LECTURES 1 - 10 INTRODUCTION TO CLASSICAL MECHANICS

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6.3.3 Momentum in the CM frame

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7.1 Two-body collisions - general concepts

- 7.1.1 Momentum exchange and impulse
- 7.1.2 An off-axis collision in 2D

7.2 Elastic collisions in the Lab frame

7.2.1 Elastic collisions in 1D in the Lab frame7.2.2 Special case in 1D where target particle is at rest7.2.3 Collision in 2D : equal masses, target at rest

8.1 Elastic collisions in the CM frame

8.2 Lab to CM : 2-body 1D elastic collision

8.2.1 Collision in 1D : numerical example

8.3 Relationship between speeds in CM in 2D

8.4 Lab to CM : 2-body 2D elastic collision

9.1 Examples of 2D elastic collisions

9.1.1 Example 1: Equal masses, target at rest

9.1.2 Example 2: Elastic collision, $m_2 = 2m_1$, $\theta_1 = 30^{\circ}$

9.2 Inelastic collisions in the Lab frame in $1D(u_2 = 0)$

9.2.1 Coefficient of restitution

9.3 Inelastic collisions viewed in the CM frame
9.3.1 Kinetic energy in the CM : alternative treatment
9.3.2 Coefficient of restitution in the CM
9.3.3 Example of inelastic process

10.1 Resisted motion and limiting speed

10.2 Air resistance

10.3 Example 1 : Resistive force, $F_R \propto v$

10.4 Example 2: Resistive force, $F_R \propto v^2$ 10.4.1 Work done on the body by the force for $F_R \propto v^2$

1.1 Outline of lectures

Two groups of lectures

- ▶ 10 in MT mostly 1D & 2D linear motion.
- 19 in HT 3D full vector treatment of Newtonian mechanics, rotational dynamics, orbits, introduction to Lagrangian dynamics

Info on the course is on the web:

http://www.physics.ox.ac.uk/users/harnew/lectures/

- Synopsis and suggested reading list
- Problem sets
- Lecture slides

1.2 Book list

- Introduction to Classical Mechanics A P French & M G Ebison (Chapman & Hall)
- Introduction to Classical Mechanics D. Morin (CUP) (good for Lagrangian Dynamics and many examples).
- Classical Mechanics : a Modern Introduction, M W McCall (Wiley 2001)
- Mechanics Berkeley Physics Course Vol I C Kittel et al. (McGraw Hill)
- Fundamentals of Physics Halliday, Resnick & Walker (Wiley)
- Analytical Mechanics 6th ed, Fowles & Cassidy (Harcourt)
- Physics for Scientists & Engineers, (Chapters on Mechanics) P.A Tipler & G. Mosca (W H Freeman)
- Classical Mechanics T W B Kibble & F H H Berkshire (Imperial College Press)

1.3 What is Classical Mechanics?

Classical mechanics is the study of the motion of bodies in accordance with the general principles first enunciated by Sir Isaac Newton in his Philosophiae Naturalis Principia Mathematica (1687). Classical mechanics is the foundation upon which all other branches of Physics are built. It has many important applications in many areas of science:

- Astronomy (motion of stars and planets)
- Molecular and nuclear physics (collisions of atomic and subatomic particles)
- Geology (e.g., the propagation of seismic waves)
- Engineering (eg structures of bridges and buildings)

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Classical Mechanics covers:

- The case in which bodies remain at rest
- Translational motion- by which a body shifts from one point in space to another
- Oscillatory motion- e.g., the motion of a pendulum or spring
- Circular motion-motion by which a body executes a circular orbit about another fixed body [e.g., the (approximate) motion of the earth about the sun]
- More general rotational motion–orbits of planets or bodies that are spinning
- Particle collisions (elastic and inelastic)

Forces in mechanics

Relative magnitude of forces:

- Strong force nuclear : \sim 1
- Electromagnetism charged particles : ¹/₁₃₇
- Weak force β decay : $\sim 10^{-5}$
- $\blacktriangleright\,$ Gravitational important for masses, relative strength : $\sim 10^{-39}$

Not too fast!

Classical Mechanics valid on scales which are:

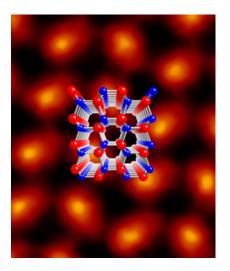
- Not too fast
- eg. high energy particle tracks from CERN
- v << c [speed of light in vacuo]
- If too fast, time is no longer absolute - need special relativity.



Not too small!

Classical Mechanics valid on scales which are:

- Not too small!
- Images of atom planes in a lattice by scanning tunneling electron microscope
- Particles actually have wave-like properties :
 λ = h/p (h = 6.6 × 10⁻³⁴ Js)
- Hence for scales >> λ, wave properties can be ignored

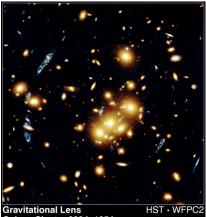


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Not too large!

Classical Mechanics valid on scales which are:

- Not too large!
- Gravitational lens produced by a cluster of galaxies
- Space is "flat" in classical mechanics - curvature of space is ignored
- Also in Newtonian mechanics, time is absolute



Galaxy Cluster 0024+1654 PRC96-10 · ST Sci OPO · April 24, 1996 W.N. Colley (Princeton University), E. Turner (Princeton University), J.A. Tyson (RTST Bell Labs) and NaSA

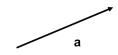
1.4 Vectors in mechanics

The use of vectors is essential in the formalization of classical mechanics.

- A scalar is characterised by magnitude only: energy, temperature.
- A vector is a quantity characterised by magnitude and direction: eg. Force, momentum, velocity.

Notation:

► Vector: **a** (bold); in components $\underline{\mathbf{a}} = (a_x, a_y, a_z)$



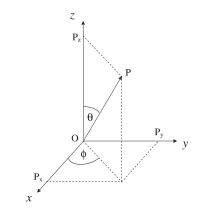
- Magnitude of $\underline{\mathbf{a}}$ is $|\underline{\mathbf{a}}|$ or simply a.
 - Two vectors are equal if they have the same magnitude and direction (i.e. parallel)

$$\underline{\mathbf{a}} = \underline{\mathbf{b}}$$
 gives $a_x = b_x$, $a_y = b_y$, $a_z = b_z$

1.4.1 Vector components in 3D

Projecting the components: $\underline{\mathbf{p}} = (p_x, p_y, p_z)$

- x-component $p_x = |\underline{\mathbf{p}}| \sin(\theta) \cos(\phi)$
- ► y-component $p_y = |\mathbf{p}| \sin(\theta) \sin(\phi)$
- z-component $p_z = |\underline{\mathbf{p}}| \cos(\theta)$
- Magnitude $|\underline{\mathbf{p}}| = \sqrt{(p_x^2 + p_y^2 + p_z^2)}$
- Direction $\tan(\phi) = (p_y/p_x)$ $\cos(\theta) = (p_z/|\underline{\mathbf{p}}|)$

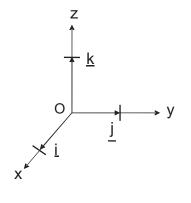


1.4.2 Unit vectors

- A unit vector is a vector with magnitude equal to one.
- e.g. Three unit vectors defined by orthogonal components of the Cartesian coordinate system:

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$$\underline{\mathbf{j}} = (0,1,0), |\underline{\mathbf{j}}| = 1$$

- $\mathbf{\underline{k}} = (0,0,1), \ \overline{|\mathbf{\underline{k}}|} = 1$
- A unit vector in the direction of general vector <u>a</u> is written <u>â</u> = <u>a</u>/|<u>a</u>|
- <u>a</u> is written in terms of unit vectors <u>a</u> = $a_x \underline{i} + a_y \underline{j} + a_z \underline{k}$



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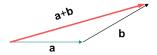
1.4.3 Vector algebra

Sum of two vectors

- To calculate the sum of two vectors <u>c</u> = <u>a</u> + <u>b</u> Triangle rule: Put the second vector nose to tail with the first and the resultant is the vector sum.
- $\mathbf{\underline{c}} = \mathbf{\underline{a}} + \mathbf{\underline{b}}$: in (x, y, z) components $(c_x, c_y, c_z) = (a_x + b_x, a_y + b_y, a_z + b_z)$

Alternatively
$$\underline{\mathbf{c}} = \underline{\mathbf{a}} + \underline{\mathbf{b}}$$

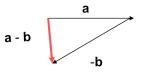
 $c_x \underline{\mathbf{i}} + c_y \underline{\mathbf{j}} + c_z \underline{\mathbf{k}} =$
 $(a_x + b_x) \underline{\mathbf{i}} + (a_y + b_y) \underline{\mathbf{j}} + (a_z + b_z) \underline{\mathbf{k}}$



Vector algebra laws

- $\underline{\mathbf{a}} + \underline{\mathbf{b}} = \underline{\mathbf{b}} + \underline{\mathbf{a}}$: commutative law
- $\underline{\mathbf{a}} + (\underline{\mathbf{b}} + \underline{\mathbf{c}}) = (\underline{\mathbf{a}} + \underline{\mathbf{b}}) + \underline{\mathbf{c}}$: associative law
- Can treat vector equations in same way as ordinary algebra <u>a + b = c</u> ⇒ <u>a = c - b</u>
- Note that vector -<u>b</u> is equal in magnitude to <u>b</u> but in the opposite direction.

SO $\underline{\mathbf{a}} - \underline{\mathbf{b}} = \underline{\mathbf{a}} + (-\underline{\mathbf{b}})$



$$a - b = a + (-b)$$

Multiplication of a vector by a scalar

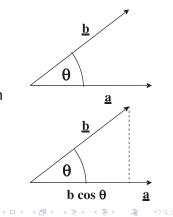
- This gives a vector in the same direction as the original but of proportional magnitude.
- For any scalars α and β and vectors $\underline{\mathbf{a}}$ and $\underline{\mathbf{b}}$
 - (αβ) <u>a</u> = α(β <u>a</u>) = β(α <u>a</u>) = <u>a</u> (αβ) : associative & commutative
 - $(\alpha + \beta)\mathbf{\underline{a}} = \alpha\mathbf{\underline{a}} + \beta\mathbf{\underline{a}}$: distributive
 - $\bullet \ \alpha(\underline{\mathbf{a}} + \underline{\mathbf{b}}) = \alpha \underline{\mathbf{a}} + \alpha \underline{\mathbf{b}}$

2.1.1 The scalar (dot) product

Scalar (or dot) product definition: $\underline{\mathbf{a}}.\underline{\mathbf{b}} = |\underline{\mathbf{a}}|.|\underline{\mathbf{b}}|\cos\theta \equiv ab\cos\theta$

(write shorthand $|\underline{\mathbf{a}}| = \mathbf{a}$).

- Scalar product is the magnitude of <u>a</u> multiplied by the projection of <u>b</u> onto <u>a</u>.
- Obviously if <u>a</u> is perpendicular to <u>b</u> then <u>a</u>.<u>b</u> = 0
- Also <u>a.a</u> = |a|² (since θ =0°) Hence a = √(<u>a.a</u>)



Properties of scalar product

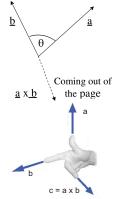
- (i) $\underline{i}.\underline{i} = \underline{j}.\underline{j} = \underline{k}.\underline{k} = 1$ and $\underline{i}.\underline{j} = \underline{j}.\underline{k} = \underline{k}.\underline{i} = 0$
- (ii) This leads to $\underline{\mathbf{a}} \cdot \underline{\mathbf{b}} = (a_x \underline{\mathbf{i}} + a_y \underline{\mathbf{j}} + a_z \underline{\mathbf{k}}) \cdot (b_x \underline{\mathbf{i}} + b_y \underline{\mathbf{j}} + b_z \underline{\mathbf{k}})$ = $a_x b_x + a_y b_y + a_z b_z$
- $\begin{array}{ll} \mbox{iii}) & \underline{\mathbf{a}}.\underline{\mathbf{b}} = \underline{\mathbf{b}}.\underline{\mathbf{a}} : \mbox{commutative} \\ & \underline{\mathbf{a}}.(\underline{\mathbf{b}}+\underline{\mathbf{c}}) = \underline{\mathbf{a}}.\underline{\mathbf{b}} + \underline{\mathbf{a}}.\underline{\mathbf{c}} : \mbox{distributive} \end{array}$
- (iv) Parentheses are important Note $(\underline{\mathbf{u}}.\underline{\mathbf{v}}) \underline{\mathbf{w}} \neq \underline{\mathbf{u}} (\underline{\mathbf{v}}.\underline{\mathbf{w}})$ because one is a vector along $\underline{\hat{\mathbf{w}}}$, the other is along $\underline{\hat{\mathbf{u}}}$.

2.1.2 The vector (cross) product

Vector (or cross) product of two vectors, definition: $\underline{\mathbf{a}} \times \underline{\mathbf{b}} = |\underline{\mathbf{a}}||\underline{\mathbf{b}}| \sin\theta \ \underline{\hat{\mathbf{n}}}$

where $\underline{\hat{n}}$ is a *unit vector* in a direction *perpendicular* to both \underline{a} and \underline{b} . To get direction of $\underline{a} \times \underline{b}$ use right hand rule:

- ► i) Make a set of directions with your *right* hand→ thumb & first index finger, and with middle finger positioned perpendicular to plane of both
- \blacktriangleright ii) Point your thumb along the first vector $\underline{\mathbf{a}}$
- ► iii) Point your 1st index finger along <u>b</u>, making the smallest possible angle to <u>a</u>
- ► iv) The direction of the middle finger gives the direction of $\underline{\mathbf{a}} \times \underline{\mathbf{b}}$.



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Properties of vector product

$$\bullet \ \underline{\mathbf{i}} \times \underline{\mathbf{j}} = \underline{\mathbf{k}} \quad ; \quad \underline{\mathbf{j}} \times \underline{\mathbf{k}} = \underline{\mathbf{i}} \quad ; \quad \underline{\mathbf{k}} \times \underline{\mathbf{i}} = \underline{\mathbf{j}} \quad ; \quad \underline{\mathbf{i}} \times \underline{\mathbf{i}} = \mathbf{0} \text{ etc.}$$

- $(\underline{\mathbf{a}} + \underline{\mathbf{b}}) \times \underline{\mathbf{c}} = (\underline{\mathbf{a}} \times \underline{\mathbf{c}}) + (\underline{\mathbf{b}} \times \underline{\mathbf{c}})$: distributive
- $\underline{\mathbf{a}} \times \underline{\mathbf{b}} = -\underline{\mathbf{b}} \times \underline{\mathbf{a}}$: NON-commutative
- $(\underline{\mathbf{a}} \times \underline{\mathbf{b}}) \times \underline{\mathbf{c}} \neq \underline{\mathbf{a}} \times (\underline{\mathbf{b}} \times \underline{\mathbf{c}})$: NON-associative
- ► If *m* is a scalar, $m(\underline{\mathbf{a}} \times \underline{\mathbf{b}}) = (m\underline{\mathbf{a}}) \times \underline{\mathbf{b}} = \underline{\mathbf{a}} \times (m\underline{\mathbf{b}}) = (\underline{\mathbf{a}} \times \underline{\mathbf{b}})m$

•
$$\underline{\mathbf{a}} \times \underline{\mathbf{b}} = 0$$
 if vectors are parallel (0^o)
i.e $\underline{\mathbf{a}} \times \underline{\mathbf{a}} = 0$

Vector product in components

Cross product written out in components:

•
$$\underline{\mathbf{a}} \times \underline{\mathbf{b}} = (a_x, a_y, a_z) \times (b_x, b_y, b_z)$$

= $(a_x \underline{\mathbf{i}} + a_y \underline{\mathbf{j}} + a_z \underline{\mathbf{k}}) \times (b_x \underline{\mathbf{i}} + b_y \underline{\mathbf{j}} + b_z \underline{\mathbf{k}})$
• Since $\underline{\mathbf{i}} \times \underline{\mathbf{i}} = \underline{\mathbf{j}} \times \underline{\mathbf{j}} = \underline{\mathbf{k}} \times \underline{\mathbf{k}} = 0$ and $\underline{\mathbf{i}} \times \underline{\mathbf{j}} = \underline{\mathbf{k}}$ etc.
• $\underline{\mathbf{a}} \times \underline{\mathbf{b}} = (a_y b_z - a_z b_y) \underline{\mathbf{i}} - (a_x b_z - a_z b_x) \underline{\mathbf{j}} + (a_x b_y - a_y b_x) \underline{\mathbf{k}}$

This is much easier when we write in *determinant* form:

$$\underline{\mathbf{a}} \times \underline{\mathbf{b}} = \begin{vmatrix} \underline{\mathbf{i}} & \underline{\mathbf{j}} & \underline{\mathbf{k}} \\ a_x & a_y & a_z \\ b_x & b_y & b_z \end{vmatrix}$$
(1)

2.1.3 Examples of scalar & vector products in mechanics

a) Scalar product

Work done on a body by a force through distance $\underline{\mathbf{dr}}$ from position 1 to 2

 $W_{12} = \int_1^2 \underline{\mathbf{F}} \cdot \underline{\mathbf{dr}}$

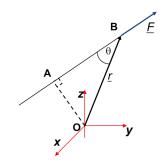
Only the component of force parallel to the line of displacement does work.

b) Vector product

A *torque* about O due to a force $\underline{\mathbf{F}}$ acting at B :

 $\underline{\tau} = \underline{\mathbf{r}} \times \underline{\mathbf{F}}$

Torque is a vector with direction perpendicular to both $\underline{\mathbf{r}}$ and $\underline{\mathbf{F}}$, magnitude of $|\underline{\mathbf{r}}||\underline{\mathbf{F}}|\sin\theta$.



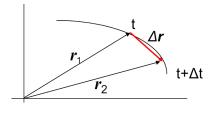
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2.2 Differentiation of vectors

Notation: a dot above the function indicates derivative wrt time. A "dash" indicates derivative wrt a spatial coordinate.

$$\dot{y} \equiv \frac{dy}{dt} \qquad y' \equiv \frac{dy}{dx}$$
$$\dot{\underline{a}} = \lim_{\Delta t \to 0} \frac{\underline{a}(t + \Delta t) - \underline{a}(t)}{\Delta t} = \lim_{\Delta t \to 0} \frac{\Delta \underline{a}}{\Delta t}$$
$$\underline{a}(t) = a_x(t)\underline{i} + a_y(t)\underline{j} + a_z(t)\underline{k}$$
$$\dot{\underline{a}} = \lim_{\Delta t \to 0} \left(\frac{a_x(t + \Delta t) - a_x(t)}{\Delta t} \underline{i} + \dots \right) \qquad \mathbf{a}(t + \Delta t) \qquad \mathbf{a}(t)$$
Hence $\underline{\dot{a}} = \dot{a}_x \underline{i} + \dot{a}_y \underline{j} + \dot{a}_z \underline{k}$

2.2.1 Vector velocity



$$\Delta \underline{\mathbf{r}} = \underline{\mathbf{r}}_2 - \underline{\mathbf{r}}_1$$

$$\underline{\mathbf{v}} = \lim_{\Delta t \to 0} \frac{\Delta \underline{\mathbf{r}}}{\Delta t}$$

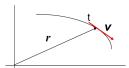
 Velocity at any point is tangent to the path at that point

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•
$$\underline{\mathbf{v}} = \frac{d\underline{\mathbf{r}}}{dt} = \underline{\dot{\mathbf{r}}}$$

In one dimension: Abandon vector notation and simply write $v = \frac{dx}{dt} = \dot{x}$, (+v in +x direction, -v in -x direction).



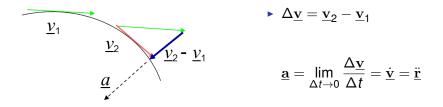
Example - 1D motion

A body has velocity $v_0 = -15 ms^{-1}$ at position $x_0 = 20m$ and has a time-dependent acceleration $a(t) = 6t - 4 [ms^{-2}]$. Find the value of *x* for which the body instantaneously comes to rest.



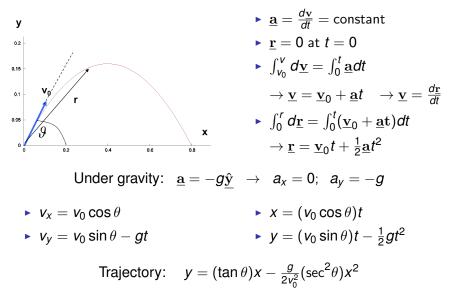
•
$$a(t) = 6t - 4 \ [ms^{-2}]$$
; $x_0 = 20 \ m$; $v_0 = -15 \ ms^{-1}$
• $\dot{v} = 6t - 4 \rightarrow v = \int a(t)dt = 3t^2 - 4t + c$
• At $t = 0, v = -15 \ ms^{-1} \rightarrow c = -15 \ ms^{-1}$
• $v = 3t^2 - 4t - 15$
• $v = 0$ for $3t^2 - 4t - 15 = 3(t - 3)(t + \frac{5}{3}) = 0$
 $\rightarrow t = 3s \ (also - \frac{5}{3}s)$
• $x = \int v(t)dt = t^3 - 2t^2 - 15t + c' \rightarrow x = 20 \ m \text{ at } t = 0, c' = 20 \ m$
• $x(t) = 27 - 18 - 45 + 20 = -16 \ m$

2.2.2 Vector acceleration



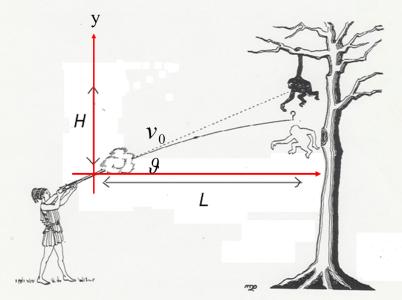
In one dimension: Abandon vector notation and simply write $a = \frac{dv}{dt} = \dot{v} = \ddot{x}$, (+a in +x direction, -a in -x direction).

Example: constant acceleration - projectile motion in 2D



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The monkey and the hunter



3.1 Dimensional analysis

- A useful method for determining the units of a variable in an equation
- Useful for checking the correctness of an equation which you have derived after some algebraic manipulation. Dimensions need to be correct !
- Determining the form of an equation itself

Most physical quantities can be expressed in terms of combinations of basic dimensions. These are certainly not unique :

- mass (M)
- length (L)
- time (T)
- electric charge (Q)
- temperature (θ)

Note: The term "dimension" is not quite the same as "unit", but obviously closely related.

Quantity	Unit		Dimension
Frequency	Hertz (Hz)	= (cycles) s^{-1}	$T^{-1} \\ MLT^{-2} \\ ML^2T^{-2} \\ ML^2T^{-3} \\ QT^{-1} \\ ML^2T^{-2}Q^{-1}$
Force	Newton (N)	= kg m s^{-2}	
Energy	Joule (J)	= N m = kg m ² s^{-2}	
Power	Watt (W)	= J s^{-1} = kg m ² s^{-3}	
Current	Ampere (A)	= Cs ⁻¹	
EMF	Volt (V)	= Nm C ⁻¹ = kg m ² $s^{-2}C^{-1}$	

Dimensional analysis is best illustrated with examples.

3.1.1 The period of a pendulum

How does the period of a pendulum depend on its length?

- Variables: period P, mass m, length I, acceleration due to gravity g
- Guess the form: let $P = k m^a \ell^b g^c$

(k is a dimensionless constant)

•
$$T^1 = M^a L^b (LT^{-2})^c = M^a L^{b+c} T^{-2c}$$

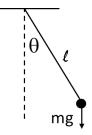
Compare terms:

$$a = 0, b + c = 0, -2c = 1$$

$$\rightarrow c = -1/2, b = 1/2$$

$${\pmb P}={\pmb k}\sqrt{rac{\ell}{g}}$$

We know that $P = 2\pi \sqrt{\frac{\ell}{g}}$: we obtained this form using dimensions and without using equation of motion: IMPRESSIVE !



3.1.2 Kepler's third law

How does the period of an orbiting mass depend on its radius?

- Variables: period P, central mass M₀, orbit radius r, Gravitational constant G
- Guess the form: let P = k M₀^ar^bG^c
 (k is a dimensionless constant)
- Dimensions of $G \rightarrow (MLT^{-2}).L^2M^{-2}$

►
$$T^1 = M^a L^b (M^{-1} L^3 T^{-2})^c$$

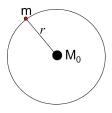
= $M^{(a-c)} L^{b+3c} T^{-2c}$

Compare terms:

$$a-c = 0, b+3c = 0, -2c = 1$$

 $\rightarrow a = -1/2, c = -1/2, b = 3/2$

 $P = k M_0^{-1/2} r^{3/2} G^{-1/2} \rightarrow P^2 = \frac{k^2}{GM_0} r^3$



$$\blacktriangleright \ \frac{GmM_0}{r^2} = \frac{mv^2}{r}$$

•
$$V = \frac{2\pi r}{P}$$

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►
$$P^2 = \frac{4\pi^2}{GM_0}r^3$$

3.1.3 The range of a cannon ball

A cannon ball is fired with V_y upwards and V_x horizontally, assume no air resistance.

- Variables: V_x, V_y, distance travelled along x (range) R, acceleration due to gravity g
- First with no use of directed length dimensions
- Let $R = kV_x^a V_y^b g^c$.

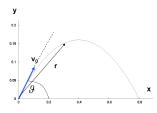
(k is a dimensionless constant)

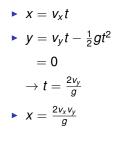
- Dimensionally $L = (L/T)^{a+b} (L/T^2)^c$
- Compare terms:

a+b+c=1 and a+b+2c=0, which leaves one exponent undetermined.

- ▶ Now use directed length dimensions, then V_x will be dimensioned as L_x/T , V_y as L_y/T , R as L_x and g as L_y/T^2
- ► The dimensional equation becomes: $L_x = (L_x/T)^a (L_y/T)^b (L_y/T^2)^c$ $\rightarrow a = 1, b = 1 \text{ and } c = -1.$

$$R = k rac{v_x v_y}{g}$$





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3.1.4 Example of limitations of the method

- Let $y = f(x_1, x_2, ..., x_n)$ where $x_1, x_2, ..., x_n$ have independent dimensions
- However in general $y = (x_1^a x_2^b \dots x^n) \phi(u_1, \dots u_k)$ where u_i are dimensionless variables

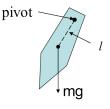
Extend to how the period of a *rigid* pendulum depends on length pivot to CM.

In actual fact P ≡ P(g, ℓ, m, I) where I is the moment of inertia

•
$$[I] = ML^2 \rightarrow \text{ can define } u = \frac{I}{m\ell^2}$$

$${\cal T}=\sqrt{rac{\ell}{g}}\,\phi({\it u})$$

i.e. Equation is not reproduced



$$T=2\pi\sqrt{rac{1}{mg\ell}}$$

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3.2 Newton's Laws of motion

- NI: Every body continues in a state of rest or in uniform motion (constant velocity in straight line) unless acted upon by an external force.
- ► NII: The rate of change of momentum is equal to the applied force; where the momentum is defined as the product of mass and velocity (<u>p</u> = m<u>v</u>). [i.e. the applied force <u>F</u> on a body is equal to its mass *m* multiplied by its acceleration <u>a</u>.]
- NIII: When one body exerts a force on a second body, the second body simultaneously exerts a force equal in magnitude and opposite in direction on the first body [i.e. action and reaction forces are equal in magnitude and opposite in direction.]

3.3 Frames of reference

- A frame of reference is an environment which is used to observe an event or the motion of a particle.
- A coordinate system is associated with the frame to observe the event (eg the body's location over time).
- The observer is equipped with measuring tools (eg rulers and clocks) to measure the positions and times of events.
- In classical mechanics, time intervals between events is the same in all reference frames (time is absolute).
- In relativity, we will need to use space-time frames.
- A reference frame in which NI is satisfied is called an inertial reference frame.

Inertial reference frames

A frame in which Newton's first law is satisfied:

- Deep space
- The Earth? [Only in circumstances where we can ignore gravity & the spin of the Earth.]

Principle of Relativity : The laws of Physics are the same in all inertial frames of reference.

$$\bigcup_{0}^{S} x \qquad \bigcup_{0' \qquad x'}^{V_0 \qquad S'}$$

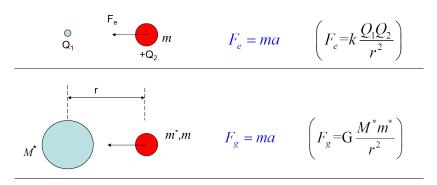
At t = 0, x = 0, x' = 0 and S and S' are coincident.

Galilean Transformation of coordinates:

• $x' = x - v_0 t$, y' = y, z' = z, t' = t

- Velocity of a body v in S; velocity measured in $S' = v v_0$
- Acceleration measured in S' = a
- ► Hence F' = F (consistent with the principle of relativity)

3.4 The Principle of Equivalence



- The Principle of Equivalence dictates that $m = m^*$.
- Inertial mass = Gravitational mass
- This may seem obvious, but it was not an original postulate of Newton

4.1 Newton's Second Law

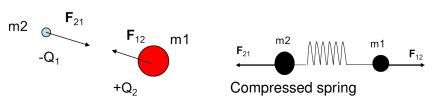
The rate of change of momentum of a body is equal to the applied force on the body.

- $\underline{\mathbf{F}} = \frac{d\mathbf{p}}{dt} = m\underline{\mathbf{a}}$ where $\underline{\mathbf{p}} = m\underline{\mathbf{v}}$
- In components: $(F_x, F_y, F_z) = m(a_x, a_y, a_z)$
- ► Assuming constant mass, we can define the equation of motion in 1D : $F = \frac{d(mv)}{dt} = m \frac{d^2x}{dt^2}$
- We require two initial conditions for a unique solution: e.g. v = v₀ at t = 0 and x = x₀ at t = 0

We shall later solve the EOM for three examples: (i) F = constant, (ii) $F \propto -v$, (iii) $F \propto -x$

4.2 Newton's Third Law

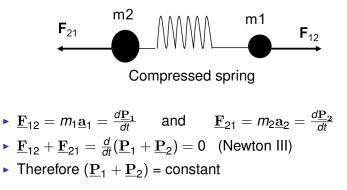
Action and reaction forces are equal in magnitude and opposite in direction.



Electrostatic interaction

$$\underline{\mathbf{F}}_{12} = -\underline{\mathbf{F}}_{21}$$

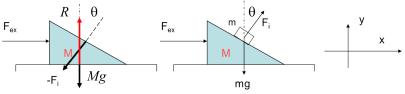
Conservation of momentum



In an isolated system, the total momentum is conserved.

Newton II : Example 1. E.O.M. under constant force

How fast should we accelerate the triangular wedge to keep the block *m* stationary on the wedge?



Forces on wedge:

- Horizontal: $F_{ex} F_i \sin \theta = MA_x$
- Vertical: $R F_i \cos \theta Mg = 0$

Forces on block:

- Horizontal: $F_i \sin \theta = ma_x$
- Vertical: $F_i \cos \theta mg = ma_y$

For block to remain at the same place $A_x = a_x$ and $a_y = 0$

•
$$F_i = \frac{mg}{\cos\theta}$$
 and $a_x = g \tan\theta = A_x$

• Hence $F_{ex} = Mg \tan \theta + mg \tan \theta = (m + M)g \tan \theta$

Example constant force, continued

What is the internal force that the blocks apply on each other and the reaction force by the ground on M?

From before:

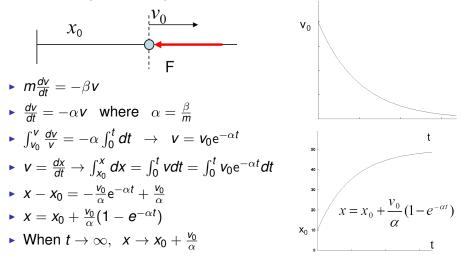
•
$$F_i = \frac{mg}{\cos\theta}$$

$$\bullet R - F_i \cos \theta - Mg = 0$$

• Hence: $R = F_i \cos \theta + Mg = (m + M)g$

Example 2. Force proportional to velocity

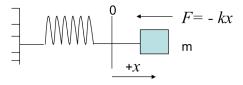
Solve the equation of motion for the case $F = -\beta v$ ($\beta > 0$) with $x = x_0$ and $v = v_0$ at t = 0



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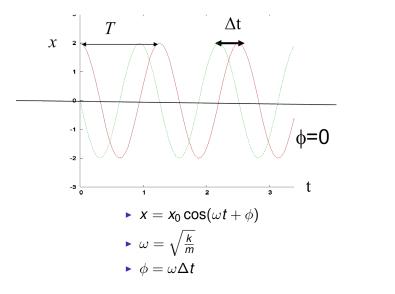
Example 3. Force proportional to position: simple harmonic oscillator

Solving the equation of motion for the case $F = m \frac{d^2 x}{dt^2} = -kx$



• $m\frac{d^2x}{dt^2} + \frac{k}{m}x = 0$; trial solution $x = A\cos\omega t + B\sin\omega t$ $\rightarrow \dot{x} = -A\omega\sin\omega t + B\omega\cos\omega t$; $\ddot{x} = -A\omega^2\cos\omega t - B\omega^2\sin\omega t$ • $\ddot{x} = -\omega^2 x \rightarrow \omega^2 = \frac{k}{m}$ • Alternatively $x = x_0\cos(\omega t + \phi)$ (or $x = x_0Re[e^{i(\alpha t + \phi)}]$) • Expand : $x = x_0(\cos(\omega t)\cos\phi - \sin(\omega t)\sin\phi)$ $A = x_0\cos\phi$; $B = -x_0\sin\phi \rightarrow x_0^2 = A^2 + B^2$; $\tan\phi = -B/A$ • $x_0 = \text{amplitude}, \phi = \text{phase}, \omega = \text{angular frequency}$ ($T = \frac{2\pi}{\omega}$)

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4.3 Energy conservation in one dimension Work done on a body by a force F

•
$$W = \int_{x_1}^{x_2} F(x) dx = m \int_{x_1}^{x_2} \frac{dv}{dt} dx$$

- ► We can write: $\frac{dv}{dt}dx = \frac{dx}{dt}dv = vdv$ hence $\int_{x_1}^{x_2} F(x)dx = m \int_{v_1}^{v_2} vdv = \frac{1}{2}m(v_2^2 - v_1^2) = T_2 - T_1$
- Now introduce an arbitrary reference point x₀

 $\int_{x_1}^{x_2} F dx = \int_{x_0}^{x_2} F dx - \int_{x_0}^{x_1} F dx \text{ defines a conservative force}$ hence $T_2 + [-\int_{x_0}^{x_2} F dx] = T_1 + [-\int_{x_0}^{x_1} F dx]$

▶ We define the *potential energy* U(x) at a point x :

 $U(x) - U(x_0) = -\int_{x_0}^x F dx$ and hence $T_2 + U_2 = T_1 + U_1$ (total energy PE + KE conserved)

► Note the minus sign. The potential energy (relative to a reference point) is always the *negative* of the work done by the force $\rightarrow F(x) = -\frac{dU}{dx}$

5.1 Conservative forces

$$W_{ab} = \int_a^b \underline{\mathbf{F}} \cdot d\underline{\mathbf{r}} = U(a) - U(b)$$

For a conservative field of force, the work done depends only on the initial and final positions of the particle independent of the path.

The conditions for a conservative force (all equivalent) are:

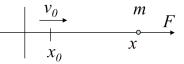
- ► The force is derived from a (scalar) potential function: $\underline{\mathbf{F}}(\underline{\mathbf{r}}) = -\nabla U \rightarrow F(x) = -\frac{dU}{dx}$ etc.
- ► There is zero net work by the force when moving a particle around any closed path: $W = \oint_c \mathbf{F} \cdot d\mathbf{r} = 0$
- In equivalent vector notation $\underline{\nabla} \times \underline{\mathbf{F}} = \mathbf{0}$

For any force: $W_{ab} = \frac{1}{2}mv_b^2 - \frac{1}{2}mv_a^2$

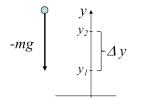
Only for a *conservative* force: $W_{ab} = U(a) - U(b)$

Conservative force: example 1. Constant acceleration

Consider a particle moving under constant force (in 1-D).



- F = ma. Say at $t = 0 \rightarrow x = x_0$ and $v = v_0$
- $T_2 + U_2 = T_1 + U_1$ (the total energy is conserved)
- $\frac{1}{2}mv^2 (ma)x = \frac{1}{2}mv_0^2 (ma)x_0 = \text{constant}$ • $v^2 = v_0^2 + 2a(x - x_0)$



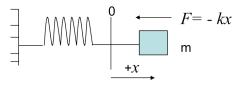
Gravitational potential energy

•
$$U(\Delta y) = -\int_{y_1}^{y_2} F(y) dy$$

•
$$U(\Delta y) = -\int_{y_1}^{y_2} (-mg) dy = mg(y_2 - y_1)$$

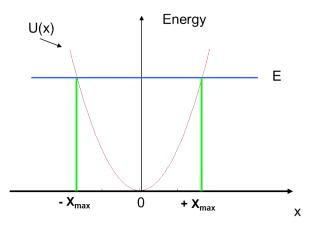
Example 2. Simple harmonic oscillator

Equation of motion: $F = m \frac{d^2 x}{dt^2} = -kx$



- Potential energy: $U(x) = -\int_0^x F dx = -\int_0^x (-kx) dx = \frac{kx^2}{2}$
- Total energy: $E = T(x) + U(x) = \frac{1}{2}m\dot{x}^2 + \frac{kx^2}{2}$
- ► Check conservation of energy: EOM : $m\ddot{x} + kx = 0$ \rightarrow [multiply by \dot{x}] $m\ddot{x}\dot{x} + kx\dot{x} = 0$ $\rightarrow \frac{1}{2}m\frac{d}{dt}(\dot{x}^2) + \frac{1}{2}k\frac{d}{dt}(x^2) = 0$
- ► Integrate wrt $t: \frac{1}{2}m\dot{x}^2 + \frac{1}{2}kx^2 = \text{constant} \rightarrow \text{i.e. energy conserved.}$

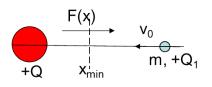
SHM potential energy curve



$$\bullet \ E = U(x) + \frac{1}{2}mv^2$$

• The particle can only reach locations x that satisfy U < E

Example 3. Minimum approach of a charge

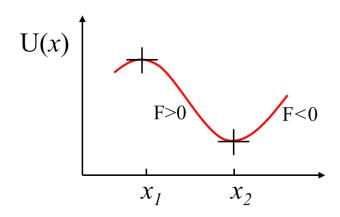


A particle of mass *m* and charge $+Q_1$ starts from $x = +\infty$ with velocity v_0 . It approaches a fixed charge +Q. Calculate its minimum distance of approach x_{min} .

Force on charge
$$+Q_1$$
: $F(x) = +\frac{QQ_1}{4\pi\epsilon_0 x^2}$ (+ve direction)

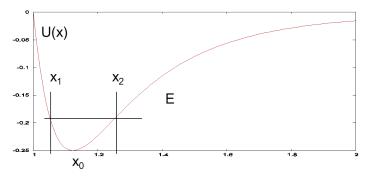
- ► Potential energy at point *x*: $U(x) = -\int_{\infty}^{x} F(x) dx = +\frac{QQ_1}{4\pi\epsilon_0 x}$ (where PE = 0 at $x = \infty$)
- Conservation of energy : $\frac{1}{2}mv_0^2 + 0 = \frac{1}{2}mv^2 + U(x)$
- Min. dist. when v = 0: $\frac{1}{2}mv_0^2 = \frac{QQ_1}{4\pi\epsilon_0 x_{min}} \rightarrow x_{min} = \frac{QQ_1}{2\pi m\epsilon_0 v_0^2}$

5.2 Potential with turning points



- U is a maximum: unstable equilibrium
- ► U is a minimum: *stable* equilibrium

5.2.1 Oscillation about stable equilibrium

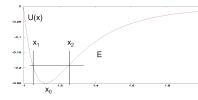


- For SHM : $U(x) = \frac{1}{2}k(x x_0)^2$
- Taylor expansion about x₀:

$$U(x) = U(x_0) + \underbrace{\left[\frac{dU}{dx}\right]_{x=x_0}}_{=0} (x-x_0) + \frac{1}{2!} \underbrace{\left[\frac{d^2U}{dx^2}\right]_{x=x_0}}_{=k} (x-x_0)^2 + \dots$$

Example: The Lennard-Jones potential

The Lennard-Jones potential describes the potential energy between two atoms in a molecule: $U(x) = \epsilon[(x_0/x)^{12} - 2(x_0/x)^6]$ (ϵ and x_0 are constants and x is the distance between the atoms).

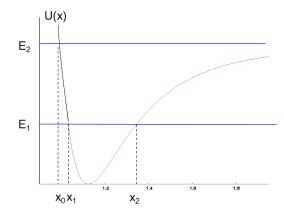


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Show that the motion for small displacements about the minimum is simple harmonic and find its frequency.

• $U(x) = U(x_0) + \left[\frac{dU}{dx}\right]_{x=x_0}(x-x_0) + \frac{1}{2!}\left[\frac{d^2U}{dx^2}\right]_{x=x_0}(x-x_0)^2 + \dots$ • $U(x_0) = \epsilon[(x_0/x)^{12} - 2(x_0/x)^6]_{x=x_0} = -\epsilon$ • $\frac{dU(x)}{dx}|_{x=x_0} = 12\epsilon[-\frac{1}{x_0}(x_0/x)^{13} + \frac{1}{x_0}(x_0/x)^7]_{x=x_0} = 0$ as expected. • $\frac{d^2U(x)}{dx^2}|_{x=x_0} = \frac{12\epsilon}{x_0^2}[13(x_0/x)^{14} - 7(x_0/x)^8]_{x=x_0} = \frac{72\epsilon}{x_0^2}$ • Hence $U(x) \approx -\epsilon + \frac{72\epsilon}{2!x_0^2}(x-x_0)^2$ • $F(x) = -\frac{dU}{dx} \approx -\frac{1}{2} \times 2(\frac{72\epsilon}{x_0^2})(x-x_0) = -k(x-x_0)$ SHM about x_0 • Angular frequency of small oscillations : $\omega^2 = \frac{k}{m} = \frac{72\epsilon}{mx^2}$

5.2.2 Bounded and unbounded potentials



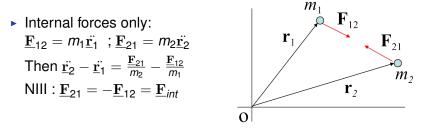
- ▶ Bounded motion : $E = E_1$: x constrained $x_1 < x < x_2$
- ▶ Unbounded motion : $E = E_2$: x unconstrained at high x $x_0 < x < \infty$

6.1 Lab & CM frames of reference

From hereon we will deal with 2 inertial frames:

- The Laboratory frame: this is the frame where measurements are actually made
- The centre of mass frame: this is the frame where the centre of mass of the system is at rest and where the total momentum of the system is zero

6.2 Internal forces and reduced mass



• Define $\underline{\mathbf{r}} = \underline{\mathbf{r}}_2 - \underline{\mathbf{r}}_1 \rightarrow \dot{\underline{\mathbf{r}}} = \dot{\underline{\mathbf{r}}}_2 - \dot{\underline{\mathbf{r}}}_1$

•
$$\underline{\mathbf{F}}_{int}\left(\frac{1}{m_1}+\frac{1}{m_2}\right)=\underline{\ddot{\mathbf{r}}}$$

• Define
$$\frac{1}{\mu} = \frac{1}{m_1} + \frac{1}{m_2} \rightarrow \underline{\mathbf{F}}_{int} = \mu \underline{\ddot{\mathbf{F}}}$$

 $\mu = \frac{m_1 m_2}{m_1 + m_2}$ is the *reduced mass* of the system

This defines the equation of motion of the relative motion of the particles under internal forces, with position vector \underline{r} & mass μ

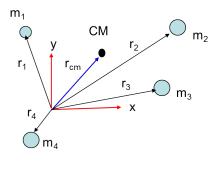
6.3 The Centre of Mass

The centre of mass (CM) is the point where the mass-weighted position vectors (moments) relative to the point sum to zero ; the CM is the mean location of a distribution of mass in space.

Take a system of *n* particles, each with mass *m_i* located at positions <u>r</u>_i, the position vector of the CM is defined by:

$$\sum_{i=1}^{n} m_i(\underline{\mathbf{r}}_i - \underline{\mathbf{r}}_{cm}) = \mathbf{0}$$

• Solve for $\underline{\mathbf{r}}_{cm}$: $\underline{\mathbf{r}}_{cm} = \frac{1}{M} \sum_{i=1}^{n} m_i \underline{\mathbf{r}}_i$ where $M = \sum_{i=1}^{n} m_i$



Example : SHM of two connected masses in 1D

SHM between two masses m_1 and m_2 connected by a spring

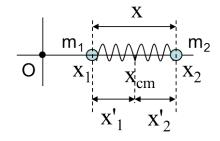
• $x = x_2 - x_1$; Natural length L

►
$$F_{int} = -k (x - L) = \mu \ddot{x}$$

 $(\mu = \frac{m_1 m_2}{m_1 + m_2} = \text{reduced mass})$
► $\ddot{x} + \frac{k}{\mu} (x - L) = 0$
Solution: $x = x_0 \cos(\omega t + \phi) + L$
where $\omega = \sqrt{\frac{k}{\mu}}$

With respect to the CM:

•
$$x_{CM} = \frac{m_1 x_1 + m_2 x_2}{M}$$
 where $M = m_1 + m_2$
• $x'_1 = x_1 - x_{cm} = \frac{M x_1 - m_1 x_1 - m_2 x_2}{M} = -\frac{m_2 x}{M}$
• $x'_2 = x_2 - x_{cm} = \frac{M x_2 - m_1 x_1 - m_2 x_2}{M} = \frac{m_1 x}{M}$
Eq. take $m_1 = m_2 = m \rightarrow \omega = \sqrt{\frac{2k}{m}}$; $x'_1 = -\frac{1}{2}x$, $x'_2 = \frac{1}{2}x$



6.3.1 CM of a continuous volume

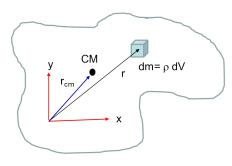
If the mass distribution is continuous with density $\rho(\underline{\mathbf{r}})$ inside a volume *V*, then:

$$\sum_{i=1}^{n} m_i(\underline{\mathbf{r}}_i - \underline{\mathbf{r}}_{cm}) = 0$$

becomes
$$\int_V \rho(\underline{\mathbf{r}})(\underline{\mathbf{r}} - \underline{\mathbf{r}}_{cm}) dV = 0$$

where $dm = \rho(\underline{\mathbf{r}}) dV$

► Solve for r_{cm} $\underline{\mathbf{r}}_{cm} = \frac{1}{M} \int_{V} \rho(\underline{\mathbf{r}}) \underline{\mathbf{r}} \, dV$ where *M* is the total mass in the volume



Example: the CM of Mount Ranier

Mount Ranier has approximately the shape of a cone (assume uniform density) and its height is 4400 m. At what height is the centre of mass?

We have cylindrical symmetry - just need to consider the *y* direction. Integrate from top (y = 0) to bottom (y = h)

•
$$y_{cm} = \frac{\int_0^h y dm}{M} = \frac{\int_0^h y dm}{\int_0^h dm}$$

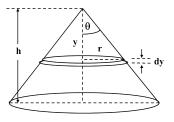
 $dm = \rho(\pi r^2) dy = \rho(\pi y^2 \tan^2 \theta) dy$

$$\flat \ y_{cm} = \frac{\int_0^h y \rho(\pi y^2 \tan^2 \theta) dy}{\int_0^h \rho(\pi y^2 \tan^2 \theta) dy}$$

$$y_{cm} = \frac{\int_0^h y^3 dy}{\int_0^h y^2 dy} = \frac{3h^4}{4h^3} = \frac{3h}{4}$$
 (measured from the top)

• $h = 4400 m \rightarrow y_{cm} = 1100 m$ above the base





6.3.2 Velocity in the Centre of Mass frame

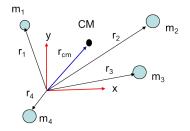
- ► The position of the centre of mass is given by: $\underline{\mathbf{r}}_{cm} = \frac{1}{M} \sum_{i=1}^{n} m_i \underline{\mathbf{r}}_i$ where $M = \sum_{i=1}^{n} m_i$
- The velocity of the CM:

$$\mathbf{\underline{v}_{cm}} = \mathbf{\underline{\dot{r}}_{cm}} = \frac{1}{M} \sum_{i=1}^{n} m_i \mathbf{\underline{\dot{r}}_i}$$

• In the Lab frame, total momentum $\underline{\mathbf{p}}_{tot}$:

$$\underline{\mathbf{p}}_{tot} = \sum_{i=1}^{n} m_i \, \underline{\mathbf{\dot{r}}}_i = M \underline{\mathbf{v}}_{cm}$$

Hence the total momentum of a system in the Lab frame is equivalent to that of a single particle having a mass Mand moving at a velocity $\underline{\mathbf{v}}_{cm}$



6.3.3 Momentum in the CM frame

- ► Velocity of the CM: $\underline{\mathbf{v}}_{\mathbf{cm}} = \underline{\mathbf{\dot{r}}}_{\mathbf{cm}}^n = \frac{\sum_{i=1}^n m_i \, \underline{\mathbf{\dot{r}}}_i}{\sum_i m_i} = \frac{\sum_{i=1}^n m_i \, \underline{\mathbf{v}}_i}{\sum_i m_i}$
- \blacktriangleright Velocity of a body in the CM relative to Lab $\ \underline{\mathbf{v}}_i' = \underline{\mathbf{v}}_i \underline{\mathbf{v}}_{cm}$
- The total momentum of the system of particles in the CM:

$$\sum_{i} \underline{\mathbf{p}}_{i}' = \sum_{i} m_{i} \underline{\mathbf{v}}_{i}' = \sum_{i} m_{i} (\underline{\mathbf{v}}_{i} - \underline{\mathbf{v}}_{cm})$$

$$= \sum_{i} m_{i} \underline{\mathbf{v}}_{i} - \sum_{i} m_{i} \frac{\sum_{j} m_{j} \underline{\mathbf{v}}_{j}}{\sum_{j} m_{j}} = \sum_{i} m_{i} \underline{\mathbf{v}}_{i} - \sum_{j} m_{j} \underline{\mathbf{v}}_{j} = \mathbf{0}$$

Hence the total momentum of a system of particles in the CM frame is equal to zero

In addition, the total energy of the system is a minimum compared to all other inertial reference frames.

6.3.4 Motion of CM under external forces Force on particle *i*: $m_i \ddot{\mathbf{r}}_i = \mathbf{F}_i^{ext} + \mathbf{F}_i^{int}$ $\sum_{i=1}^{n} m_{i} \underline{\ddot{\mathbf{r}}}_{i} = \sum_{i=1}^{n} \underline{\mathbf{F}}_{i}^{ext} + \sum_{i=1}^{n} \underline{\mathbf{F}}_{i}^{int} = \sum_{i=1}^{n} \underline{\mathbf{F}}_{i}^{ext}$ all masses external forces internal forces = zero $\mathbf{F}_{1}^{\text{int}}$ m_2 • $\underline{\mathbf{r}}_{CM} = \sum_{i} \frac{m_{i} \underline{\mathbf{r}}_{i}}{M}$ F^{int} r_i where $M = \sum_{i} m_{i}$ • $\underline{\ddot{\mathbf{r}}}_{CM} = \sum_{i} \frac{m_{i} \underline{\ddot{\mathbf{r}}}_{i}}{M}$ m_1 \mathbf{E}^{ext} \mathbf{r}_1 $\rightarrow M \ddot{\mathbf{r}}_{CM} = \sum_{i} \mathbf{F}_{i}^{ext}$

The motion of the system is equivalent to that of a single particle having a mass *M* acted on by the sum of external forces

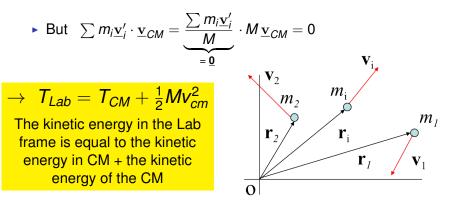
(The CM moves at constant velocity if no external forces)

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6.4 Kinetic energy and the CM

► Lab kinetic energy : $T_{Lab} = \frac{1}{2} \sum m_i \underline{\mathbf{v}}_i^2$; $\underline{\mathbf{v}}_i = \underline{\mathbf{v}}_i - \underline{\mathbf{v}}_{CM}$ where $\underline{\mathbf{v}}_i'$ is velocity of particle *i* in the CM

$$T_{Lab} = \frac{1}{2} \sum m_i \, \underline{\mathbf{v}}_i'^2 + \sum m_i \, \underline{\mathbf{v}}_i' \cdot \underline{\mathbf{v}}_{CM} + \frac{1}{2} \sum m_i \, \underline{\mathbf{v}}_{CM}^2$$



7.1 Two-body collisions - general concepts



No external forces. Collision via massless springs or other force type.



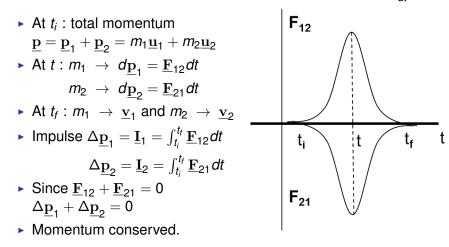


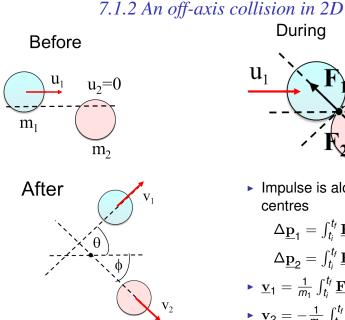


- ► t_i : collision starts. All energy is kinetic. $T_i = \frac{1}{2}m_1u_1^2 + \frac{1}{2}m_2u_2^2$
- ► t : collision peaks. Some kinetic is converted into potential (of the spring). $E = \frac{1}{2}m_1\dot{x_1}^2 + \frac{1}{2}m_2\dot{x_2}^2 + E_{int}$
- ► t_f : collision ends. All energy is kinetic again. $T_f = \frac{1}{2}m_1v_1^2 + \frac{1}{2}m_2v_2^2;$ $T_i = T_f + \Delta E$ (\leftarrow inelastic)

7.1.1 Momentum exchange and impulse

During collision: internal force causes change of momentum $\underline{\mathbf{F}} = \frac{d\mathbf{p}}{dt}$





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During

Impulse is along line of centres

$$\Delta \underline{\mathbf{p}}_{1} = \int_{t_{i}}^{t_{f}} \underline{\mathbf{F}}_{12} dt$$

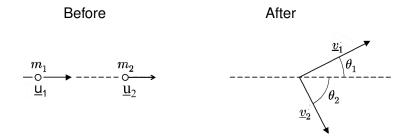
$$\Delta \underline{\mathbf{p}}_2 = \int_{t_i}^{t_f} \underline{\mathbf{F}}_{21} dt$$

$$\mathbf{v}_1 = \frac{1}{m_1} \int_{t_i}^{t_f} \mathbf{\underline{F}}_{12} dt + \mathbf{\underline{u}}_1$$

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$$\mathbf{v}_2 = -\frac{1}{m_2} \int_{t_i}^{t_f} \mathbf{F}_{12} dt$$

7.2 Elastic collisions in the Lab frame

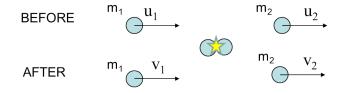


Conservation of momentum: $m_1 \underline{\mathbf{u}}_1 + m_2 \underline{\mathbf{u}}_2 = m_1 \underline{\mathbf{v}}_1 + m_2 \underline{\mathbf{v}}_2$

Conservation of energy: $\frac{1}{2}m_1\underline{\mathbf{u}}_1^2 + \frac{1}{2}m_2\underline{\mathbf{u}}_2^2 = \frac{1}{2}m_1\underline{\mathbf{v}}_1^2 + \frac{1}{2}m_2\underline{\mathbf{v}}_2^2$

[Note that the motion is in a plane, and the 2D representation can be trivially extended into 3D by rotation of the plane].

7.2.1 Elastic collisions in 1D in the Lab frame



• Momentum : $m_1 u_1 + m_2 u_2 = m_1 v_1 + m_2 v_2$ (1)

 $\rightarrow m_1(v_1-u_1) = m_2(u_2-v_2)$ (2)

- Energy : $\frac{1}{2}m_1u_1^2 + \frac{1}{2}m_2u_2^2 = \frac{1}{2}m_1v_1^2 + \frac{1}{2}m_2v_2^2$
 - $\rightarrow m_1(v_1-u_1)(v_1+u_1) = m_2(u_2-v_2)(u_2+v_2) \quad (3)$
- ► Divide (2) & (3) : \rightarrow ($v_1 + u_1$) = ($u_2 + v_2$) \rightarrow ($u_1 - u_2$) = ($v_2 - v_1$) (4)

 \rightarrow Relative speed before collision = Relative speed after

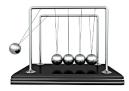
7.2.2 Special case in 1D where target particle is at rest

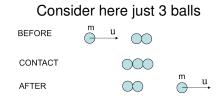
►
$$u_2 = 0$$
; From (1) & (4):
 $m_1 u_1 = m_1 v_1 + m_2(u_1 + v_1)$ BEFORE $m_1 u_1$ m_2 At rest
► $v_1 = \frac{(m_1 - m_2)u_1}{m_1 + m_2}$
► Similarly:
 $m_1 u_1 = m_1(v_2 - u_1) + m_2 v_2$ AFTER $m_1 u_1$ $m_2 v_2$
► $v_2 = \frac{2m_1 u_1}{m_1 + m_2}$
► $v_1 = \frac{m_1 - m_2}{m_1 + m_2} u_1$ and $v_2 = \frac{2m_1}{m_1 + m_2} u_1$

Special cases:

- $m_1 = m_2$: \rightarrow $v_1 = 0, v_2 = u_1$ (complete transfer of momentum)
- $m_1 >> m_2$: Gives the limits $v_1 \rightarrow u_1, v_2 \rightarrow 2u_1$ (m_2 has double u_1 velocity)
- ▶ $m_1 \ll m_2$: Gives the limits $v_1 \rightarrow -u_1, v_2 \rightarrow 0$ ("brick wall" collision)

Example: Newton's cradle





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- ► If the balls are touching, the most general case is: Momentum after collision : $mu = mv_1 + mv_2 + mv_3$ Energy after collision : $\frac{1}{2}mu^2 = \frac{1}{2}mv_1^2 + \frac{1}{2}mv_2^2 + \frac{1}{2}mv_3^2$ 2 equations, 3 unknowns
- The obvious solution: $v_1 = v_2 = 0$, $v_3 = u$
- ▶ But other solution(s) possible: Momentum : mu = -¹/₃mu + ²/₃mu + ²/₃mu Energy : ¹/₂mu² = ¹/₁₈mu² + ⁴/₁₈mu² + ⁴/₁₈mu²
 ▶ So why does the simple solution always prevail?

7.2.3 Collision in 2D : equal masses, target at rest

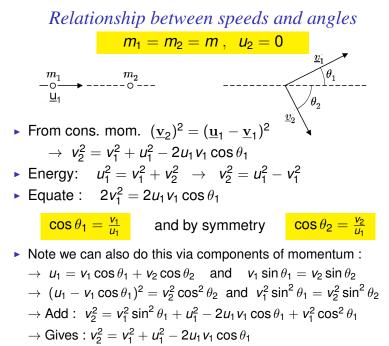
$$m_1 = m_2 = m , \ u_2 = 0$$

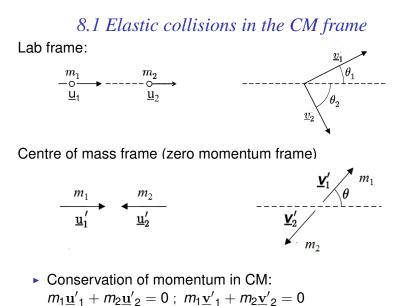


- ► Momentum: $m\underline{\mathbf{u}}_1 = m\underline{\mathbf{v}}_1 + m\underline{\mathbf{v}}_2 \rightarrow \underline{\mathbf{u}}_1 = \underline{\mathbf{v}}_1 + \underline{\mathbf{v}}_2$ Squaring $\rightarrow \underline{\mathbf{u}}_1^2 = \underline{\mathbf{v}}_1^2 + \underline{\mathbf{v}}_2^2 + 2\underline{\mathbf{v}}_1 \cdot \underline{\mathbf{v}}_2$
- Energy: $\frac{1}{2}m\underline{\mathbf{u}}_1^2 = \frac{1}{2}m\underline{\mathbf{v}}_1^2 + \frac{1}{2}m\underline{\mathbf{v}}_2^2 \rightarrow \underline{\mathbf{u}}_1^2 = \underline{\mathbf{v}}_1^2 + \underline{\mathbf{v}}_2^2$
- Hence $2\underline{\mathbf{v}}_1 \cdot \underline{\mathbf{v}}_2 = 0$

 \rightarrow EITHER $\underline{\mathbf{v}}_1 = \mathbf{0} \& \underline{\mathbf{v}}_2 = \underline{\mathbf{u}}_1 \text{ OR } (\theta_1 + \theta_2) = \frac{\pi}{2}$

Either a head-on collision or opening angle is 90°

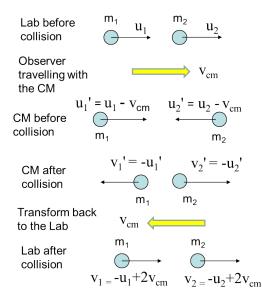




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• Conservation of energy in CM: $\frac{1}{2}m_1\underline{\mathbf{u}'}_1^2 + \frac{1}{2}m_2\underline{\mathbf{u}'}_2^2 = \frac{1}{2}m_1\underline{\mathbf{v}'}_1^2 + \frac{1}{2}m_2\underline{\mathbf{v}'}_2^2$

8.2 Lab to CM : 2-body 1D elastic collision



- $V_{CM} = \frac{(m_1 u_1 + m_2 u_2)}{(m_1 + m_2)}$
- Before in CM :
 - $m_1u_1' + m_2u_2' = 0$
- After in CM : $m_1 v'_1 + m_2 v'_2 = 0$
- From last lecture $u'_1 u'_2 = v'_2 v'_1$
- Sub for u'_2, v'_2 :
 - $u_1'(1 + m_1/m_2) =$
 - $-v_1'(1+m_1/m_2)$

•
$$V'_1 = -U'_1$$

 $V'_2 = -U'_2$

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8.2.1 Collision in 1D : numerical example $m_1 = 6 \ kg \qquad \text{Lab} \qquad m_2 = 4 \ kg \qquad u_2 = 0.5 \ ms^{-1}$ 1) Find CM velocity relative to laboratory frame : $v_{cm} = \frac{m_1 u_1 + m_2 u_2}{m_1 + m_2} = \frac{6 \times 3 + 0.5 \times 4}{10} = 2 \ ms^{-1}$

2) Transform initial velocities into CM : $u' = u - v_{cm}$

$$u'_{1} = 3 - 2 = 1 \text{ ms}^{-1}$$
; $u'_{2} = 0.5 - 2 = -1.5 \text{ ms}^{-1}$
 $v_{cm} = 2 \text{ ms}^{-1}$
 $u'_{1} = 1 \text{ ms}^{-1}$
frame
 $u'_{2} = -1.5 \text{ ms}^{-1}$

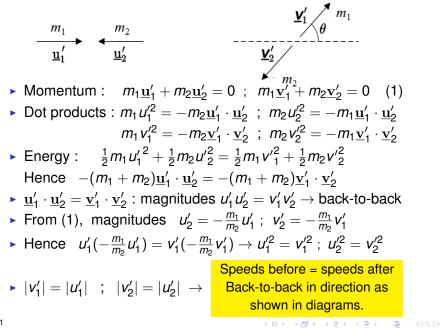
3) Conservation of energy : $v'_1 = -u'_1$: $v'_2 = -u'_2$

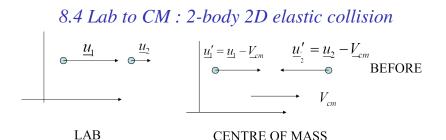
 $v_1' = -1 \text{ ms}^{-1}$; $v_2' = 1.5 \text{ ms}^{-1}$ after collision

4) Transform final velocities back to Laboratory frame : $v = v' + v_{cm}$ $v_1 = -1 + 2 = 1 \text{ ms}^{-1}$; $v_2 = 1.5 + 2 = 3.5 \text{ ms}^{-1}$

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8.3 Relationship between speeds in CM in 2D





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1) Find centre of mass velocity v_{CM}

$$(\underline{\mathbf{u}}_1 - \underline{\mathbf{v}}_{CM})m_1 + (\underline{\mathbf{u}}_2 - \underline{\mathbf{v}}_{CM})m_2 = 0 \rightarrow \underline{\mathbf{v}}_{CM} = \frac{m_1\underline{\mathbf{u}}_1 + m_2\underline{\mathbf{u}}_2}{m_1 + m_2}$$

2) Transform initial Lab velocities to CM

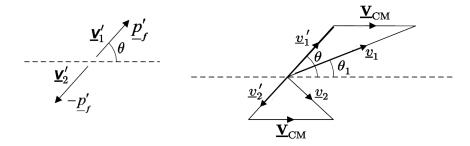
•
$$\underline{\mathbf{u}}_1' = \underline{\mathbf{u}}_1 - \underline{\mathbf{v}}_{CM}$$
 , $\underline{\mathbf{u}}_2' = \underline{\mathbf{u}}_2 - \underline{\mathbf{v}}_{CM}$

3) Get final CM velocities

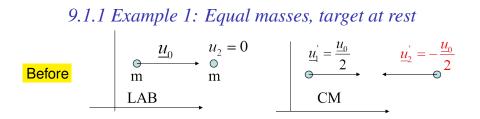
$$\bullet |\underline{\mathbf{v}}_1'| = |\underline{\mathbf{u}}_1'| \quad ; \quad |\underline{\mathbf{v}}_2'| = |\underline{\mathbf{u}}_2'|$$

4) Transform vectors back to the Lab frame

$$\bullet \underline{\mathbf{v}}_1 = \underline{\mathbf{v}}_1' + \underline{\mathbf{v}}_{CM} \quad ; \quad \underline{\mathbf{v}}_2 = \underline{\mathbf{v}}_2' + \underline{\mathbf{v}}_{CM}$$



5) Can then use trigonometry to solve



Magnitude of velocities:

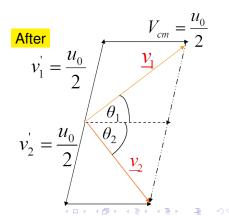
$$V_{CM} = \frac{m_1 u_1 + m_2 u_2}{m_1 + m_2} = \frac{u_0}{2}$$

$$U'_1 = u_0 - V_{CM} = \frac{u_0}{2}$$

$$U'_2 = -V_{CM} = -\frac{u_0}{2}$$

$$|v'_1| = |u'_1| = \frac{u_0}{2}$$

$$|v'_2| = |u'_2| = \frac{u_0}{2}$$



Relationships between angles and speeds

Angles:

Cosine rule:

$$(\frac{u_0}{2})^2 = (\frac{u_0}{2})^2 + v_1^2 - 2v_1\frac{u_0}{2}\cos\theta_1$$

- $v_1 u_0 \cos \theta_1 = v_1^2$
- $\cos \theta_1 = \frac{v_1}{u_0}$ as before
- $\cos \theta_2 = \frac{v_2}{u_0}$

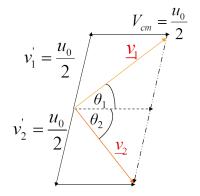
Opening angle:

Cosine rule:

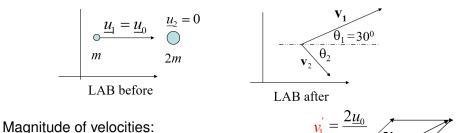
$$u_0^2 = v_1^2 + v_2^2 - 2v_1v_2\cos(\theta_1 + \theta_2)$$

- But u₀² = v₁² + v₂² (conservation of energy)
- $\cos(\theta_1 + \theta_2) = 0 \rightarrow \theta_1 + \theta_2 = \frac{\pi}{2}$

NB: Lines joining opposite corners of rhombus cross at 90°



9.1.2 *Example 2: Elastic collision*, $m_2 = 2m_1$, $\theta_1 = 30^{\circ}$ Find the velocities v_1 and v_2 and the angle θ_2



$$V_{CM} = \frac{m_1 u_1 + m_2 u_2}{m_1 + m_2} = \frac{u_0}{3}$$

$$U'_1 = u_0 - V_{CM} = \frac{2u_0}{3}$$

$$U'_2 = -V_{CM} = -\frac{u_0}{3}$$

$$|V'_1| = |U'_1| = \frac{2u_0}{3}$$

$$|V'_2| = |U'_2| = \frac{u_0}{3}$$

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Relationships between angles and speeds

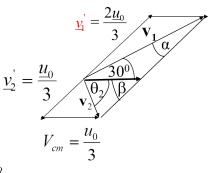
• Sine rule: $(\sin 30/\frac{2u_0}{3}) = (\sin \alpha/\frac{u_0}{3})$ $\rightarrow \sin \alpha = \frac{1}{4} \rightarrow \alpha = 14.5^{\circ}$ • $\beta = 30 + \alpha = 44.5^{\circ}$ • $\sin 30/\frac{2u_0}{3} = \sin(180 - 44.5)/v_1$ $\rightarrow v_1 = 0.93u_0$ • Cosine rule: $v_2^2 = (\frac{u_0}{2})^2 + (\frac{u_0}{2})^2 - 2(\frac{u_0}{2})^2 \cos \beta$

$$v_2^2 = (\frac{u_0}{3})^2 + (\frac{u_0}{3})^2 - 2(\frac{u_0}{3})^2 \cos \theta$$

$$\rightarrow v_2 = 0.25u_0$$

Sine rule:

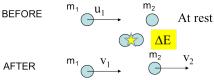
 $(\sin 44.5/v_2) = (\sin \theta_2/\frac{u_0}{3})$ $\rightarrow \ \theta_2 = 68.0^{\circ}$



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9.2 Inelastic collisions in the Lab frame in 1D ($u_2 = 0$)

An *inelastic* collision is where energy is lost (or there is internal excitation).



- Take m_2 at rest & in 1D. Momentum : $m_1u_1 = m_1v_1 + m_2v_2$ (1)
- Energy : $\frac{1}{2}m_1u_1^2 = \frac{1}{2}m_1v_1^2 + \frac{1}{2}m_2v_2^2 + \Delta E$ (2)
- Square Equ.(1) and subtract $2m_1 \times$ Equ.(2)

$$\rightarrow \quad m_2(m_2-m_1)v_2^2+2m_1m_2v_1v_2-2m_1\Delta E=0$$

Substitute for m_1v_1 from Equ.1 to get quadratic in v_2

$$\rightarrow \quad m_2(m_2+m_1)v_2^2-2m_1m_2u_1v_2+2m_1\Delta E=0$$

Solve, taking consistent solutions with elastic case ($\Delta E = 0$)

$$\rightarrow V_{2} = \frac{2m_{1}m_{2}u_{1} + \sqrt{4m_{1}^{2}m_{2}^{2}u_{1}^{2} - 8m_{1}m_{2}(m_{1} + m_{2})\Delta E}}{2m_{2}(m_{1} + m_{2})}$$
(3)

$$\rightarrow V_{1} = \frac{2m_{1}^{2}u_{1} - \sqrt{4m_{1}^{2}m_{2}^{2}u_{1}^{2} - 8m_{1}m_{2}(m_{1} + m_{2})\Delta E}}{2m_{1}(m_{1} + m_{2})}$$
(4)

1D inelastic collisions viewed in the Lab frame ($u_2 = 0$)

• We see from Equ. (3) & (4) there is a *limiting case*: $4m_1^2m_2^2u_1^2 - 8m_1m_2(m_1 + m_2)\Delta E \ge 0$

• i.e.
$$\Delta E \leq \frac{m_1 m_2 u_1^2}{2(m_1 + m_2)}$$

- ► This corresponds to the two bodies sticking together in a single object of mass (m₁ + m₂) → v₁ = v₂
- From momentum cons. $m_1u_1 = m_1v_1 + m_2v_2$

if
$$v_1 = v_2 = v$$
, then $v = \frac{m_1 u_1}{(m_1 + m_2)}$ (the CM velocity)

For equal mass $m_1 = m_2$

$$v_2 \ , \ v_1 = rac{u_1}{2} \left[1 \pm \sqrt{1 - rac{4\Delta E}{mu_1^2}}
ight]$$

9.2.1 Coefficient of restitution

General definition : $e = \frac{|\underline{\mathbf{v}}_2 - \underline{\mathbf{v}}_1|}{|\underline{\mathbf{u}}_1 - \underline{\mathbf{u}}_2|} = \frac{\text{Speed of relative separation}}{\text{Speed of relative approach}}$

From Equ.(3) & (4) previously

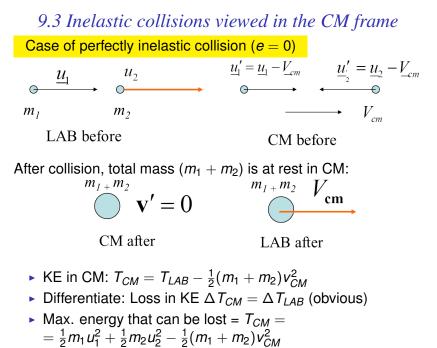
$$v_2 - v_1 = \frac{2m_1m_2u_1 + \sqrt{4m_1^2m_2^2u_1^2 - 8m_1m_2(m_1 + m_2)\Delta E}}{2m_2(m_1 + m_2)} - \frac{2m_1^2u_1 - \sqrt{4m_1^2m_2^2u_1^2 - 8m_1m_2(m_1 + m_2)\Delta E}}{2m_1(m_1 + m_2)}$$

Factorizing, then simplifying, then dividing by u₁ gives

$$e = \sqrt{1 - rac{2(m_1 + m_2)\Delta E}{m_1 m_2 u_1^2}} = \sqrt{1 - rac{\Delta E}{T'}}$$

where $T' = \frac{1}{2}\mu u_1^2$ with $\mu = \frac{m_1 m_2}{m_1 + m_2}$ (the *reduced mass*)

- We see later that T' is the initial energy in the CM frame, hence e is related to the fractional energy loss in this frame
- ► e = 1 completely elastic; e = 0 perfectly inelastic, in general 0 < e < 1</p>



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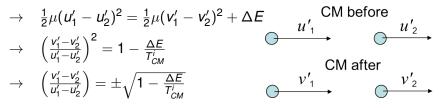
9.3.1 Kinetic energy in the CM : alternative treatment Revisit kinetic energy in the CM frame: $T_{Lab} = T_{CM} + \frac{1}{2}Mv_{CM}^2$ • $T_{CM} = \frac{1}{2}m_1u_1^{\prime 2} + \frac{1}{2}m_2u_2^{\prime 2}$ • $X'_1 = -\frac{m_2}{m_1 + m_2} X = -\frac{m_2}{M} X$, $X'_2 = \frac{m_1}{M} X$ $m_1 x'_1 x'_2$ m_2 • $U'_1 = -\frac{m_2}{M}\dot{X}$, $U'_2 = \frac{m_1}{M}\dot{X}$ \mathbf{X}_1 CM • $T_{CM} = \frac{1}{2} \left(m_1 \left(-\frac{m_2}{M} \right)^2 + m_2 \left(\frac{m_1}{M} \right)^2 \right) \dot{x}^2$ • $T_{CM} = \frac{1}{2} \frac{m_1 m_2}{M^2} (m_2 + m_1) \dot{x}^2 = \frac{1}{2} \frac{m_1 m_2}{M} \dot{x}^2$ • Also $\dot{x} = \dot{x}_2 - \dot{x}_1 = u_2' - u_1'$

$$T_{CM} = \frac{1}{2} \frac{m_1 m_2}{M} \dot{x}^2 = \frac{1}{2} \mu \dot{x}^2 = \frac{1}{2} \mu (u_1' - u_2')^2 = \frac{1}{2} \mu (u_1 - u_2)^2$$

These expressions give the CM kinetic energy in terms of the relative velocities in the CM & Lab and the reduced mass μ

9.3.2 Coefficient of restitution in the CM

- Initial KE in the CM : $T_{CM}^{i} = \frac{1}{2}\mu(u_{1}^{\prime} u_{2}^{\prime})^{2}$
- Final KE in the CM : $T_{CM}^{f} = \frac{1}{2}\mu(v_{1}' v_{2}')^{2}$
- Conservation of energy : $T_{CM}^{i} = T_{CM}^{f} + \Delta E$



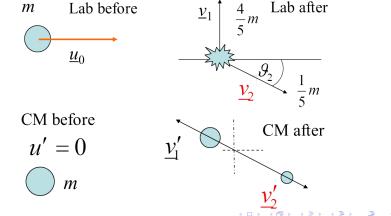
Same expression as before with $T' = T_{CM}^{i}$

Coefficient of restitution

$$\boldsymbol{e} = \frac{|\underline{\mathbf{v}}_2' - \underline{\mathbf{v}}_1'|}{|\underline{\mathbf{u}}_1' - \underline{\mathbf{u}}_2'|_{CM}} = \frac{|\underline{\mathbf{v}}_2 - \underline{\mathbf{v}}_1|}{|\underline{\mathbf{u}}_1 - \underline{\mathbf{u}}_2|_{LAB}} = \sqrt{1 - \frac{\Delta E}{T_{CM}^i}}$$

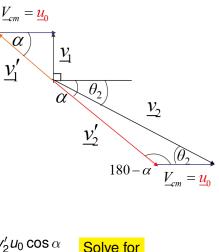
9.3.3 Example of inelastic process

A calcium nucleus (A=20), mass *m*, travels with velocity u_0 in the Lab. It decays into a sulphur nucleus (A=16), mass $\frac{4}{5}m$, and an α -particle (A=4), mass $\frac{1}{5}m$. Energy ΔT is released as KE in the calcium rest frame (CM). A counter in the Lab detects the sulphur nucleus at 90° to the line of travel. What is the speed and angle of the α -particle in the Lab?



- Energy ΔT is released as KE in the CM. $v_{CM} = u_0$
- ► Momentum in CM: $\frac{4}{5}mv'_{1} - \frac{1}{5}mv'_{2} = 0$ $\rightarrow v'_{2} = 4v'_{1}$
- Energy: $\Delta T = \frac{1}{2} (\frac{4}{5}m) v_1'^2 + \frac{1}{2} (\frac{1}{5}m) 16 v_1'^2 = 2m v_1'^2$ $\rightarrow v_1' = [\frac{\Delta T}{2m}]^{\frac{1}{2}}$ $\rightarrow v_2' = [\frac{8\Delta T}{m}]^{\frac{1}{2}}$
- Transform to Lab by boosting by v_{CM}(= u₀)
- $\cos \alpha = \frac{u_0}{v_1'} = \left[\frac{2mu_0^2}{\Delta T}\right]^{\frac{1}{2}}$
- Cosine rule: $v_2^2 = {v_2'}^2 + u_0^2 + 2v_2'u_0 \cos \alpha$

• Sine rule:
$$\frac{\sin \theta_2}{v'_2} = \frac{\sin \alpha}{v_2}$$



 V_2 , θ_2

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10.1 Resisted motion and limiting speed

- ► Newton II: $m\frac{dv}{dt} = F_{ext} + F_R$ where F_{ext} is the external force and F_R is a resistive force
- ► If $F_{ext} = 0$ and $F_R \propto$ velocity, then $v \propto \exp(-\alpha t)$ (see Lecture 4)
- ► If $F_{ext} \neq 0$ and e.g. $F_R \propto -v^n$ then there exists a *limiting speed* corresponding to $\frac{dv}{dt} = 0$ that satisfies $F_R = -F_{ext}$

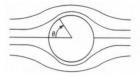
10.2 Air resistance



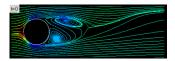
• Laminer flow : Stoke's Law $F = 6\pi\eta rv$

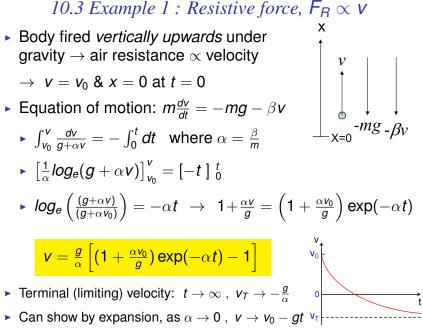
r is the radius of the sphere v is the velocity of the sphere η is the viscosity of the fluid

► Turbulent flow : S $F = \frac{1}{2}\pi\rho C_d r^2 v^2$ ρ is the density of the fluid C_d is the drag coefficient (e.g. for a smooth sphere $C_d \sim 0.47$









Maximum height and distance travelled for $F_R \propto v$

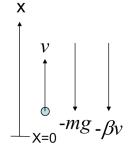
$$\mathbf{v} = \frac{g}{\alpha} \left[(\mathbf{1} + \frac{\alpha \mathbf{v}_0}{g}) \exp(-\alpha t) - \mathbf{1} \right]$$

• At maximum height $\rightarrow v = 0, t = t_{max}$

$$\rightarrow \exp(-\alpha t_{max}) = (1 + \frac{\alpha v_0}{g})^{-1}$$

$$t_{max} = rac{1}{lpha} \log_{m{ extsf{ extsf} extsf{ extsf} extsf{ extsf{ extsf} extsf} extsf{ extsf} extsf} extsf} extsf{ extsf{ extsf extsf{ extsf ex} extsf{ extsf} extsf{ extsf{ extsf{ extsf{ extsf} extsf{ e$$

Can expand log to show : $t_{max} \rightarrow \frac{v_0}{g}$ when $\alpha \rightarrow 0$



► Distance travelled :
►
$$x = \int_0^t \frac{g}{\alpha} \left[(1 + \frac{\alpha v_0}{g}) \exp(-\alpha t) - 1 \right] dt$$

 $= \frac{g}{\alpha} \left[-\frac{1}{\alpha} (1 + \frac{\alpha v_0}{g}) \exp(-\alpha t) - t \right]_0^t$
Can show by
expansion
 $x \to v_0 t - \frac{1}{2}gt^2$
when $\alpha \to 0$

10.4 Example 2: Resistive force, $F_R \propto v^2$

X=0

- ▶ Body falls *vertically downwards* under gravity with air resistance \propto [velocity]², v = 0, x = 0 at t = 0
- Equation of motion: $m\frac{dv}{dt} = mg \beta v^2$
- Terminal velocity when $\frac{dv}{dt} = 0$: $v_T = \sqrt{\frac{mg}{\beta}}$
 - Equation of motion becomes $\frac{dv}{dt} = g \left(1 v^2 / v_T^2\right)$

• Integrate
$$\int_0^v \frac{dv}{g(1-v^2/v_T^2)} = \int_0^t dt$$

• Standard integral : $\int \frac{1}{1-z^2} dz = \frac{1}{2} \log_e \left(\frac{1+z}{1-z}\right)$

$$\begin{bmatrix} \frac{v_T}{2g} \log_e \left(\frac{1+v/v_T}{1-v/v_T} \right) \end{bmatrix}_0^v = t \quad \rightarrow \quad \frac{1+v/v_T}{1-v/v_T} = \exp(t/\tau) \quad , \quad \text{where } \tau = \frac{v_T}{2g} \\ \rightarrow \quad (1 - \frac{v}{v_T}) = (1 + \frac{v}{v_T}) \exp(-\frac{t}{\tau})$$

Velocity as a function of time:

$$\mathbf{V} = \mathbf{V}_{\mathcal{T}} \begin{bmatrix} \frac{1 - \exp(-t/\tau)}{1 + \exp(-t/\tau)} \end{bmatrix}$$

Velocity as a function of distance for $F_R \propto v^2$

• Equation of motion: $\frac{dv}{dt} = g \left(1 - v^2 / v_T^2\right)$

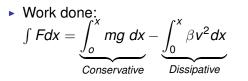
• Write
$$\frac{dv}{dt} = \frac{dv}{dx}\frac{dx}{dt} = v\frac{dv}{dx}$$

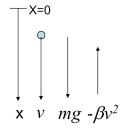
• $\int_0^v \frac{v \, dv}{g(1-v^2/v_T^2)} = \int_0^x dx$
• $\left[-\frac{v_T^2}{2g} \log_e \left(1 - v^2/v_T^2\right)\right]_0^v = x$
 $\rightarrow (1 - v^2/v_T^2) = \exp\left(-x/x_T\right)$, where $x_T = \frac{v_T^2}{2g}$
 $v^2 = v_T^2 \left[1 - \exp\left(-x/x_T\right)\right]$

To get x vs. t integrate again : $\rightarrow \int_0^t dt = \int_0^x \frac{dx}{v}$

10.4.1 Work done on the body by the force for $F_R \propto v^2$

• Equation of motion: $m\frac{dv}{dt} = mg - \beta v^2$





Conservative term : Work done = mgx

Dissipative term : Work done

$$= -\int_0^x \beta v^2 dx = -\int_0^x \beta v_T^2 [1 - \exp(-x/x_T)] dx \qquad v_T^2 = \frac{mg}{\beta}$$

$$= -\beta v_T^2 [x + \underbrace{x_T (\exp(-x/x_T) - 1)}_{= -x_T v^2/v_T^2}] \qquad x_T = \frac{v_T^2}{2g}$$

$$= -\beta v_T^2(x - v^2/2g) = -mg[x - v^2/2g]$$

Energy dissipated = $\frac{1}{2}mv^2 - mgx$ As expected.