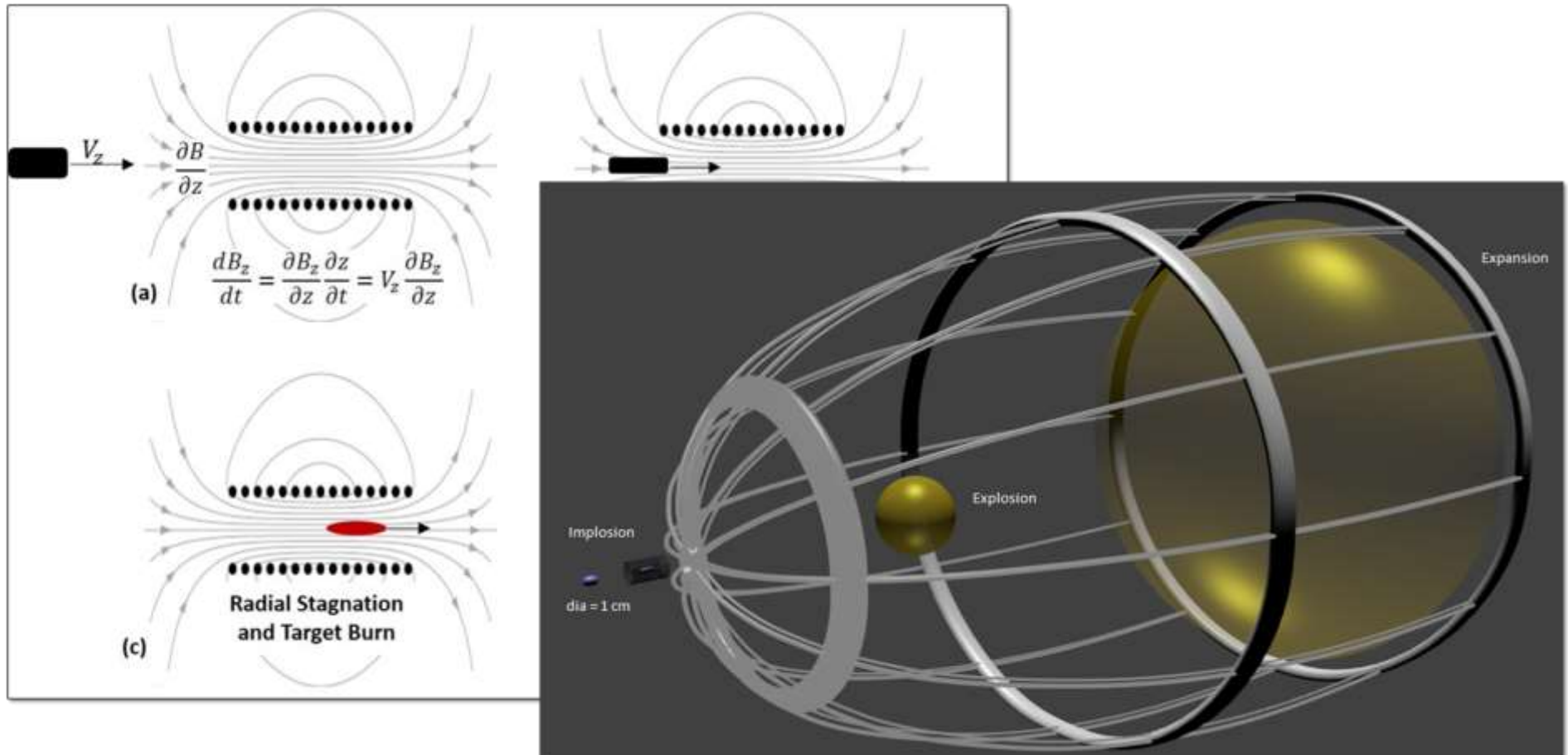
- spherical
- Cylindrical

- Physics models required

- Radiation/matter interactions
- Fast neutron fission
- Electromagnetic fields
- Equations of state with ionization and compression of solids

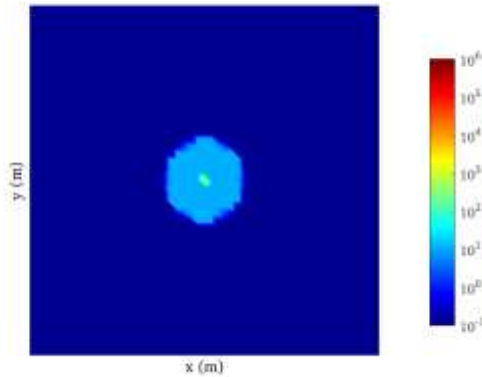


Gradient Field Imploding Liner Fusion Propulsion

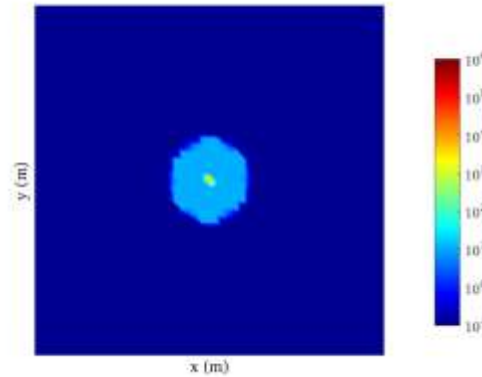


Temperature slice, center of target, $t=0$ ns

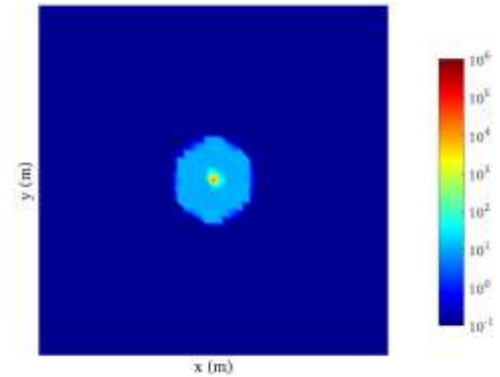
T = 2 keV



T = 5 keV

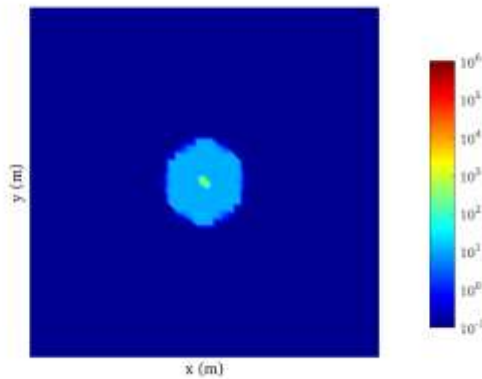


T = 10 keV

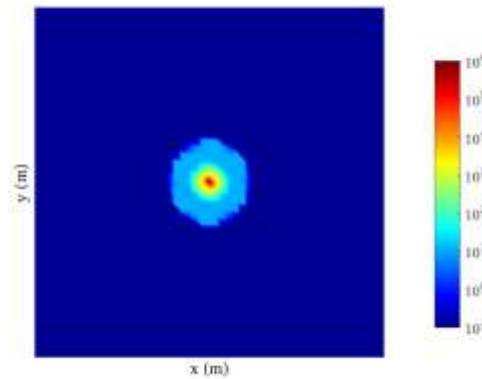


Temperature slice, center of target, $t=50$ ns

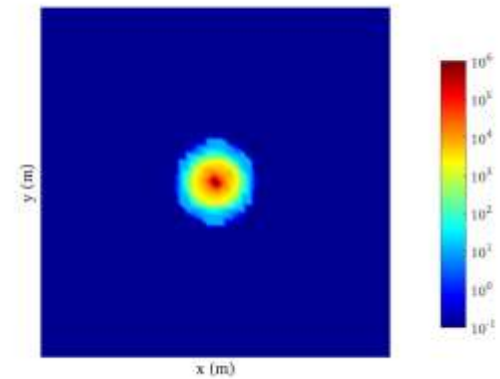
T = 2 keV



T = 5 keV

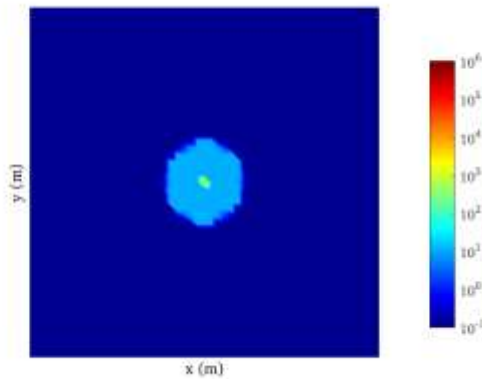


T = 10 keV

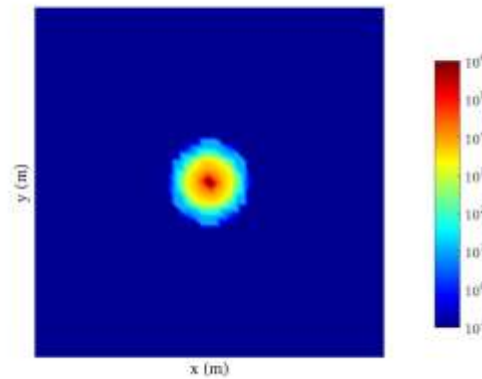


Temperature slice, center of target, $t=100$ ns

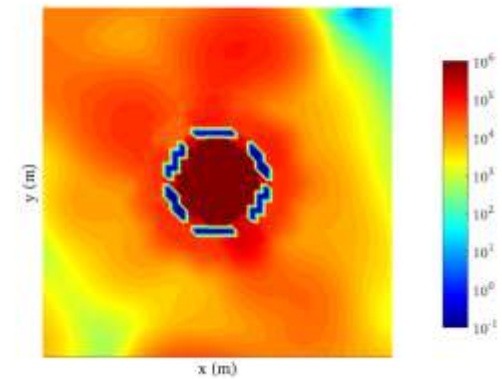
T = 2 keV



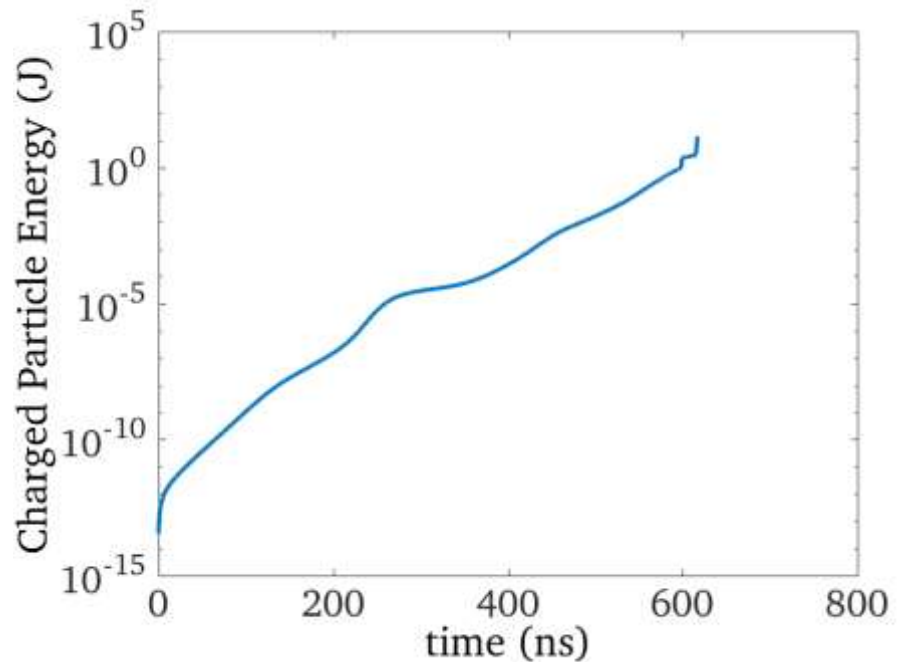
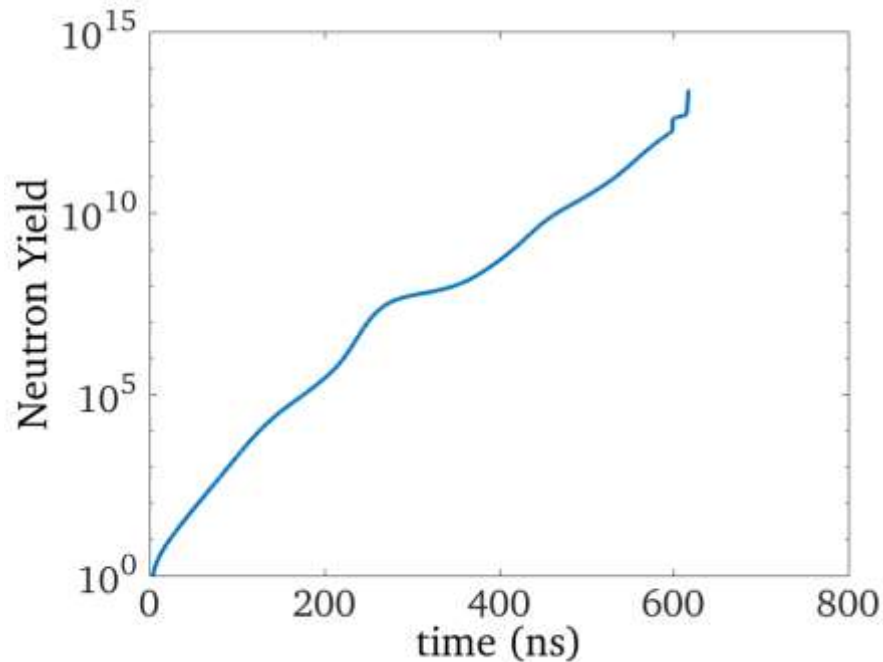
T = 5 keV



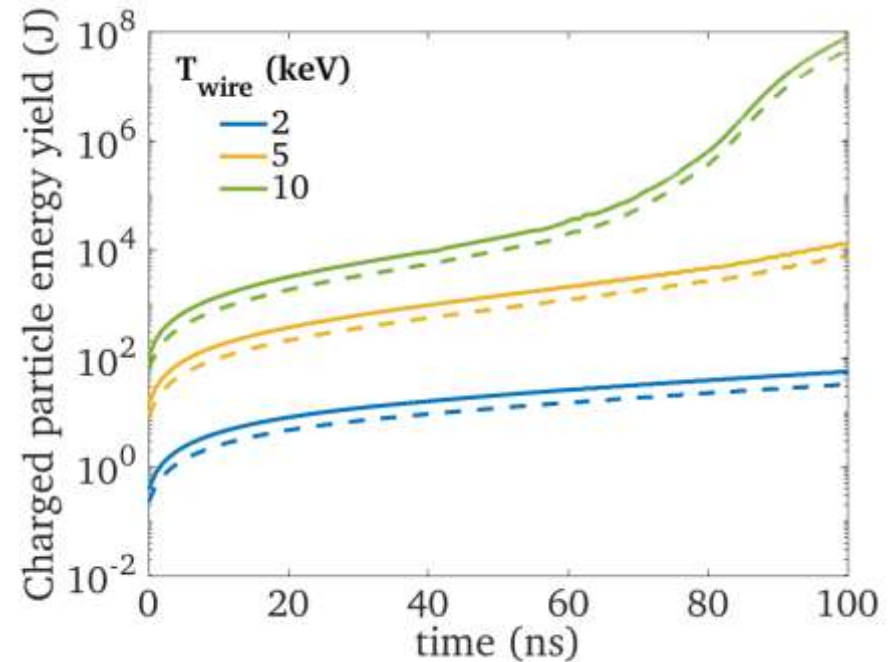
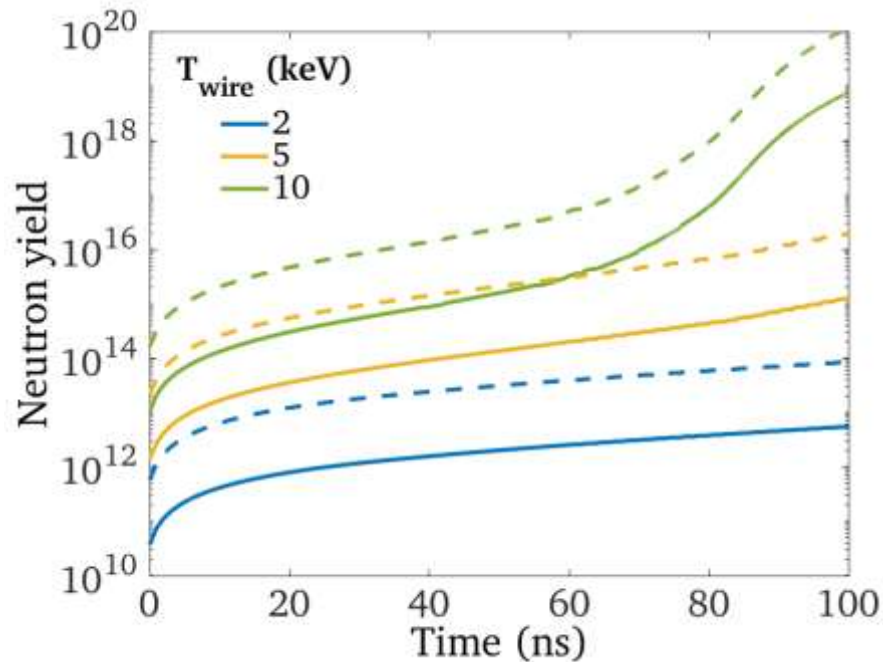
T = 10 keV



Neutron yield and charged particle energy yield vs time, prior to code crashing



High temperature wire leads to burn in secondary fusion liner at 70 ns assuming 10 keV ${}^6\text{Li}$ D wire temperature. Fission and fusion reactivity are tightly coupled.

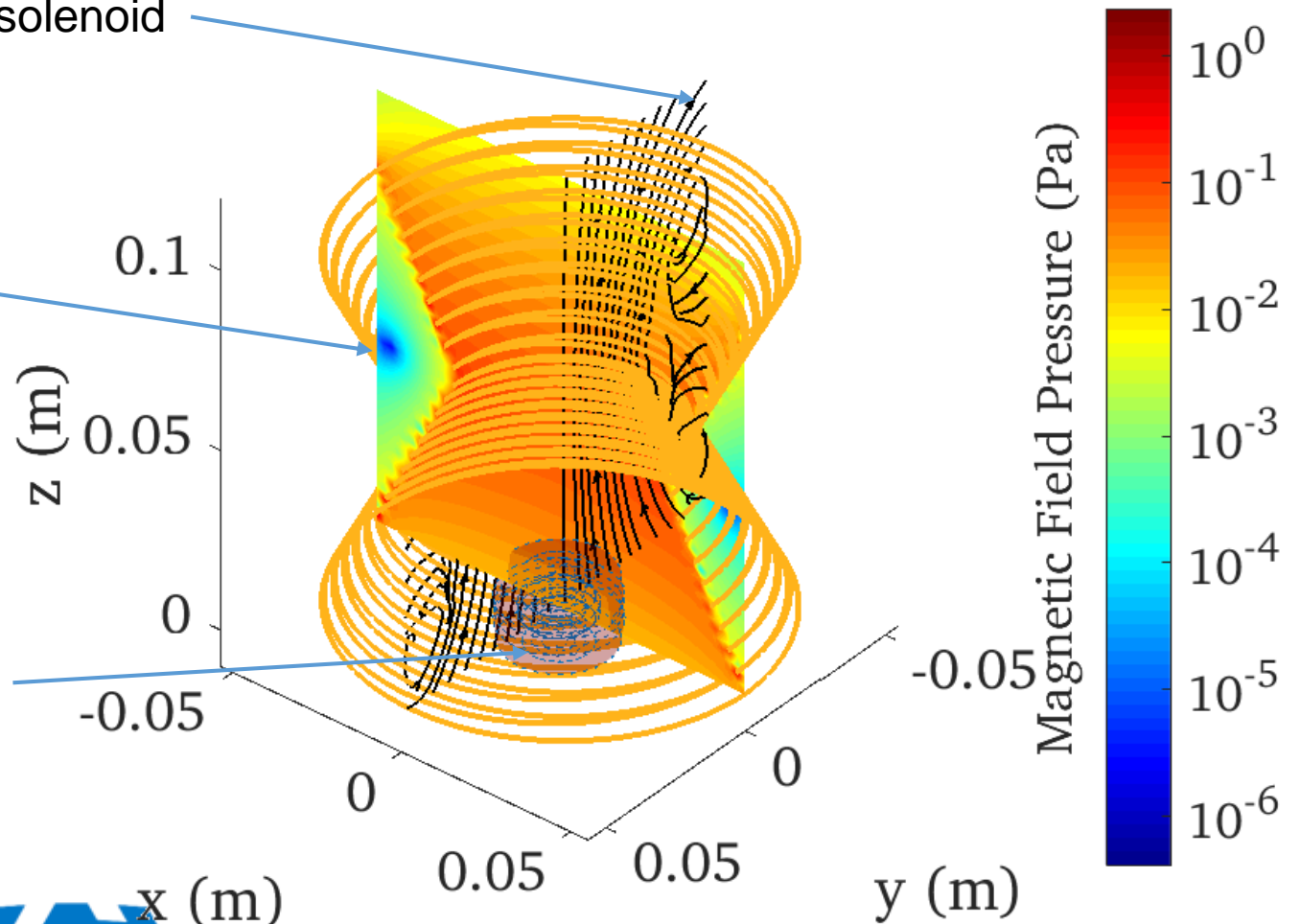


Gradient Field Fusion Propulsion system simulation

Magnetic field lines inside solenoid

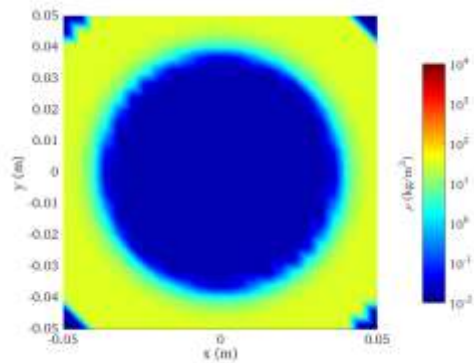
Magnetic field pressure

Fuel pellet entering mouth
Of nozzle at several km/s

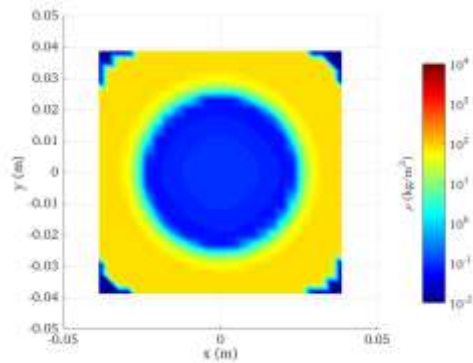


Mass Density

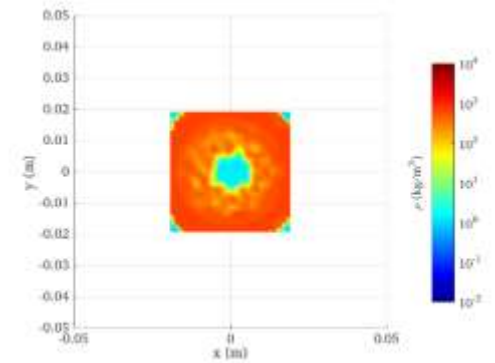
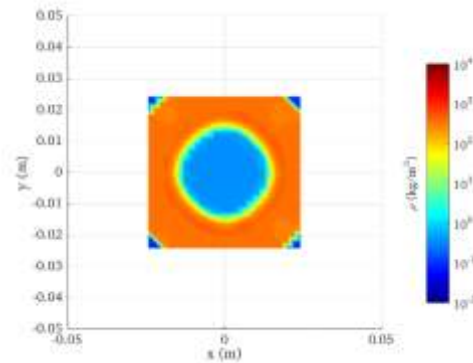
t = 0 ns



t = 250 ns

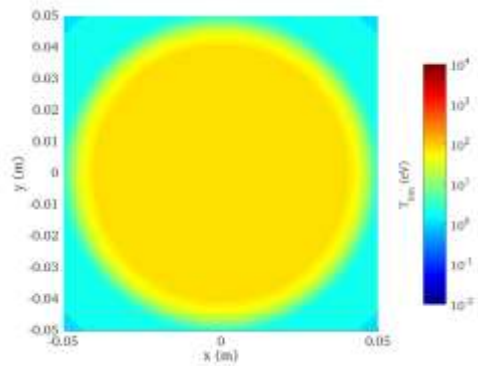


t = 500 ns

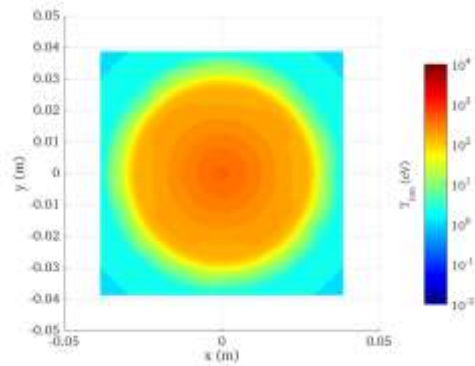


Ion Temperature

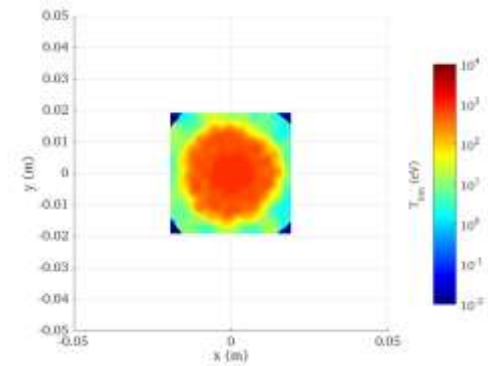
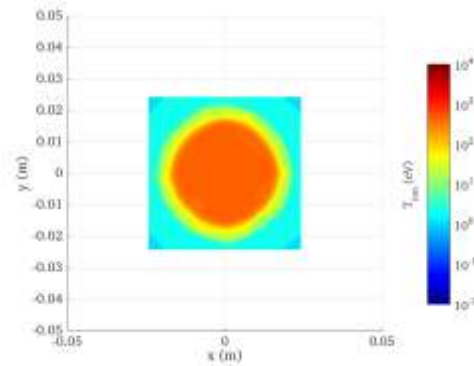
t = 0 ns

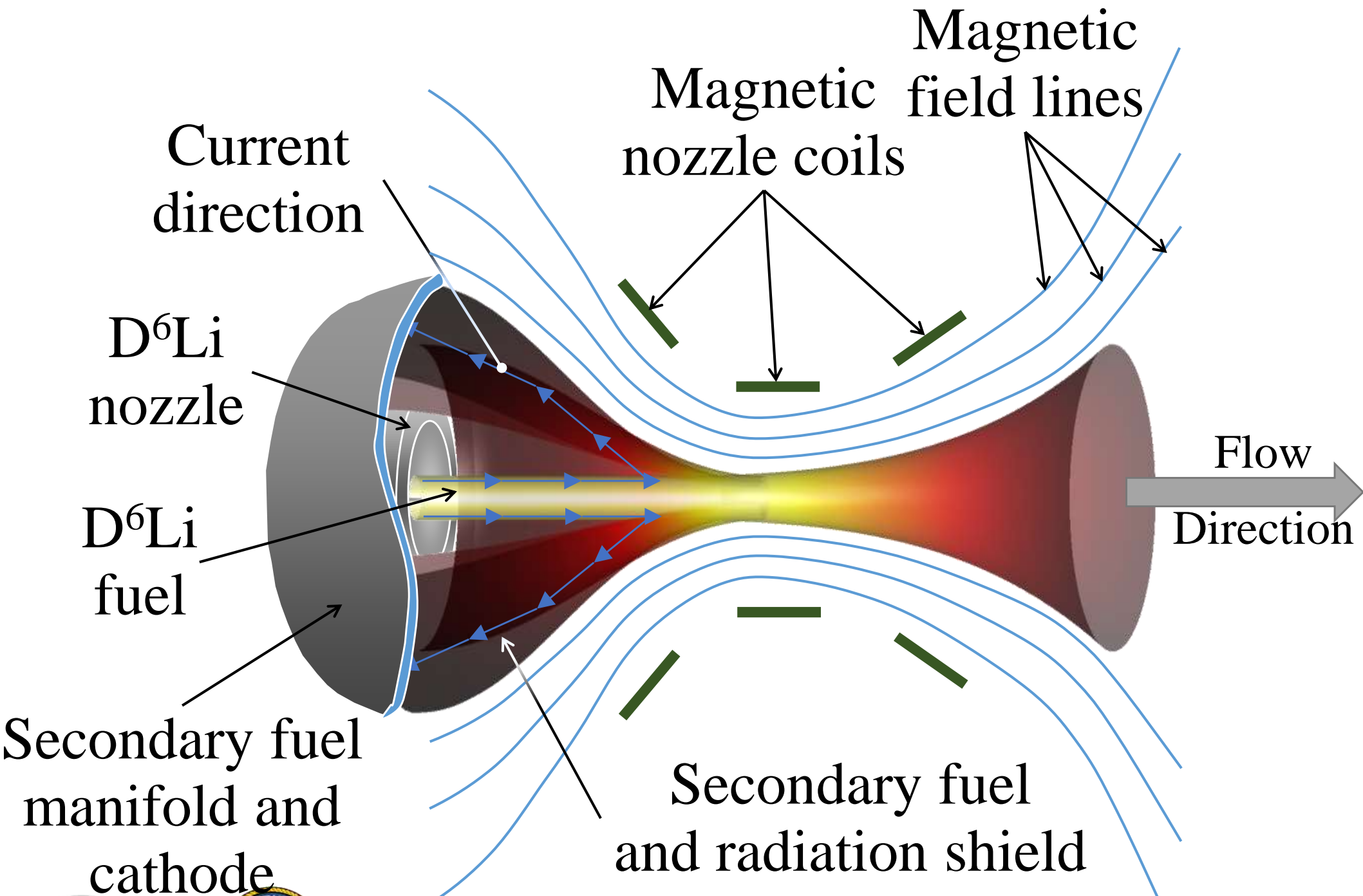


t = 250 ns



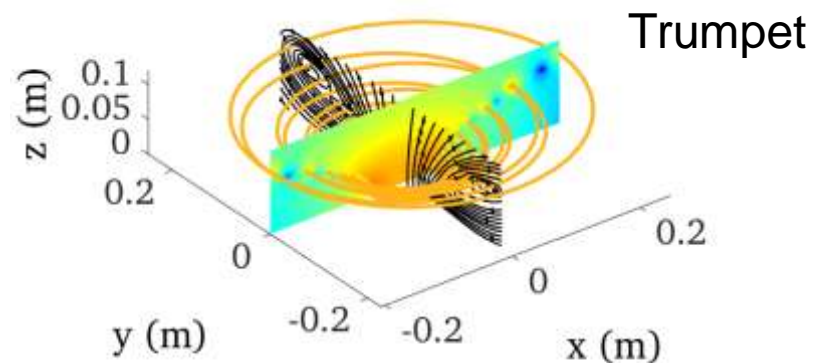
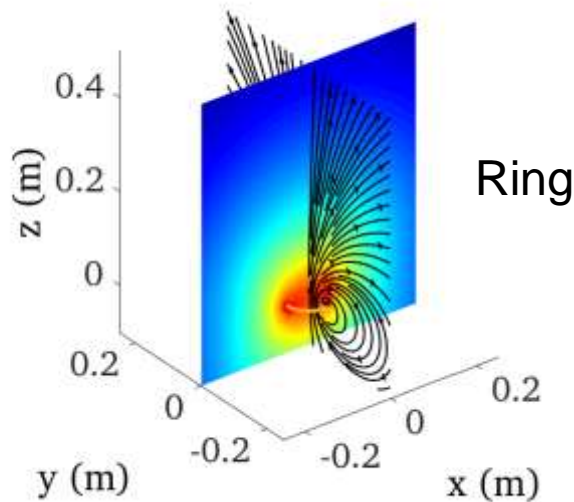
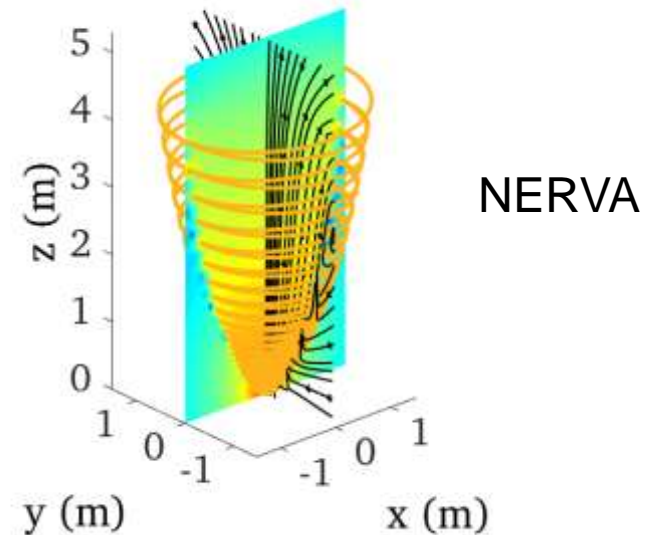
t = 500 ns



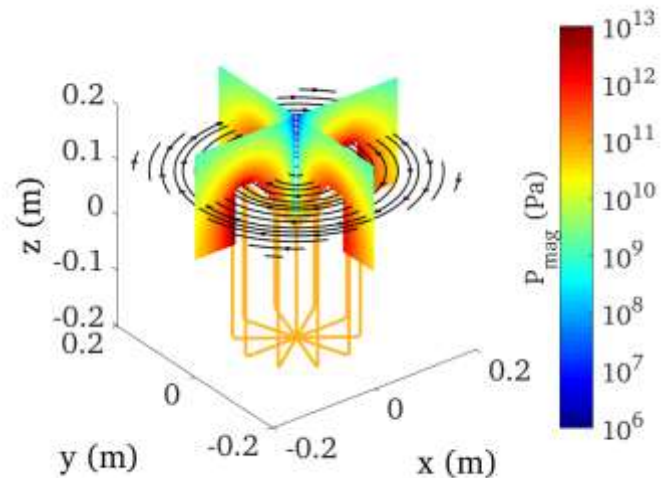
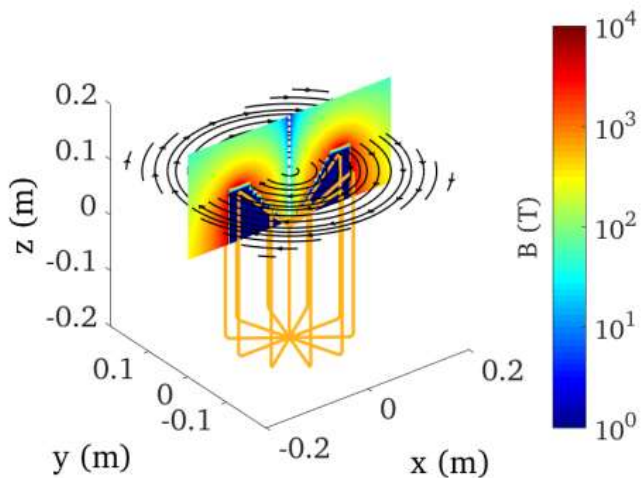
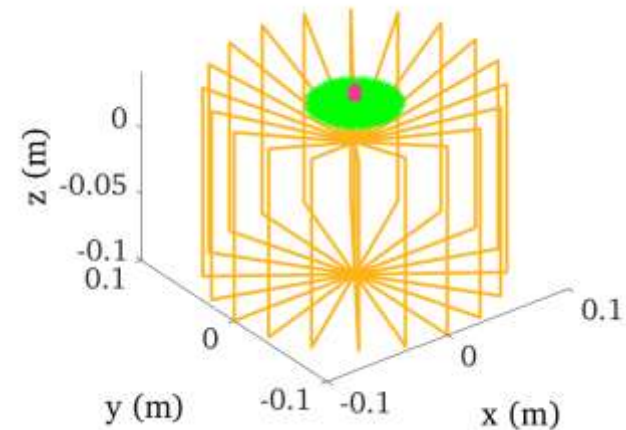
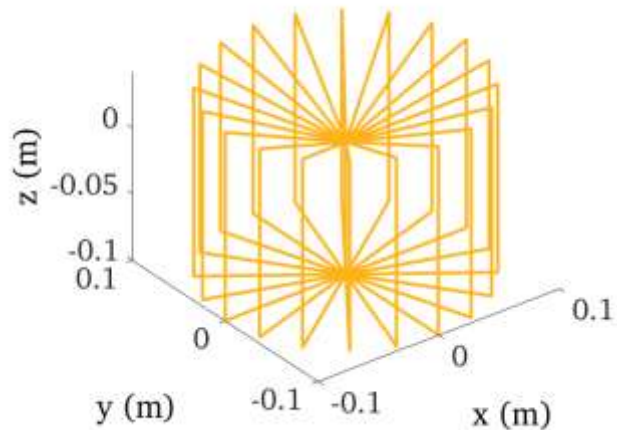


Simulation attempts in magnetic topology

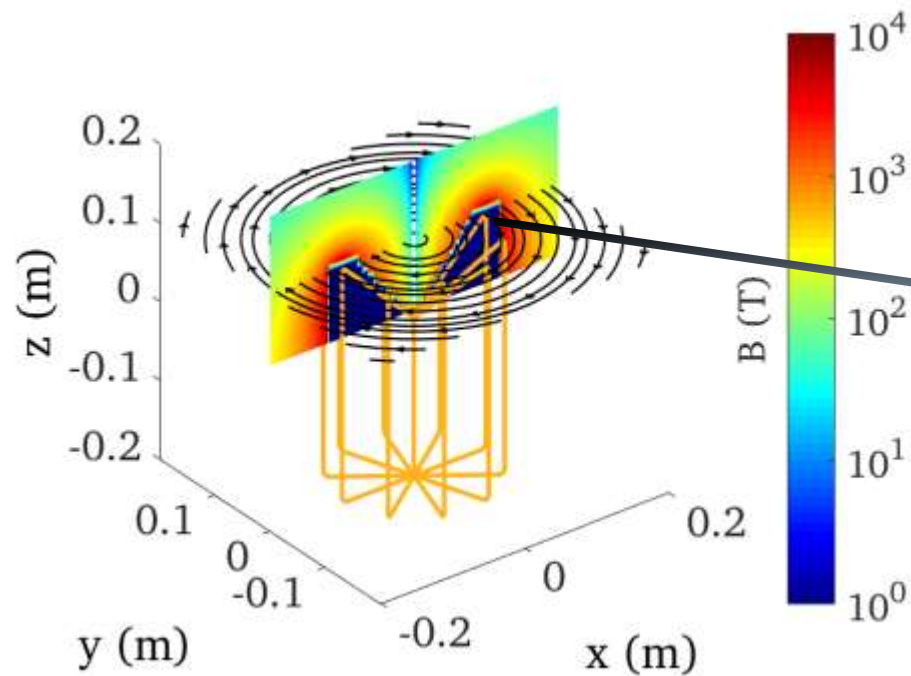
- Solenoidal winding variations
 - NERVA nozzle shape
 - Ring
 - Bell (like end of trumpet)
- Initial plasma placement
- Key parameter is finding topology that gives $\mathbf{j} \times \mathbf{B}$ Lorentz force in axial direction during expansion



Departure from solenoid to longitudinal windings provided positive results because of nearly azimuthal magnetic field generated.

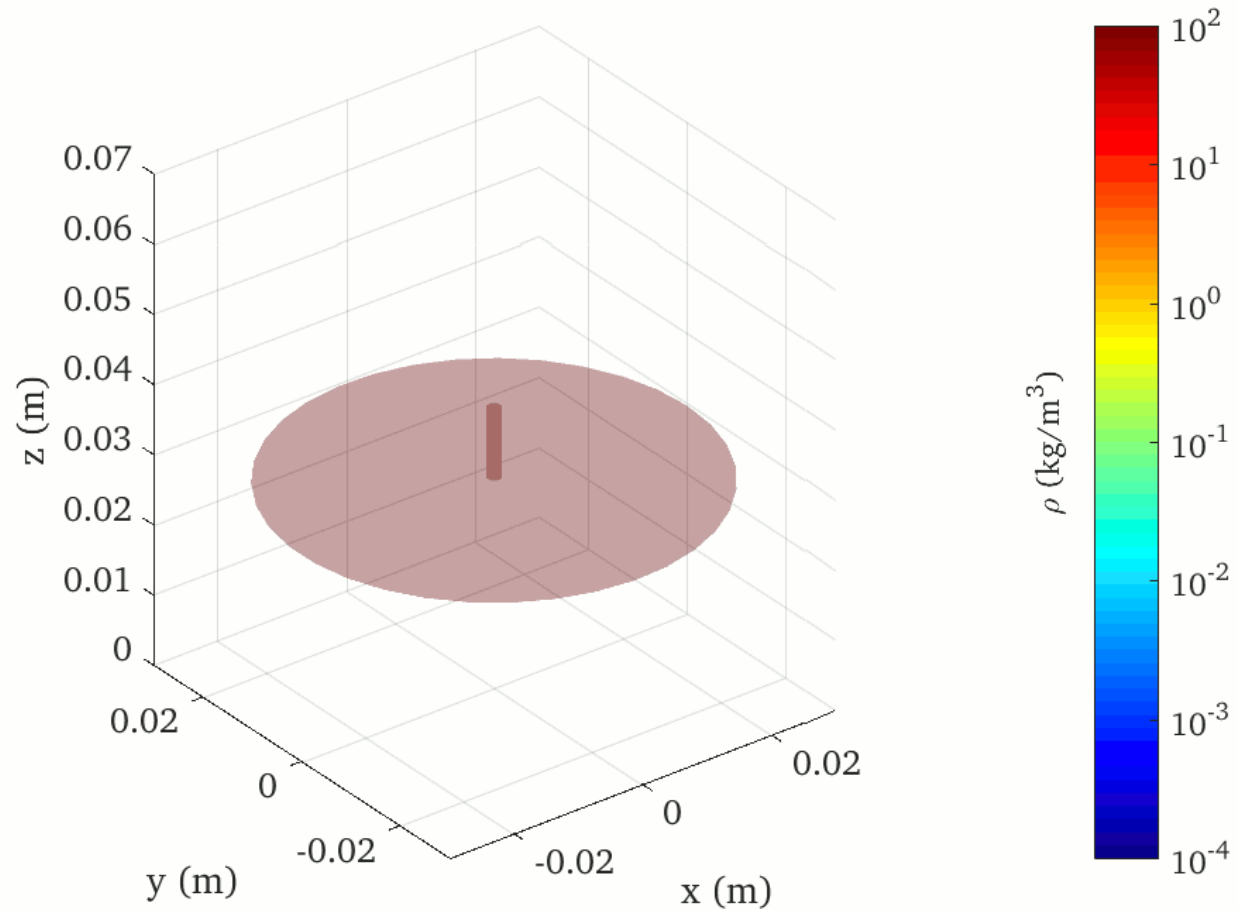


Anything new was probably invented in the last century...



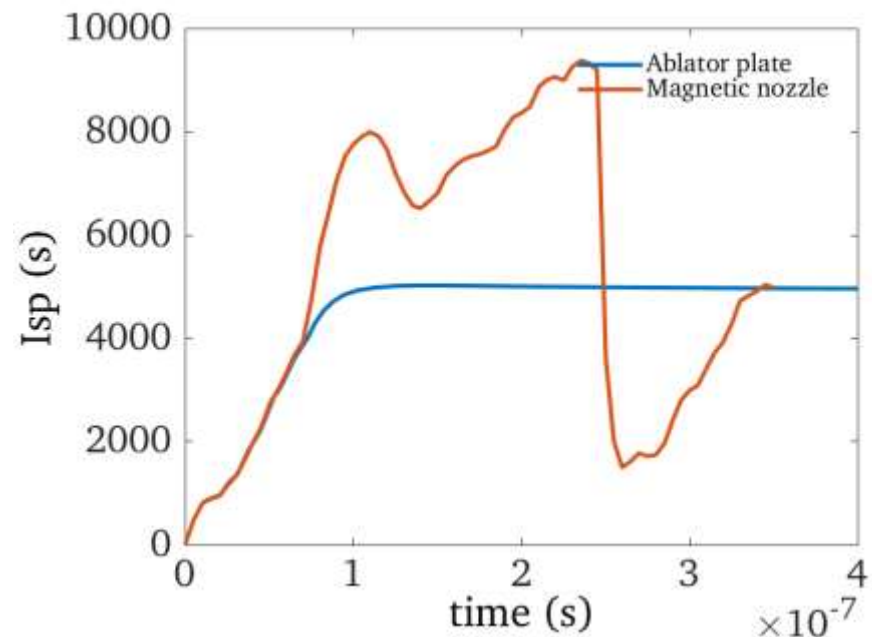
Early results for a 1 keV DT plasma

- Thermal expansion against plate provides most of the impulse
- Induced current on the surface excludes the externally provided flux from the nozzle and partially redirects the flow



Comparison of pure ablator plate and with nozzle

- Presence of field provides additional redirection of thermally expanded plasma
- Improvements are seen against pure ablator plate with no field
- At 250 ns, the circuit model crashes (next task is to explore why)
- Rapid tapering off of Isp with time during expansion caused by radiation cooling



Summary of vehicle performance parameters for very deep space missions

Destination	Trip time (years)	Isp (s)	IMLEO (metric tons)	Thrust (N)	Shot frequency (1/s)	Mass flow rate (mg/s)	ΔV (km/s)
Gravitational Lensing	10	9.4×10^4	15.3	0.66	2.9	0.72	392
Alpha Centauri	269	2.7×10^5	1000	1.39	8.8	0.48	7,280

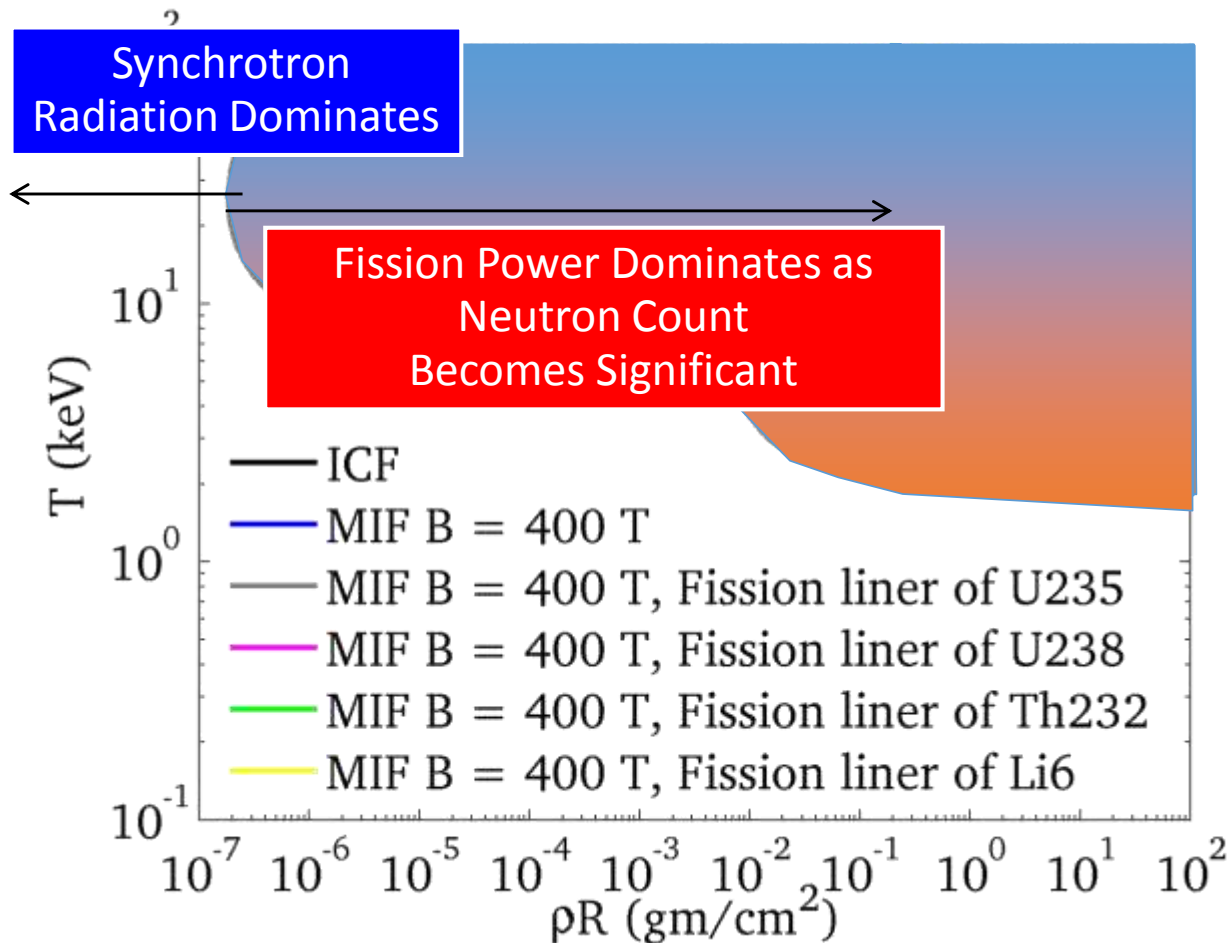


Concluding remarks

- Fusion and Fission/Fusion hybrid concepts being studied to enable rapid interplanetary space flight for human piloted and science missions.
- The need for reliable restarts requires a nuclear electric system for deep space travel
- The low thrust compromises trip times due to the slow spiral out of Earth's gravity well
- A bimodal approach has been proposed leveraging high thrust from an NTP system for rapid departure and high specific impulse for rapid interplanetary space travel



Fusion Power Balance



- Parameter space for ignition
- Greatly broadened with embedded magnetic field
- Marginally improved with ⁶Li and thorium liners
- Significantly enhanced with uranium liners (²³⁵U and ²³⁸U)



Why Send Humans to Explore Other Planets?

(Additional arguments made by Bill Gerstenmeier)

- Knowledge
 - Human exploration inspires people to seek knowledge
 - Through scientific discovery we increase our understanding of the world
- Economic Growth
 - Creation of industries
 - Job growth
 - Demand for a highly skilled workforce
- A Better Future
 - Advancing American leadership
 - Creating a path for peace, diplomacy, and global cooperation





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