Laser-Driven Magnetized Collisionless Shocks

Derek Schaeffer

Princeton University/PPPL

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Collaborators

- A. Bhattacharjee, PU, PPPL
- W. Fox, PPPL
- G. Fiksel, U. Michigan
- D. Haberberger, LLE
- D. Barnak, LLE
- S. X. Hu, LLE
- K. Germaschewski, U. NH
- C. Niemann, UCLA
- C. Constantin, UCLA
- E. Everson, UCLA
- A. Bondarenko, UCLA
- S. E. Clark, UCLA
- P. Heuer, UCLA
- D. Winske, LANL
- D. Larson, LLNL















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- Brief introduction to laboratory research on collisionless shocks driven by magnetic pistons
- Recent measurements of the formation and evolution of a high-M_A shock on the Omega EP laser facility
- Results on the formation and structure of low- M_A shocks on the LAPD
- Dimensionless criteria that describe the conditions under which shocks driven by magnetic pistons form



Collisionless Shocks are Prevalent in Many Space and Astrophysical Systems





Supernovae Remnants







Active Galactic Nuclei



- Collisionless shocks convert the ram pressure of incoming supersonic flows to thermal pressure over length scales much shorter than the collisional mean free path
- Shocks are interesting!
 - Complex nonlinear and time-dependent systems
 - Rich phenomenology sensitive to ambient conditions
 - Associated with high energy particle acceleration
- Shocks classification:
 - Electrostatic
 - Turbulent (i.e. Weibel-mediated)
 - Magnetized

Romagnani, *et al.*, PRL, 2008 Kuramitsu, *et al.*, PRL, 2011 Haberberger, *et al.*, Nature, 2011 Fox, *et al.*, PRL, 2013 Huntington, *et al.*, Nature, 2015

Laboratory Experiments can Reproduce the Physics of Space and Astrophysical Collisionless Shocks in a Controlled Setting

- Spacecraft very successful but largely limited to 1D datasets and pre-formed shocks
- Early laboratory experiments also successful but of limited relevance to space shocks
 - The shocks were limited to M<6
 - The goal of the experiments was primarily shock heating (in particular electrons)
 - The experiments were limited to strictly perpendicular magnetic geometries
 - The available diagnostics were also limited to 1D datasets
- A new class of collisionless shocks experiments that utilize a laser-driven magnetic piston is now available
 - Wide range of Mach numbers (M<40)
 - 2D and 3D datasets
 - Quasi-perpendicular and quasi-parallel magnetic geometries

Russel, ASR, 1995

Paul, *et* al., Nature, 1965 Kurtmullaev, *et* al., J. App. Mech. Tech. Phys., 1965 Goldenbaum, *et al.*, PoF, 1967 Stamper, *et* al., PoF, 1969



Drake, PoP, 2000

Laser-Driven Diamagnetic Cavity Acts as Piston, Launching Shocks in Ambient Plasma



 Piston plume drives diamagnetic current, creating magnetic cavity and leading magnetic compression

Wright, PoF, 1971 Vanzeeland, *et al.*, PoP, 2004 Schaeffer, *et al.*, PoP, 2012

Piston Energy and Momentum Transferred Collisionlessly to Ambient Plasma





Berezin, *et al.*, Intl. J. Comp. Fluid Dyn., 1998 Hewett, *et al.*, JGR, 2011 Bondarenko, *et al.*, submitted, 2016

Dissipation in Subcritical (M_A<3) Shocks Provided by Drift Instabilities



Treumann, AAR, 2009 Balogh and Treumann, *Physics of Collisionless Shocks*, 2013

Dissipation in Supercritical (M_A>3) Shocks Provided by Ion Reflection



Treumann, AAR, 2009 Balogh and Treumann, *Physics of Collisionless Shocks*, 2013

Shock Physics Indicates What Experimental Features Are Necessary

- Super-Alfvénic (M_A > 1)
 - Shock speed must be greater than upstream Alfvén speed
- Collisionless
 - Collisional mean free path > system size
- Large density and magnetic compressions
 - n/n_0 and $B/B_0 > 2$
- Steep compression widths
 - Density/magnetic jump \gtrsim ion inertial length
- Shock separates from piston
 - Shock features should be distinct from piston

Experimental Setup for Quasi-Perpendicular Shocks on Omega EP



- MIFEDS coils provide background magnetic field ~ 8T
- Heater beam ablates ambient plasma (n_{i,0}≈10¹⁸) 12 ns before drive beams
- Drive beams create supersonic piston plumes that expand into ambient plasma
- Diagnostics:
 - Angular Filter Refractometry (AFR)
 - Shadowgraphy
 - Proton radiography

AFR Diagnostic Measures Density Gradients



Null Shots Show No Shock Features



 Without background magnetic field or ambient plasma, only piston plumes observed



Propagating Density and Magnetic Compressions Observed with $B_0 \neq 0$ and $n_0 \neq 0$



Propagating Density and Magnetic Compressions Observed with $B_0 \neq 0$ and $n_0 \neq 0$



- Expanding at 700 km/s, yielding $M_{ms} \approx 12$
- Density compression **n**/**n**₀ ~ 3-4
- Magnetic compression B/B₀ ~ 2-4
- Compression width Δ > 1 c/ $\omega_{\rm pi}$



2D PIC Simulations Indicate Formation of High-M_A Shock



- CH piston plasma expanding into CH ambient plasma embedded in magnetic field
- Piston ions sweep out ambient ions and magnetic field
- At early times, this leads to the formation of a H shock, mixed with piston ions and the beginnings of a C shock
- At later times, a separate C shock forms behind the H shock, and the piston ions become trapped behind the C shock

2D PIC Simulations Indicate Formation of High-M_A Shock



- The formation of a shock leads to a double "bump" density profile, corresponding to the leading shock and the trapped piston ions
- The density compression associated with the shock is n/n₀~4
- The magnetic compression associated with the shock is B/B₀>4
- The width of the density compression $\Delta \sim 1 \rho_i$

Data Profiles Show Density Evolution that is Consistent with High- M_A Shock Formation



 Early time density compression mostly associated with pile-up of piston ions

Schaeffer, et al., submitted, 2016

 At late time clear double bump feature associated with shock and trapped piston

The UCLA Facility Uniquely Combines a kJ-Laser and a Large Magnetized Plasma

Plasma Length	18 m
Plasma Diameter	60 cm
Electron Density (Main)	2x10 ¹² cm ⁻³
Electron Density (LaB ₆)	2x10 ¹³ cm ⁻³
Electron Temp.	5 eV
lon Temp.	< 1 eV
B Field	0.2 – 1.8 kG
Rep. Rate	1 Hz



- Well-characterized ambient plasma
- Plasma large-enough for shock-piston separation → Shock physics decoupled from driver
- Flexible laser geometry: quasi-perpendicular, oblique, or quasi-parallel
- Large scale & low pressure ambient allows localized probe measurements

Goal: investigate shock formation, debris-ambient coupling, ion dynamics, etc.



Raptor Laser

Wavelength	1053 nm
Pulse Length (FWHM)	25 ns
Energy per Pulse	300 J +
Rep. Rate	45 min

Experimental Setup for Quasi-Perpendicular Shocks in the LAPD



Mach ~2 Collisionless Shock Observed Separating from Piston



Simulations Show Low-Mach Number Shock Formation



Criteria Necessary for Shock Formation

- Super-Alfvénic $(M_A > 1)$
 - Shock speed must be greater than upstream Alfvén speed
- Collisionless $(\lambda_{ii}/D_0 > 1)$
 - Collisional mean free path > system size
- Sufficient experimental size $(\rho/D_0 < 1)$
 - Ion gyro-orbits must fit within the experiment
- Sufficient piston magnetization $\left(\rho_p/R_m < \sqrt{2}\right)$
 - Piston ions must have enough energy to accelerate ambient ions
 - R_m is the characteristic length scale for piston coupling
- Sufficient ambient magnetization $(\rho_a/R_* < 1)$
 - Ambient ions must be sufficiently accelerated
 - R* is the characteristic length scale for ambient coupling

Are these criteria valid?

$$R_m = \left(3N_0 m_p / 4\pi m_a n_a\right)^{1/3}$$

$$R_* = (3N_0Z_p/4\pi Z_a n_a)^{1/3}$$

Bashurin, et al., J. App. Mech. & Tech. Phys., 1983 Drake, PoP, 2000 Clark, et al., PoP, 2013 Larson, et al., JRERE, 2016

Shock Experiments can be Predictively Organized by Formation Criteria



Shock Experiments can be Predictively Organized by Formation Criteria





- We have observed for the first time the formation and evolution of a laser-driven, high-M_A (supercritical) collisionless shock. The results agree well with 2D PIC simulations.
- We have launched low-M_A (subcritical) magnetized shocks and measured their formation and structure. Hybrid simulations reproduce the basic features observed in experiments.
- Shock experiments spanning large parameter ranges reveal that pistondriven shocks can be predictively organized by dimensionless formation criteria.

Future Directions

- High-M_A shocks and particle acceleration
 - How are particles injected into the shock acceleration process?
- Low-M_A shocks and detailed shock structure
 - What dissipation mechanisms are at work in the shock layer?
- Varying magnetic geometry
 - Can quasi-parallel shocks be generated in the laboratory?
- Shocks and other highly-driven (i.e. lasers) systems
 - What, if any, interaction do shocks have with magnetic reconnection and turbulence, such as the Weibel instability?