

PHYS 666: Solid State Physics I

INSTRUCTOR

Michel van Veenendaal

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Office Hours: I am around Tuesdays and Thursdays

Web page with lecture notes: www.niu.edu/~veenendaal/666.htm

PREREQUISITES:

This course will be tough without having done quantum mechanics (560/1 or something equivalent)

Mathematical concepts:

Fourier transforms

differential equations (Schrödinger equations)

linear algebra (matrices, eigenvalue problems)

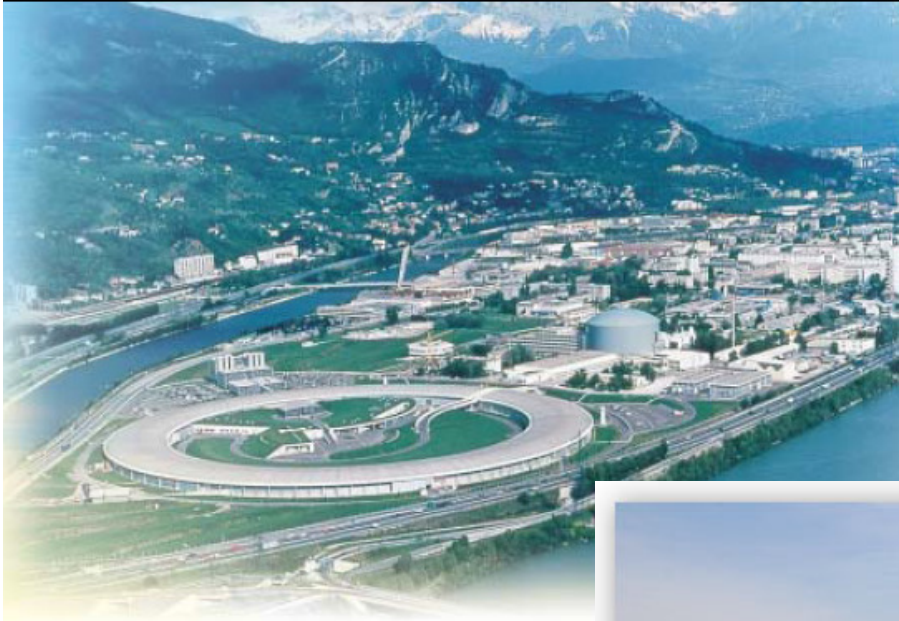
- Homework: several homework sets will be given. They will be posted on the web site.
- Midterm: one midterm will be given.

- 1986-90: undergraduate Delft University of Technology (Ir.), the Netherlands
- 1990-94: Ph.D. University of Groningen (Dr.), the Netherlands
- 1994-97: European Synchrotron Radiation Facility, Grenoble, France
- 1997-98: NIU
- 1998-2002: Philips Electronics
- 2002-present: NIU
 - 2002-2008 Associate Professor
 - 2008- Professor
 - 2009-2013 Presidential Research Professor
- 2005-present: joint with Argonne National Laboratory (Physicist)

Theoretical physicist

specialty: condensed matter physics

ESRF



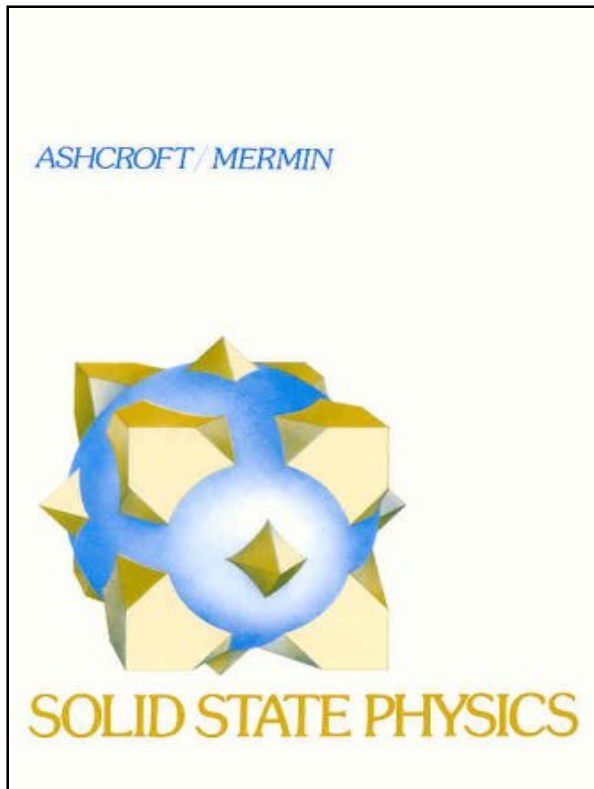
APS



REQUIRED TEXTBOOK:

[Solid State Physics](#)

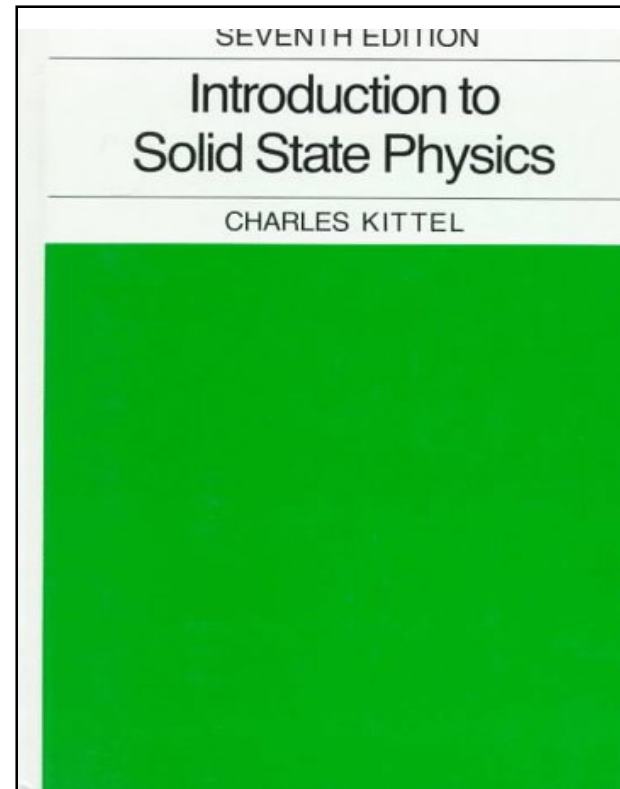
by N. W. Ashcroft and N. D. Mermin
(Harcourt, 1976).



REFERENCED TEXTBOOK

[Introduction to Solid State Physics](#)

[7th Edition](#) by C. Kittel
(John Wiley & Sons, 1996).



Solid State Physics vs. Condensed-Matter Physics

- Condensed-matter physics is the more modern term
- Condensed-matter physics is broader and applies to concepts that work in solids, but could equally applied to liquid (for example, superconductivity vs. superfluidity, soft-condensed matter)
- 1978 Division of Solid-State Physics of the American Physical Society went to the Division of Condensed-Matter Physics
- 1/3 of U.S. physicists classify themselves as Condensed-Matter Physicists
- Condensed-matter physics is closely related and overlaps with inorganic chemistry, physical chemistry, quantum chemistry, electrical and mechanical .

WHY DO WE WANT TO DO SOLID STATE PHYSICS?

<http://www.er.doe.gov/bes/reports/list.html>



Steve Chu,
Secretary of Energy
1997 Nobel Prize in Physics



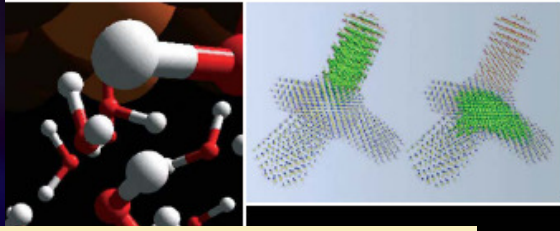
- A significant increase in the rate of discovery, innovation and technological change is needed. BES must lead a bold new initiative focused on solving the critical scientific roadblocks in next-generation, carbon-free energy technologies.
- Significant discoveries will come at the intersection of control science with advanced materials and chemical phenomena, and there is a clear first-mover advantage to those who focus their research efforts here. BES must lead U.S. energy research efforts in this direction lest the U.S. fall behind in global competition for discoveries of future energy sources and systems.
- It will take “dream teams” of highly educated talent, equipped with forefront tools, and focused on the most pressing challenges to increase the rate of discovery. BES must lead the development of these dream teams to close gaps between needs and capabilities in synthesis, characterization, theory, and computation of advanced materials.
- U.S. leadership requires BES to lead a national effort to aggressively recruit the best talent through a series of workforce development and early career programs aimed at inspiring today’s students and young researchers to be the discoverers, inventors, and innovators of tomorrow’s energy solutions.

Reports from
Basic Energy Sciences
Department of Energy

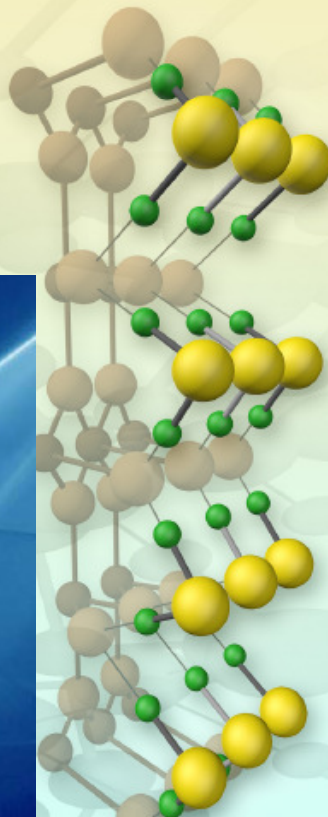
**Directing Matter and Energy:
Five Challenges for Science and the Imagination**

**Opportunities for Discovery:
Theory and Computation
in Basic Energy Sciences**

**Basic Research Needs
for Solar Energy Utilization**

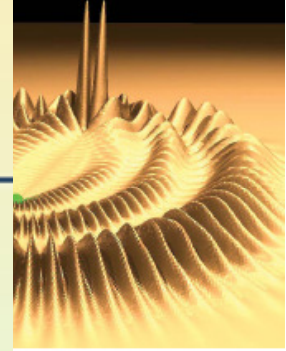


Report
Science
Solar
April



**Basic Research
Needs for
Electrical
Energy Storage**

Report of the Basic Energy
Sciences Workshop on
Electrical Energy Storage
April 2-4, 2007

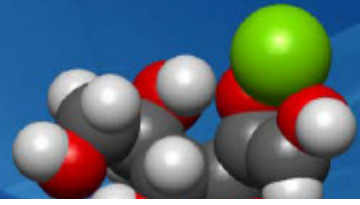


2005



**BASIC RESEARCH NEEDS:
CATALYSIS FOR ENERGY**

Report from the
U.S. Department of Energy
Basic Energy Sciences Workshop
August 6-8, 2007



Imagining a Secure, Sustainable Energy Future

Materials with Unprecedented Performance

Making Chemical Change More Selective

Solar Power

Returning Carbon to the Earth

Safer and More Efficient Nuclear Power

Solar Fuels

Let There Be (Digital) Light

Electrical Energy Storage

A Solar Economy for Buildings

Superconductivity and the 21st Century Electric Grid

A Hybrid Electrical Grid

Solid-State Lighting

Advanced Transport

Solar Fuels

Electric Transport

Advanced Nuclear Energy Systems

Geological Carbon Sequestration

Making the Stories Come True

Basic Materials Research to Commercial Impact

3

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Nanoscience

Catalysis, photosynthesis

Batteries

Superconductivity

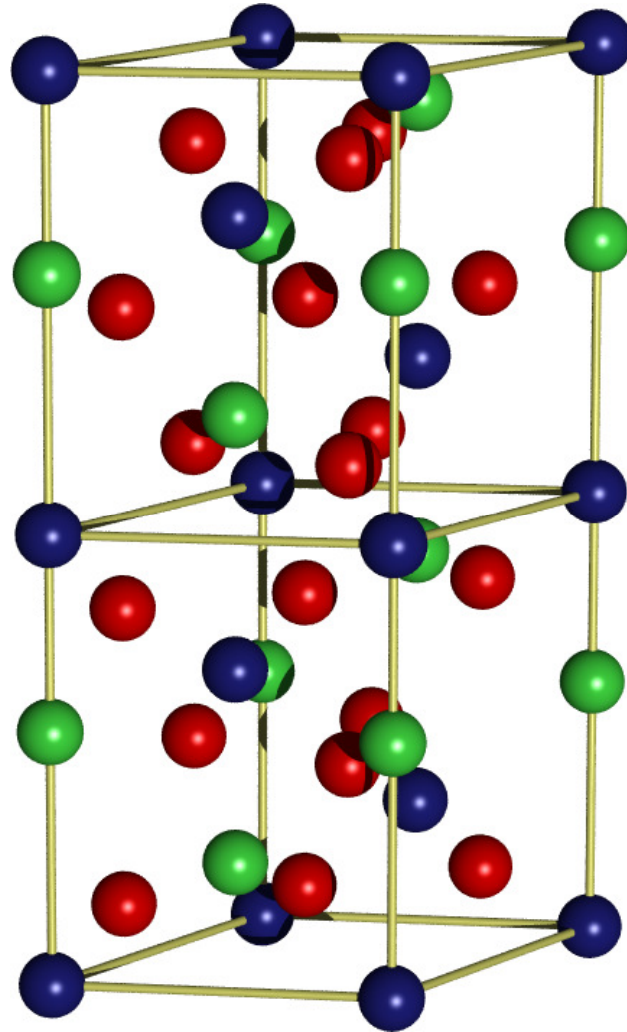
Solid-state lighting

Solar cells

New electrode materials

WHAT IS A SOLID?

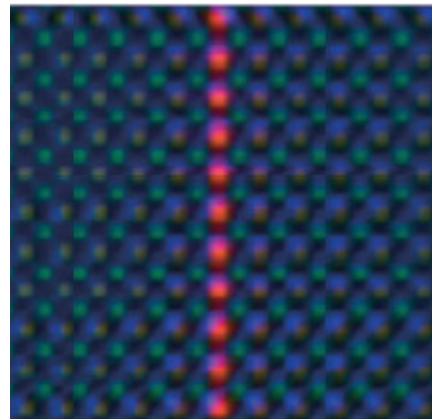
A solid is generally seen as a nice crystal made up of atom



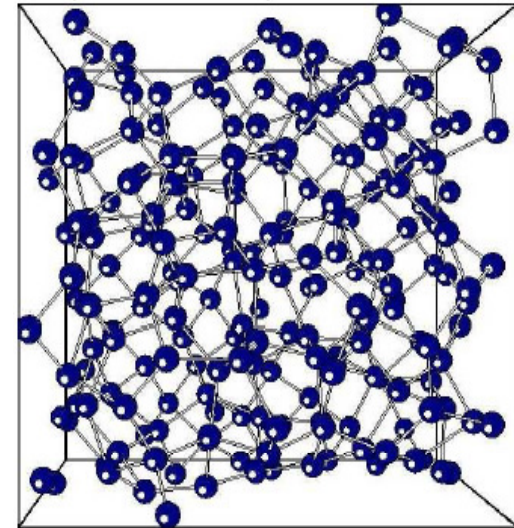
And we will generally be dealing with those

BUT ALSO

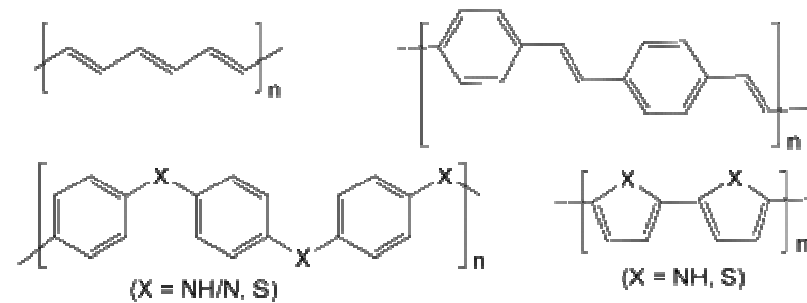
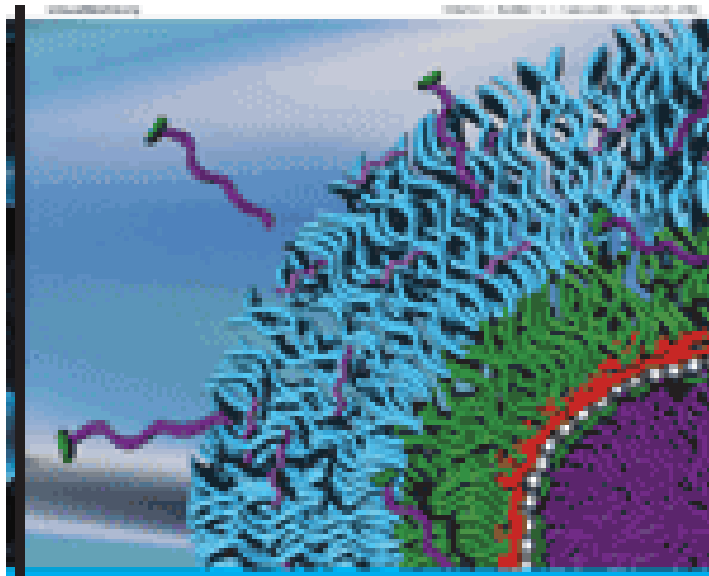
heterostuctures



Amorphous materials

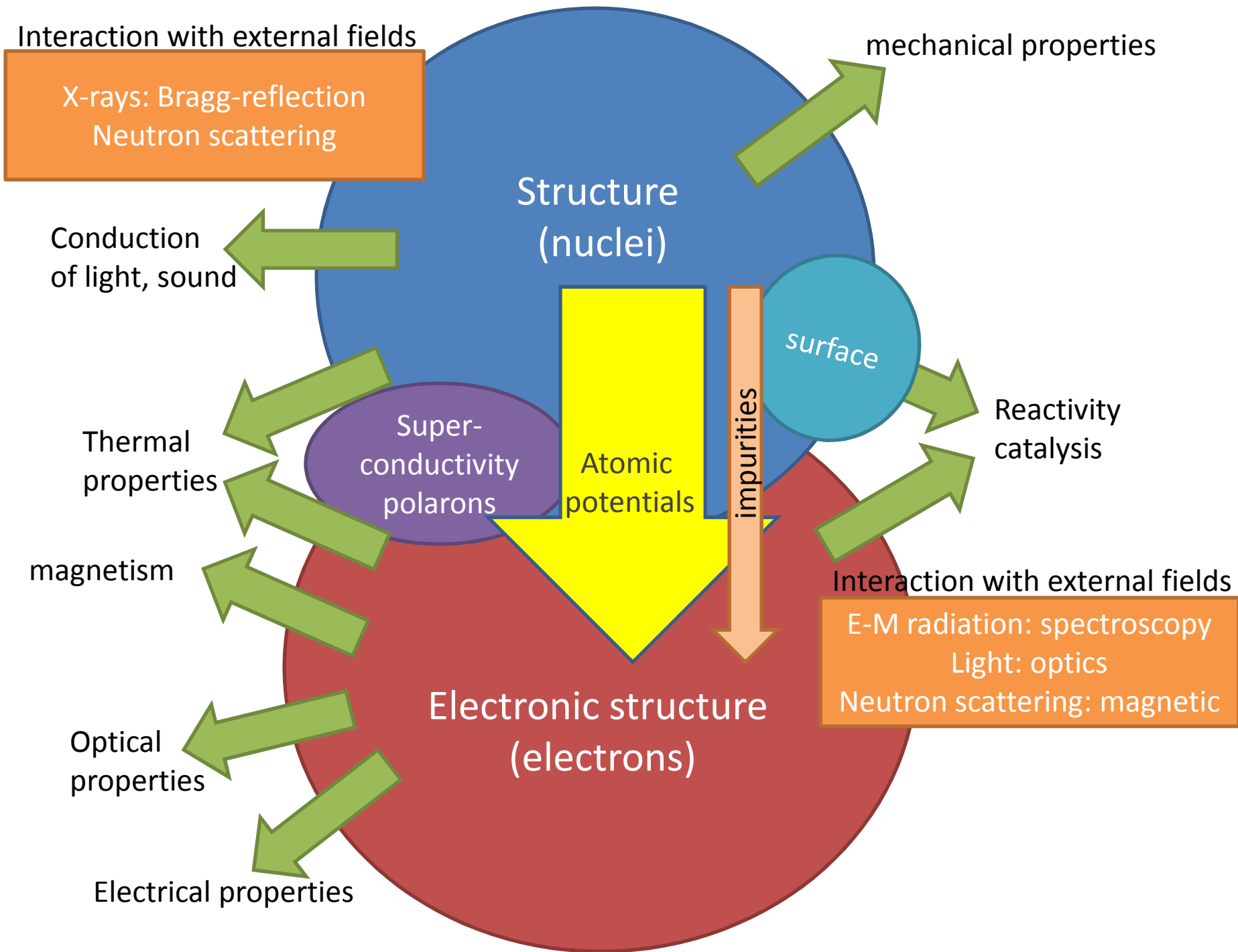


Soft condensed matter



Conducting polymers

Before you are able to deal with this, we have to get through the basics



HISTORY OF SOLIDS....

Whole ages are classified by our ability to control solids



Stone age

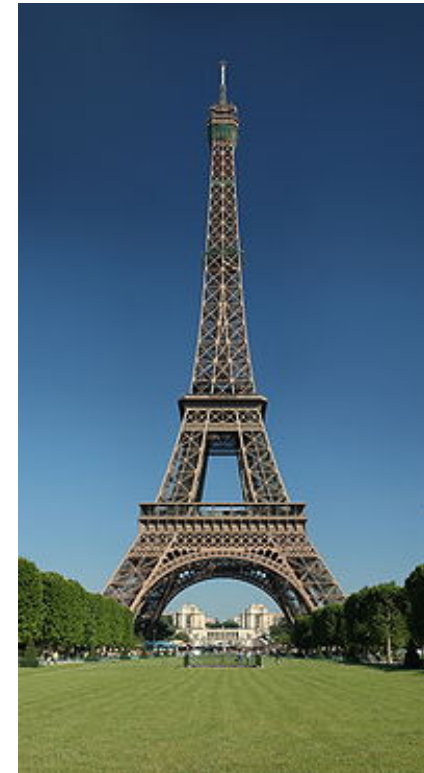


17th century BC.



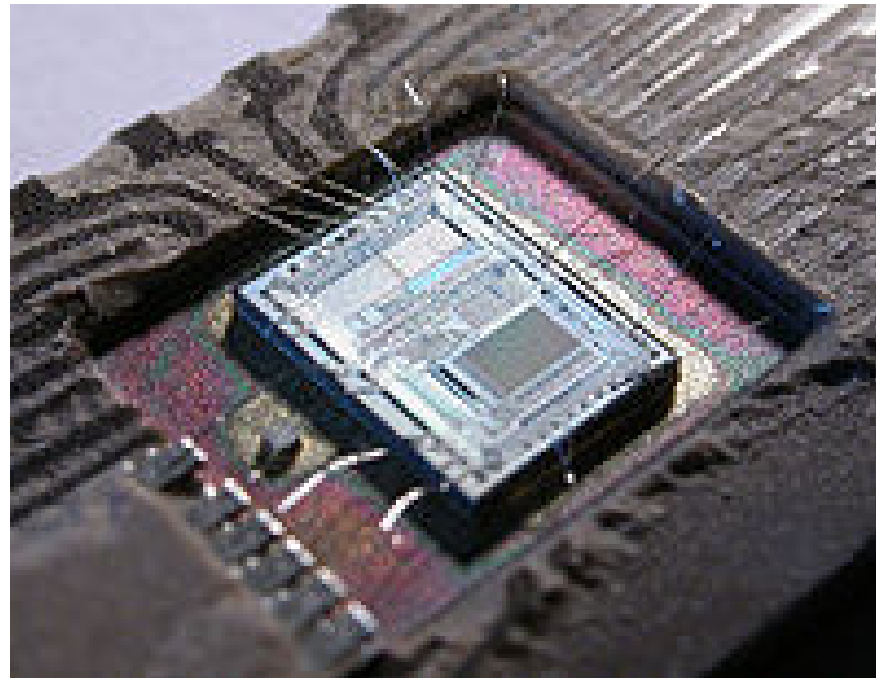
China (1600–1046 BC).

bronze age
(3300–1200 BC)

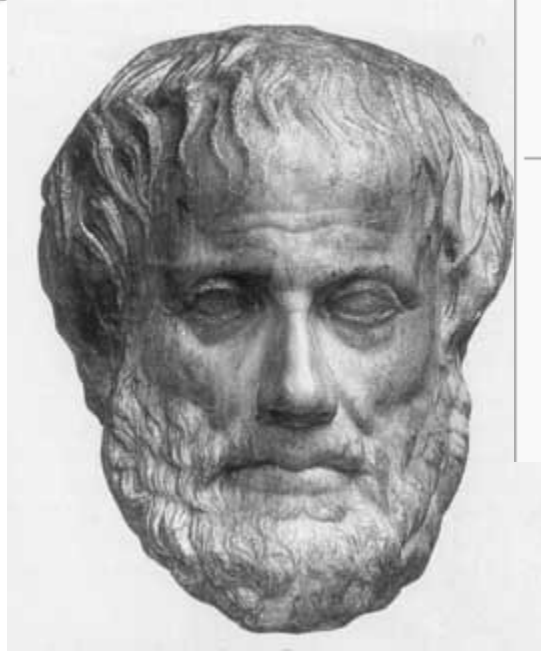
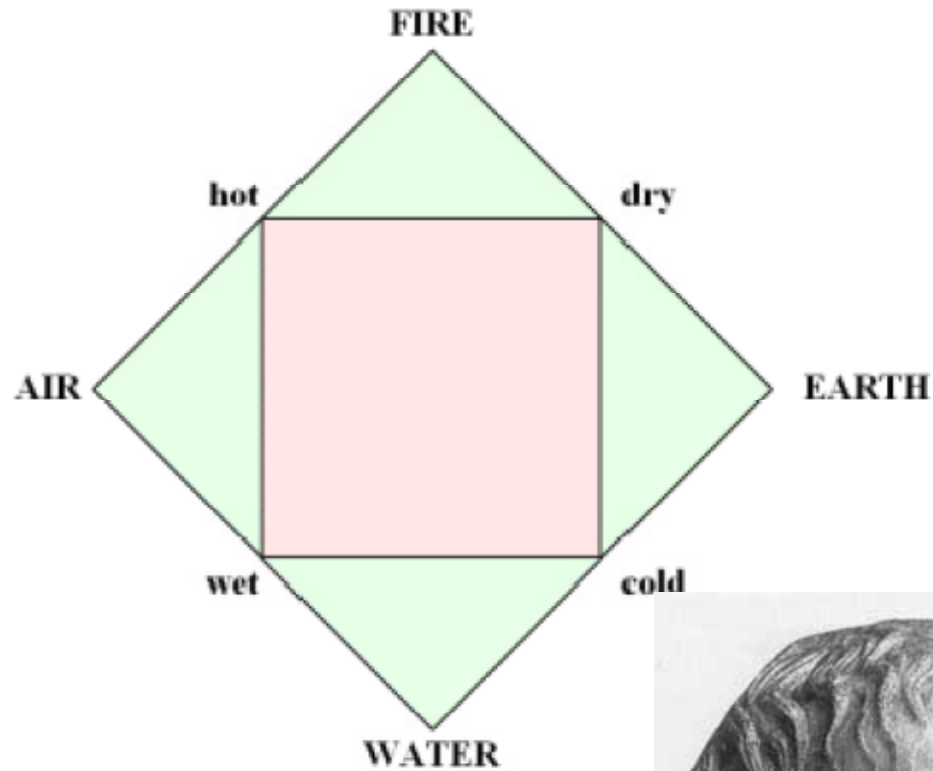


iron age
(1200 BC till present?)

Even our information age relies on our ability to manipulate materials (Si)



Ancient cultures:



Aristotle (384 BC – 322 BC)

Modern: solid, liquid, gas, combustion/chemical reactions

Hinduism (Tattva) and Buddhism (Mahābhūta)		
	Vayu/Pavan (Air/Wind)	
Ap/Jala (Water)	Akasha (Aether/Space)	Agni/Tejas (Fire)
	Prithvi/Bhumi (Earth)	
Japanese (Godai)		
	Air/Wind (風)	
Water (水)	Void/Sky/Heaven (空)	Fire (火)
	Earth (地)	
Tibetan (Bön)		
	Air	
Water	Space	Fire
	Earth	



*Alchemy workshop illustration from
Opus medico-chymicum: continens tres tractatus sive basilicas...,
Johann Mylius, Francofurtu: Hermannum a Sande, 1678*

People recognized early on the difference in properties between metals

Philosopher's stone turning common metals into gold

EARLIER THEORIES TO DESCRIBE SOLIDS

Obviously, scientists tried to deal with solids before atoms and electron

- Mechanics
- Optical properties
- Thermal conductivity
- Conductive properties

Many of these questions can be addressed without understanding the underlying nature of a material

Of great importance is the strong development of calculus and differential equations starting from Newton and Leibniz, through Euler (1707 –1783) , Gauss (1777 –1855) through the French schools (Ecole polytechnique/normale/militaire): Lagrange (1736–1813), Laplace (1749–1827), Fourier (1768–1830), Navier (1785– 1836) , Cauchy (1789 –1857), Poisson (1781–1840), etc.

Condensed-matter physics tries to connect the properties of the nuclei and electrons to the macroscopically observed quantities

Continuum mechanics

Hooke's law $\mathbf{F} = -k\mathbf{x}$ (1676)

Ut tensio, sic vis

As the extension, so the force.



Euler-Bernoulli equation:

(1750)

$$\frac{\partial^2}{\partial x^2} \left(EI \frac{\partial^2 u}{\partial x^2} \right) = w.$$

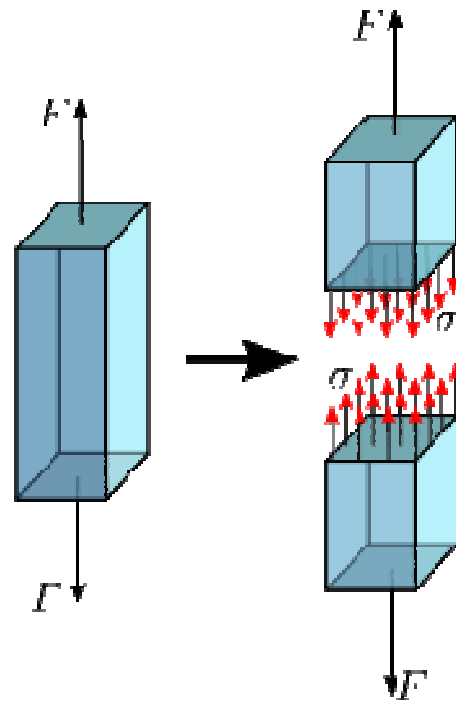
u deflection of the beam at some position x
 w is a distributed load or a force per unit length
 E is the elastic modulus
 I is the second moment of area

Cauchy stress tensor

$$\sigma_{ij} = \begin{bmatrix} \sigma_{11} & \sigma_{12} & \sigma_{13} \\ \sigma_{21} & \sigma_{22} & \sigma_{23} \\ \sigma_{31} & \sigma_{32} & \sigma_{33} \end{bmatrix} \equiv \begin{bmatrix} \sigma_{xx} & \sigma_{xy} & \sigma_{xz} \\ \sigma_{yx} & \sigma_{yy} & \sigma_{yz} \\ \sigma_{zx} & \sigma_{zy} & \sigma_{zz} \end{bmatrix} \equiv \begin{bmatrix} \sigma_x & \tau_{xy} & \tau_{xz} \\ \tau_{yx} & \sigma_y & \tau_{yz} \\ \tau_{zx} & \tau_{zy} & \sigma_z \end{bmatrix}$$



Augustin Louis Cauchy (1789 –1857)



(1822)

How are the underlying atomic properties related to the elasticity?

Interactions of radiation and matter

Reflections, color, refraction, absorption are all manifestations of interactions of radiation and matter

Euclid (~300 BC) already wrote a book on Optics



Lenses: Lippershey, Janssen, Galileo

$$\frac{\sin \theta_1}{\sin \theta_2} = \frac{v_1}{v_2} = \frac{n_2}{n_1}$$

Snell's law is a direct consequence of the electronic properties of the material



What determines the optical properties of a material: opaque, reflecting, transparent?

Late 18th and 19th century:
mechanical approach to condensed matter physics

- optical theories by Thomas Young and Jean Fresnel

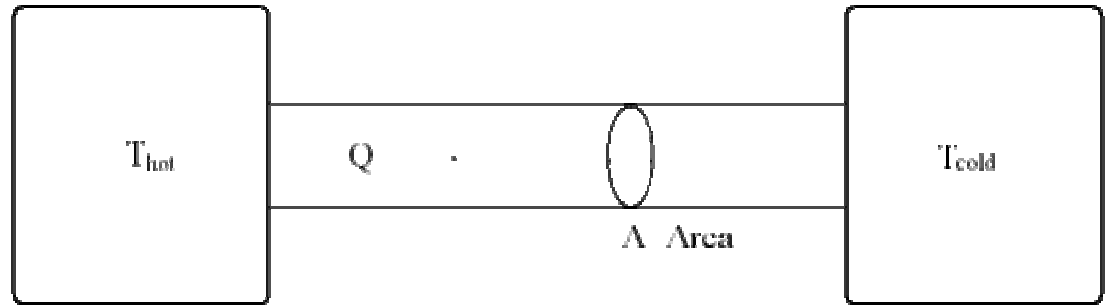


Thomas Young



- A wide variety of theories on elasticity (Navier, Cauchy)
- Theories for heat conductivity by Joseph Fourier

Thermal conductivity



Q heat [Joules]
 ΔT T_{hot} T_{cold}

Newton's law of cooling

$$\frac{dQ}{dt} = h \cdot A(T_{env} - T_0) \quad (1643-1727)$$



Fourier's law

$$q_x = -k \frac{dT}{dx} \quad (1822)$$

q is the local heat flux,
 k is the material's thermal conductivity
 dT/dx is the temperature gradient

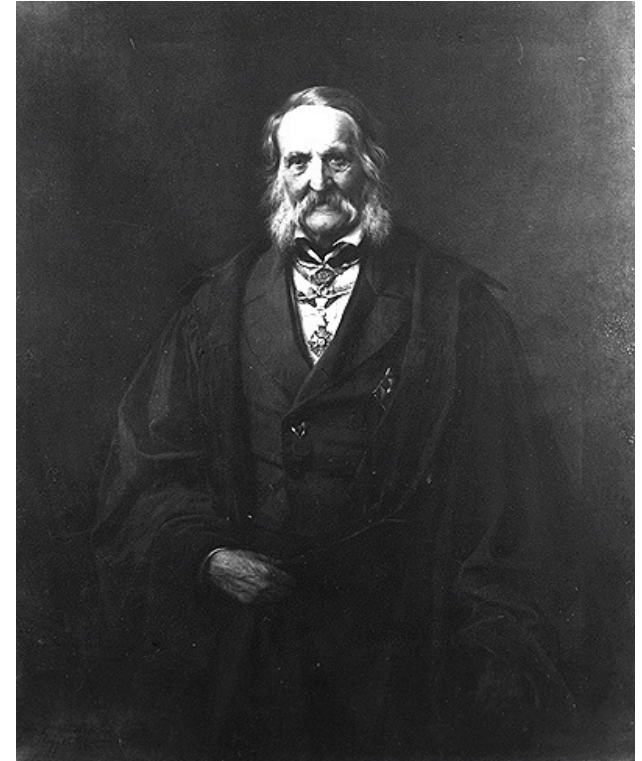


Why do materials have different thermal properties?

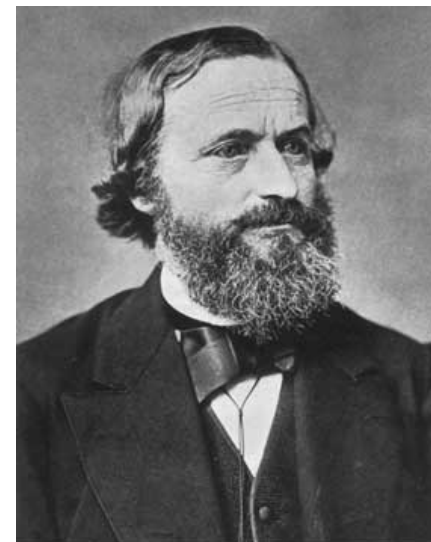
Joseph Fourier

(1768–1830)

- Interaction between light and matter, theory of birefringence by Franz Neumann



- Early theories of electrical conductivity by, among others, George Ohm and Gustav Kirchhoff



Crystal structures

- First scientific approach René-Just Haüy (1743-1822) using an atomistic picture
- Extended by Christian Samuel Weiss, introduced crystallographic axis
- Auguste Bravais: discovered the 14 space lattice types
- Woldemar Voigt classified the 230 different space groups



THE RISE (and fall) OF THE ATOMISTIC PICTURE

Leucippus (first half of 5th century BC)
Democritus (c. 460 BC – c. 370 BC)

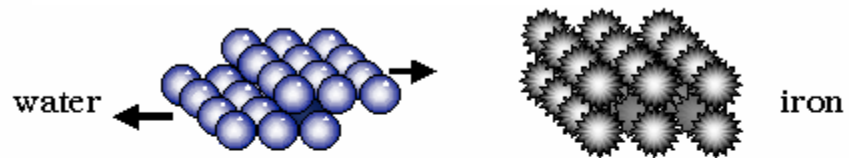


The *Atomos* Concept

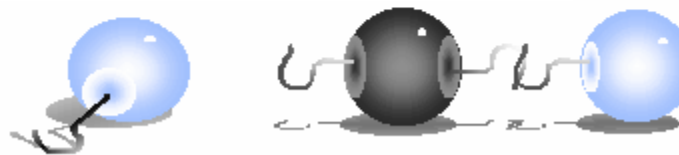
Smallest atom cannot be divided any further



Greeks: atoms determine properties



Dalton: atoms determine composition

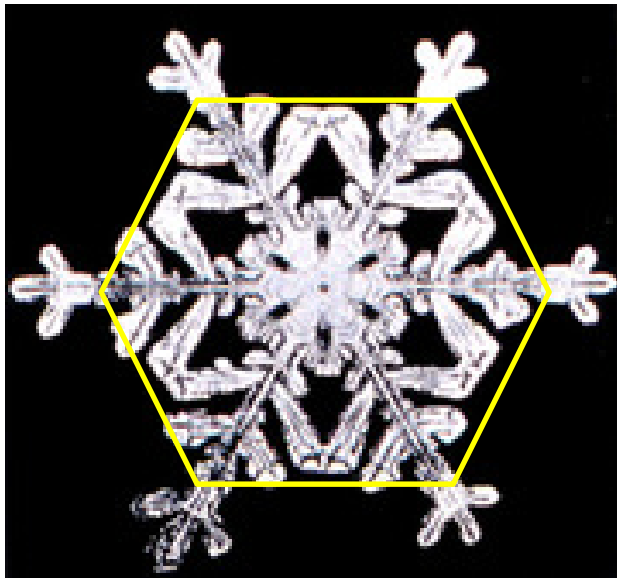


Aristotle: Horror vacui

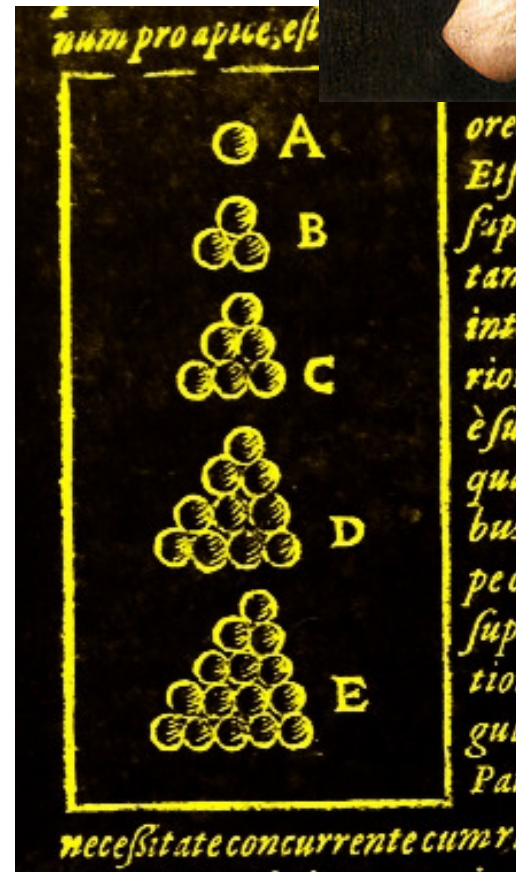
Johannes Kepler (1571 –1630)

Strena Seu de Nive Sexangula
A New Year's Gift of Hexagonal Snow

The Kepler conjecture



(1611)



Corpuscular theory



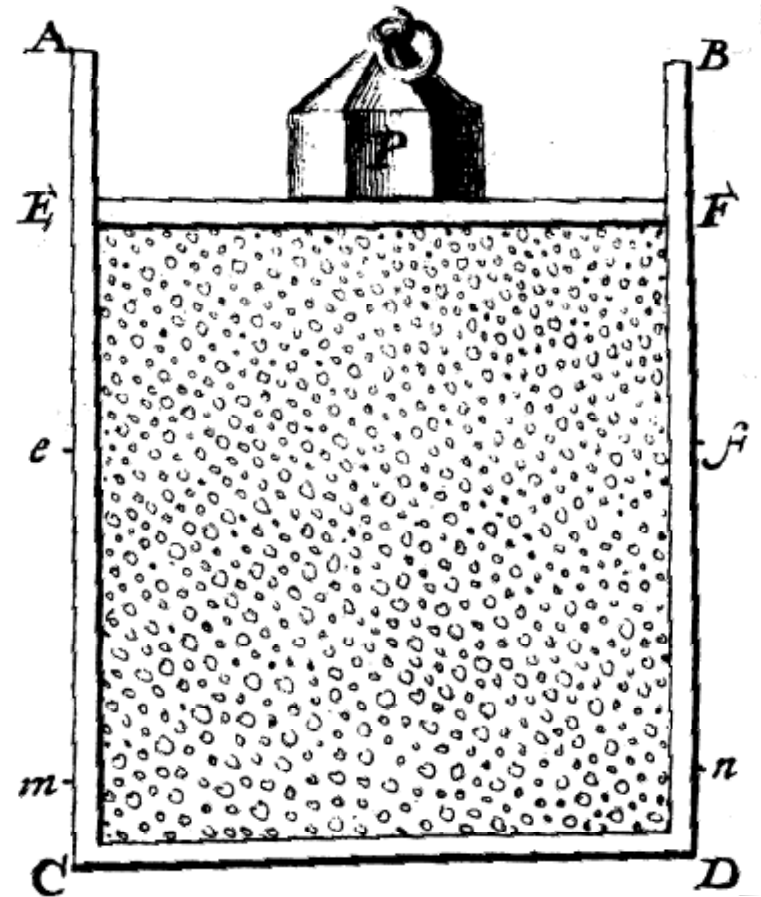
Newton



Boyle (1627 –1691)



Daniel Bernoulli (Groningen, 1700 –1782)



Let the cavity contain very minute corpuscles, which are driven hither and thither with a very rapid motion; so that these corpuscles, when they strike against the piston and sustain it by their repeated impacts, form an elastic uid which will expand of itself if the weight is removed or diminished. . . "

In physics atomistic ideas were pushed to the background in the late 18th and most of the 19th century

Unreasonable:

Not if one considers the enormous successes of continuum theories in

- Mechanics
- Thermodynamics (i.e. not statistical)
- Electricity and magnetism
- Optics

“Who needs atoms?” reigned during this period.

Not so in chemistry:

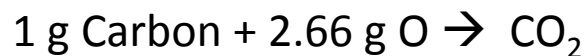
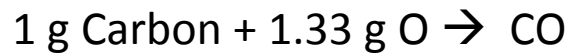
I. Law of Conservation of Mass

II. Law of Definite Proportions

Mass relationships during chemical reactions:

copper carbonate (CuCO_3) always gives 51.5% copper, 38.8% oxygen, and 9.7% carbon

III. Law of Multiple Proportions



Ratio first and second oxide 1:2

Dalton's law of partial pressures

$$P_{total} = p_1 + p_2 + \dots + p_n$$



John Dalton (1766 –1844)

ELEMENTS

	Hydrogen	1		Strontian	46
	Azote	5		Barytes	68
	Carbon	5		Iron	50
	Oxygen	7		Zinc	56
	Phosphorus	9		Copper	56
	Sulphur	13		Lead	90
	Magnesia	20		Silver	190
	Lime	24		Gold	190
	Soda	28		Platina	190
	Potash	42		Mercury	167

ELEMENTS Plate 1

Simple

1 2 3 4 5 6 7 8

9 10 11 12 13 14 15 16

17 18 19 20

Binary

21 22 23 24 25

Ternary

26 27 28 29

Quaternary

30 31 32 33

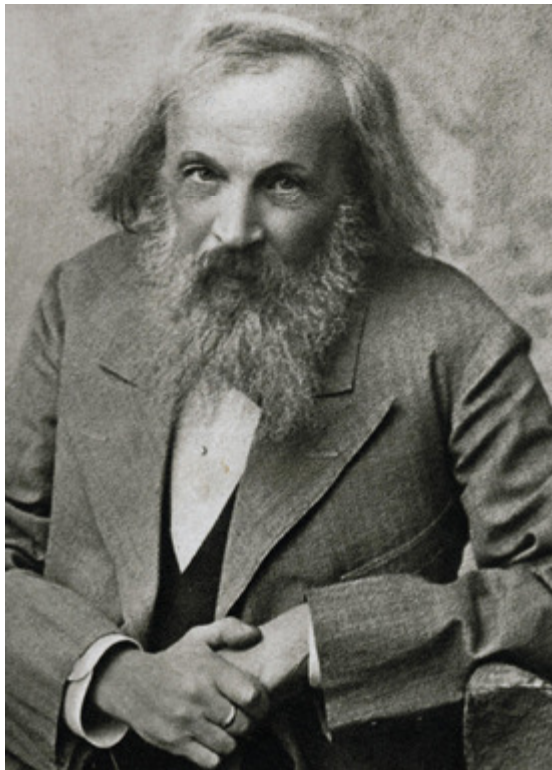
Quinary & Sextenary

34 35

Septenary

36 37

Culmination in Mendeleev's Periodic table



Dmitri Mendeleev (1834 –1907),

THE PERIODICITY OF THE ELEMENTS

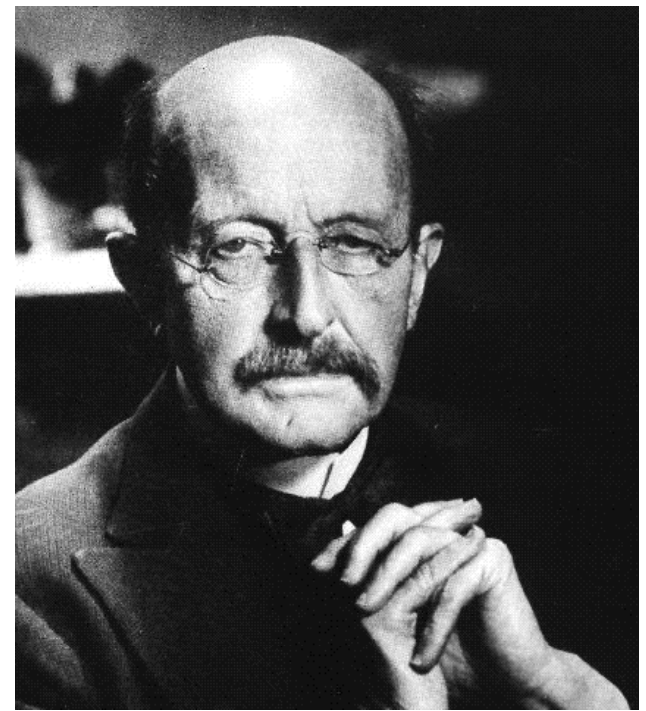
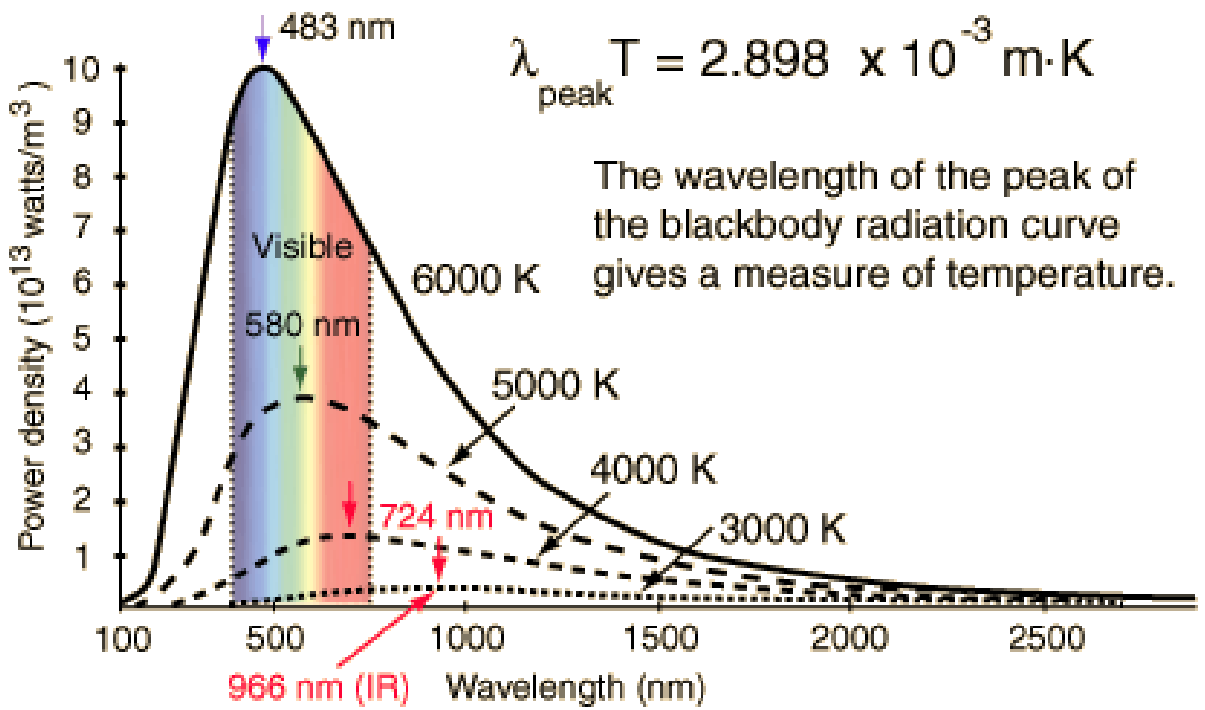
The Elements	Their Properties in the Free State				The Composition of the Hydrogen and Organo-metallic Compounds	Symbols and Atomic Weights	The Composition of the Saline Oxides	The Properties of the Saline Oxides	Small Periods or Series	
	t	a	d	A						
Hydrogen	<-200°	—	<0.03>	20	RH ₄ or R(CH ₃) _n	B	A	R ₂ O ₃	d(2A + n16) v	[11]
Lithium	180°	—	0.50	12	[5]	[6]	[7]	[8]	[9]	[10]
Beryllium	(1900°)	—	1.64	9.5	3	—	—	—	—	—
Boron	(2500°)	—	2.5	10.8	4	—	—	—	—	—
Carbon	>3500°	—	<2.0>	12	3	—	—	—	—	—
Nitrogen	<300°	—	<0.7>	14	3	—	—	—	—	—
Oxygen	<-200°	—	<1.0>	16	2	—	—	—	—	—
Fluorine	—	—	—	19	1	—	—	—	—	—
Sodium	90°	0.71	0.98	23	1	Na	23	1†	Na ₂ O	26
Magnesium	500°	0.27	1.74	24	2	Mg	24	—2†	—	—
Aluminium	600°	0.23	2.6	27	3	Al	27	—3	Al ₂ O ₃	40
Silicon	(1200°)	0.08	2.3	28	4	Si	28	—3 4	—	—
Phosphorus	44°	1.28	2.2	31	3	P	31	1 — 3* 4* 5*	—	—
Sulphur	114°	0.07	2.07	32	2	S	32	— 2 — 4* 5* 6*	—	—
Chlorine	-75°	—	1.3	35.5	1	Cl	35.5	1 — 3 — 5* — 7*	—	—
Potassium	58°	0.84	0.87	39	1	K	39	1†	—	—
Calcium	(800°)	—	1.6	40	2	Ca	40	—2†	CaO	59
Scandium	—	—	(2.5)	44	3	Sc	44	—3†	—	—
Titanium	(2500°)	—	(5.1)	48	4	Ti	48	—3 4	—	—
Vanadium	(3000°)	—	5.5	51	5	V	51	—2 3 4 5	—	—
Chromium	(3000°)	—	5.5	52	6	Cr	52	—2 3 — 6*	—	—
Manganese	(1500°)	—	7.5	55	7	Mn	55	—2† 3 4 — 6* 7*	—	—
Iron	1400°	0.12	7.8	56	6	Fe	56	—2† 3 — 6*	—	—
Cobalt	(1400°)	0.13	8.6	59	5	Co	59	—2† 3 — 6*	—	—
Nickel	1350°	0.17	8.7	58.5	5	Ni	58.5	—2† 3 — 6*	—	—
Copper	1054°	0.29	8.8	63.5	2	Cu	63.5	1† 2†	Cu ₂ O	59
Zinc	(1200°)	—	7.1	65.4	2	Zn	65.4	—2†	—	—
Gallium	30°	—	5.90	70	3	Ga	70	—2 3	Ga ₂ O ₃	59
Germanium	900°	—	5.47	72	4	Ge	72	—2 3 4	—	—
Arsenic	500°	0.06	5.7	75	3	As	75	—2 3 4 5*	—	—
Selenium	217°	—	4.8	79	2	Se	79	—2 3 4 5*	—	—
Bromine	-7°	—	3.1	80	1	Br	80	1 — 3 — 5* — 7*	—	—
Iodine	30°	—	1.5	85	1	I	85	1†	—	—
Barium	(600°)	—	2.5	87	2	Ba	87	—2†	—	—
Strontium	—	—	(2.4)	88	2	Sr	88	—2†	—	—
Yttrium	(1500°)	—	4.1	89	3	Y	89	—3†	—	—
Zirconium	—	—	7.1	91	4	Zr	91	—2 3 4	—	—
Niobium	—	—	8.6	93	5	Nb	93	—2 3 4 5*	—	—
Molybdenum	—	—	—	96	6	Mo	96	—2 3 4 — 6*	—	—
Ruthenium	(2000°)	0.10	12.2	101	8	Ru	101	—2 3 4 — 6 — 8	—	—
Rhodium	(1900°)	0.08	12.1	103	7	Rh	103	—2 3 4 — 6	—	—
Palladium	1500°	0.12	11.4	106	6	Pd	106	1† 2 — 4	—	—
Silver	950°	0.19	10.5	108	1	Ag	108	1†	Ag ₂ O	75
Cadmium	320°	0.51	8.6	112	2	Cd	112	—2†	—	—
Iodine	176°	0.46	7.4	115	3	In	115	—2 3	In ₂ O ₃	81.5
Tin	230°	0.23	7.2	118	4	Sn	118	—2 — 4	—	—
Antimony	432°	0.12	6.7	120	3	Sb	120	—3 4 5	—	—
Tellurium	452°	0.17	6.4	127	2	Te	127	—3 4 5 6*	—	—
Iodine	114°	—	4.9	127	1	I	127	1 — 3 — 5* — 7*	—	—
Cesium	27°	—	1.88	133	1	Cs	133	1†	—	—
Barium	—	—	3.75	137	2	Ba	137	—2†	—	—
Lanthanum	(600°)	—	6.1	139	3	La	139	—3†	—	—
Cerium	(700°)	—	6.6	140	4	Ce	140	—3 4	—	—
Didymium	(800°)	—	6.5	142	5	Di	142	—3 — 5	—	—
Ytterbium	—	—	(6.9)	173	14	Yb	173	—3	—	—
Tantalum	—	—	10.4	182	5	Ta	182	—4 — 6	—	—
Tungsten	(1500°)	—	19.1	184	6	W	184	—4 — 6	—	—
Osmium	(2500°)	0.07	22.5	191	8	Os	191	—3 4 — 6 — 8	—	—
Iridium	2000°	0.07	22.4	193	7	Ir	193	—3 4 — 6	—	—
Platinum	1772°	0.05	21.5	195	6	Pt	195	—2 — 4	—	—
Gold	1045°	0.14	19.3	197	3	Au	197	1 — 3	Au ₂ O (125)	111
Mercury	-39°	—	13.6	200	2	Hg	200	1† 2†	—	—
Thallium	284°	0.51	11.8	204	3	Tl	204	1† — 3	—	—
Lead	326°	0.29	11.3	206	4	Pb	206	—2† — 4	Tl ₂ O ₃ (97)	111
Bismuth	268°	0.14	9.8	208	3	Bi	208	—3 — 5	—	—
Thorium	—	—	11.1	232	4	Th	232	—4	—	—
Uranium	(800°)	—	18.7	240	6	U	240	—4 — 6	—	—

REVOLUTION IN PHYSICS

$$E = h\nu$$

frequency of radiation, sometimes written as f giving expression $E = hf$.
 Quantum energy of a photon.

$h = \text{Planck's constant} = 6.626 \times 10^{-34} \text{ Joule}\cdot\text{sec} = 4.136 \times 10^{-15} \text{ eV}\cdot\text{s}$



Energy per unit volume per unit frequency

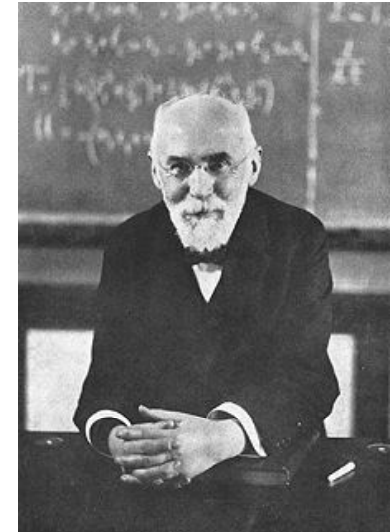
$$S_\nu = \frac{8\pi h}{c^3} \frac{\nu^3}{e^{h\nu/kT} - 1}$$

Energy per unit volume per unit wavelength

$$S_\lambda = \frac{8\pi hc}{\lambda^5} \frac{1}{e^{hc/\lambda kT} - 1}$$

Classical electron theory

- discovery of electron J. J. Thomson, Lorentz



- Drude model

Treats electrons as a gas following Boltzmann statistics as opposed to Fermi-Dirac statistics. Quantities of by several orders of magnitude (gets a lucky break with Wiedemann-Franz law)

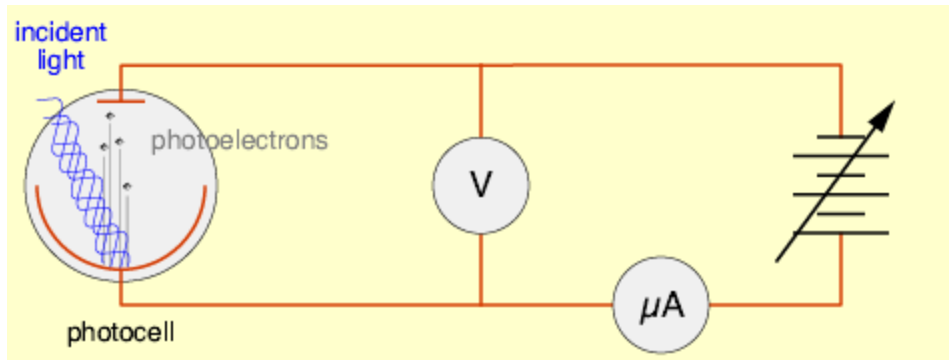
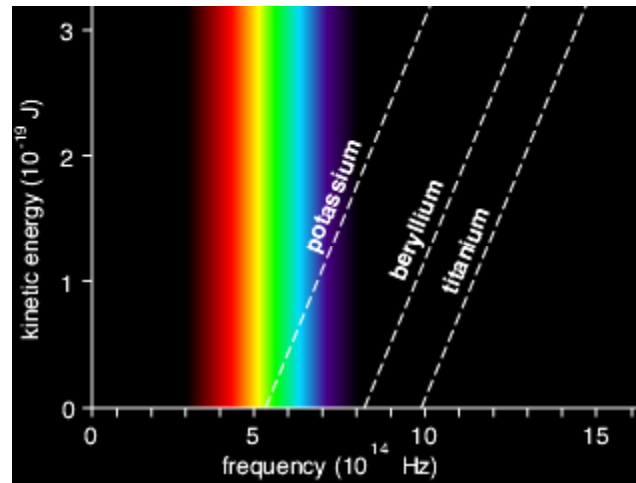
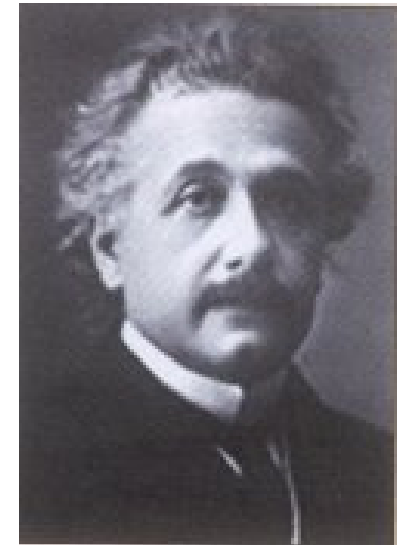


$$E = h\nu$$

frequency of radiation, sometimes written as f giving expression $E = hf$.

Quantum energy of a photon.

h = Planck's constant = 6.626×10^{-34} Joule·sec = 4.136×10^{-15} eV·s



MODERN SOLID STATE PHYSICS

Solid state physics based on atoms
generally based on quantum-mechanics
(although sometimes classical mechanics)

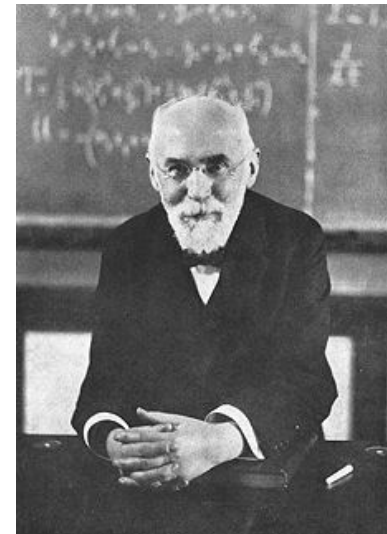
Classical electron theory

- Drude model (1900)

Treats electrons as a gas following Boltzmann statistics as opposed to Fermi-Dirac statistics. Quantities of by several orders of magnitude (gets a lucky break with Wiedemann-Franz law)



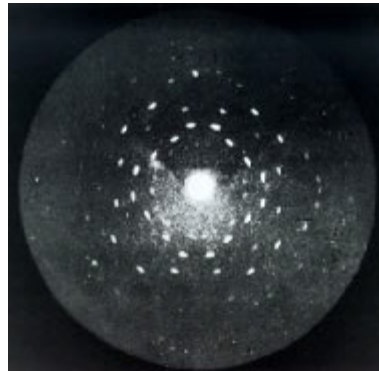
Paul Drude (1863 –1906)



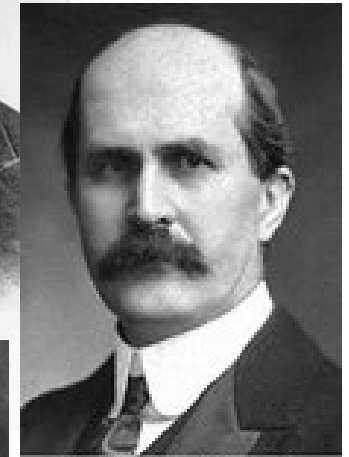
X-ray diffraction

1895: discovery of X-rays by Wilhelm Röntgen (Nobel 1901)

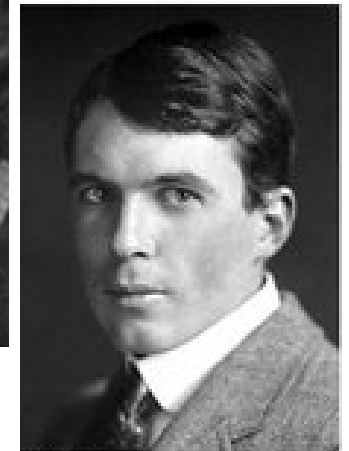
1912: discovery of X-ray diffraction by Max von Laue, Nobel 1914
(other contributors Ewald, Sommerfeld)



1913: interpretation by William and Lawrence Bragg, Nobel 1915



Sir William Henry
Bragg
(1862 - 1942)



William Lawrence
Bragg
(1890 - 1971)

First applications of quantum mechanics

Specific heat of solids: The change in internal energy with respect to temperature

experiments by Nernst

Calculations by Einstein and Debye



W. Nernst

Walther Nernst (1864 –1941)



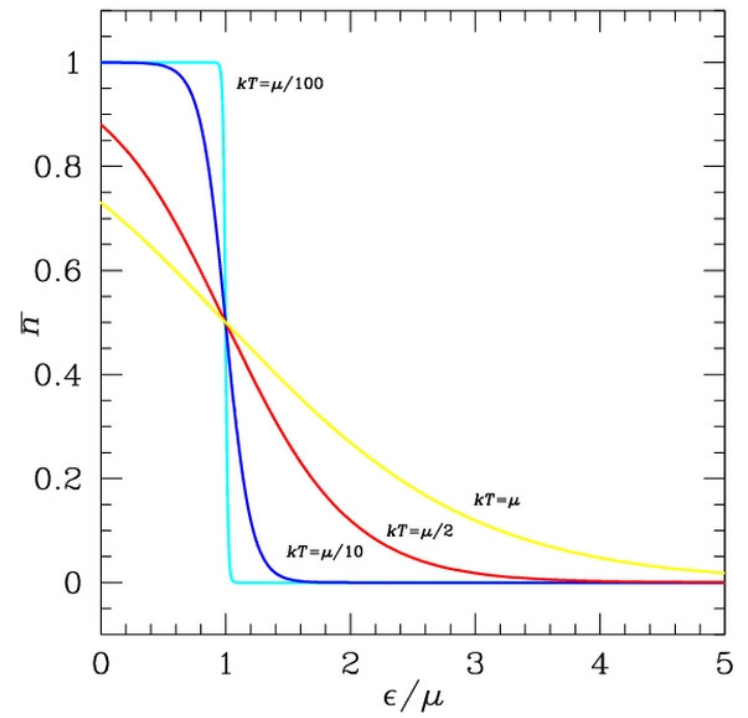
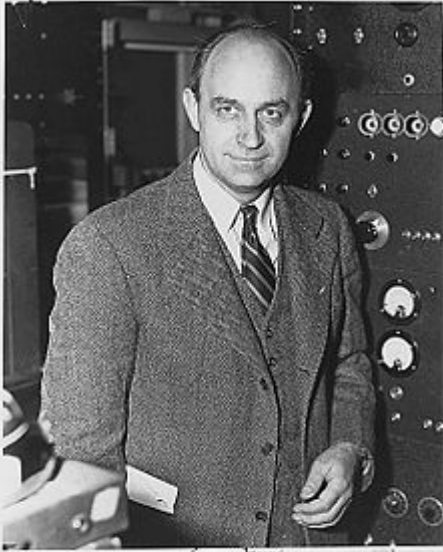
Einstein
 $E(k) = \omega_0$



Peter Debye (1884 –1966)
Nobel Chem 1936

$E(k) \sim k$

Fermi-Dirac statistics



Sommerfeld theory including Fermi-Dirac statistics

Specific heat much smaller since very few electrons participate in the conduction

Solves dilemma of Drude-Lorentz theory

Completely ignores the presence of ions!

Still fails to describe many properties...

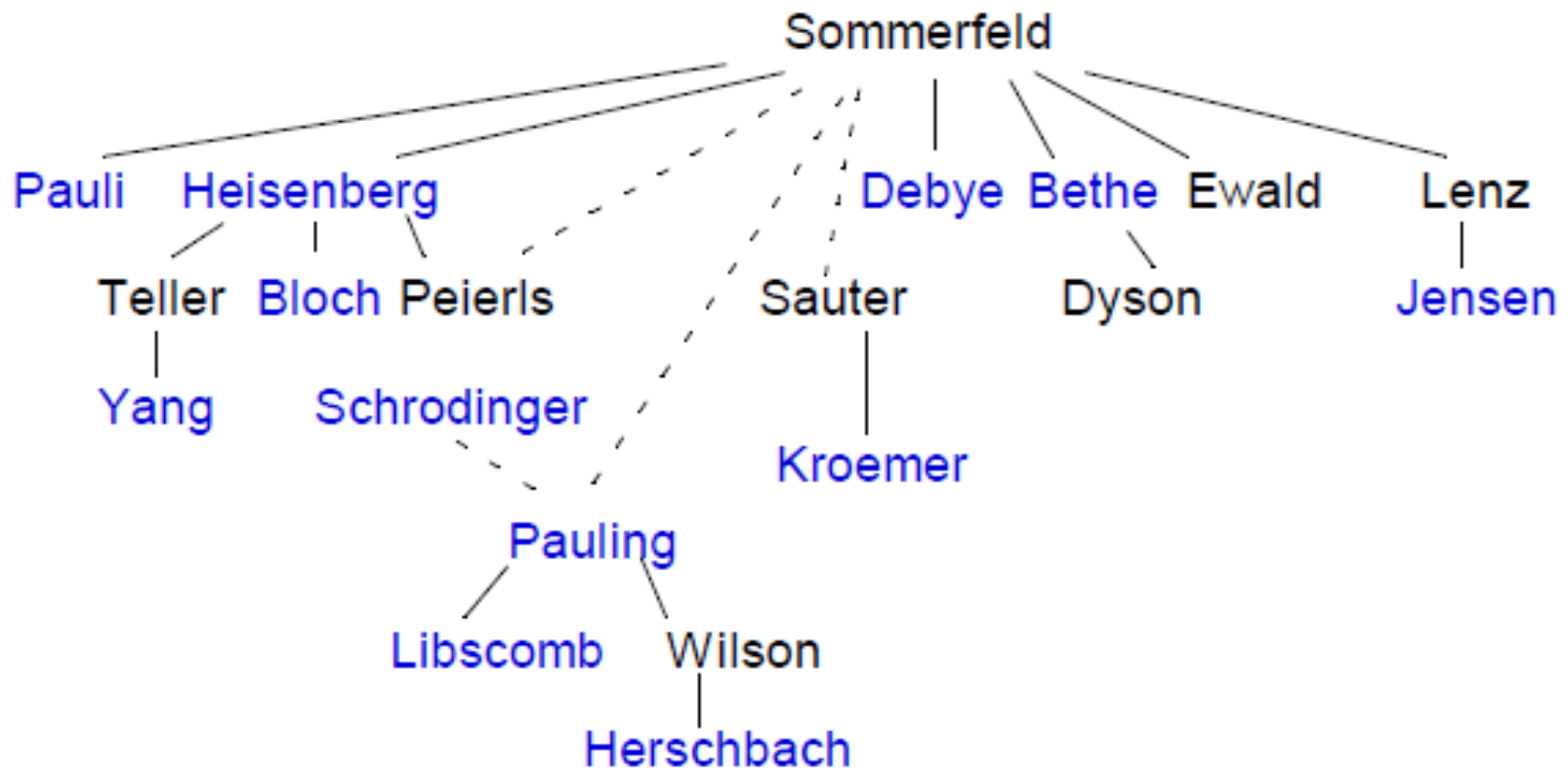


Fermi and Sommerfeld



Sommerfeld and Pauli

In addition, Sommerfeld was a star in producing world-class scientists (a selection)



Albert Einstein told Sommerfeld: “What I especially admire about you is that you have, as it were, pounded out of the soil such a large number of young talents.”

Bloch's theorem

Inclusion of the ions in the theory of metals

Inclusion of translational symmetry is essential

Bloch and Heisenberg



$$\psi_{\mathbf{k}}(\mathbf{r}) = \frac{1}{\sqrt{V}} e^{i\mathbf{k} \cdot \mathbf{r}},$$

Free electrons



$$\psi_{n\mathbf{k}}(\mathbf{r}) = e^{i\mathbf{k} \cdot \mathbf{r}} u_{n\mathbf{k}}(\mathbf{r}),$$

$$u_{n\mathbf{k}}(\mathbf{r} + \mathbf{R}) = u_{n\mathbf{k}}(\mathbf{r})$$

Bloch electrons
(Bloch's theorem)

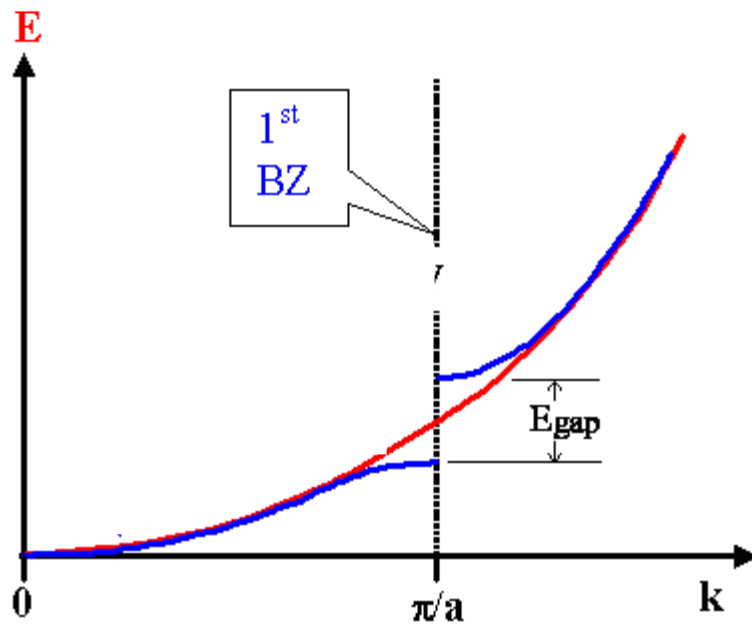
Band gaps and Brillouin zones



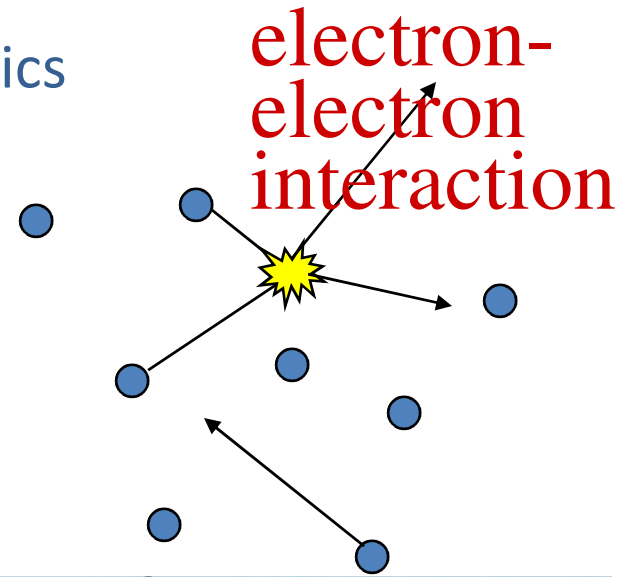
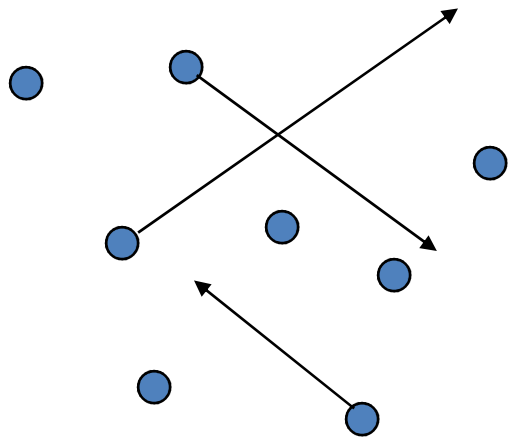
Sir Rudolf Peierls, (1907–1995),



Léon Brillouin (1889 –1969)



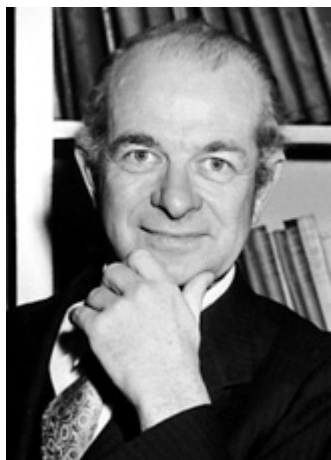
Independent-particle vs. many-body physics



This is a fundamental problem in physics that is not well understood!

There are very extensive codes based on the independent-particle approximation

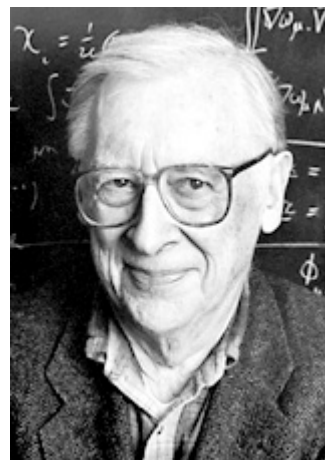
- Density functional theory
- Local Density Approximation
- Molecular orbital theory
- Quantum chemistry



Linus Pauling (1901-1994)
Nobel chem 1954



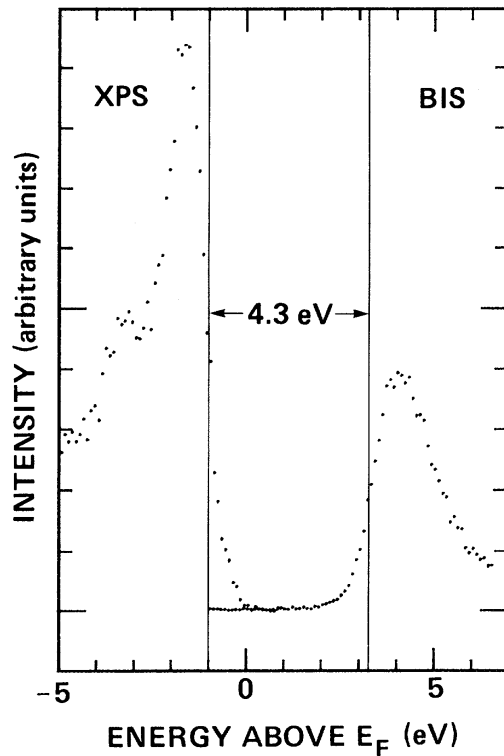
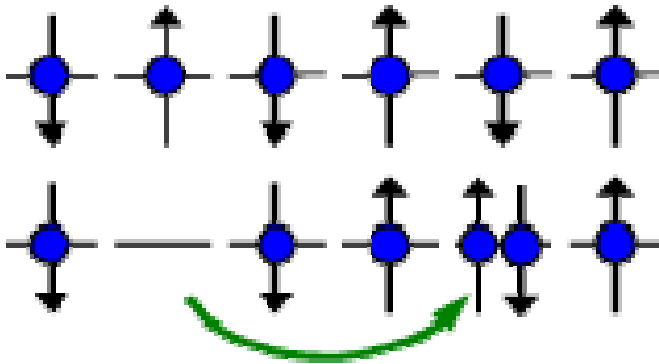
Walter Kohn (1923) John Pople (1925-2004)
Nobel chem 1998



However, often we have to deal with many-body effects

- Effects where electron-electron interactions are important

Mott-insulator state



Philip W. Anderson (1923)

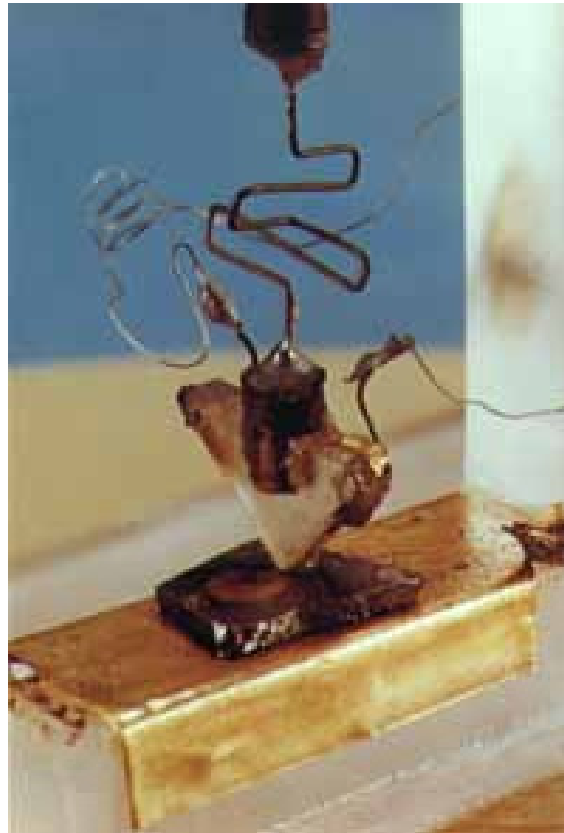
John van Vleck (1899-1980)

Sir Nevill Mott (1905-1996)

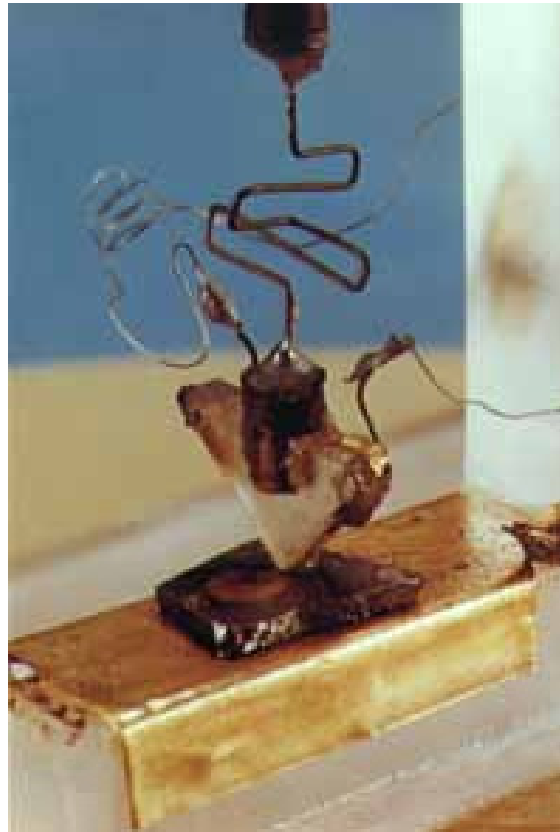
Nobel 1977

Why is NiO an insulator and not a metal?

WHAT IS THIS?

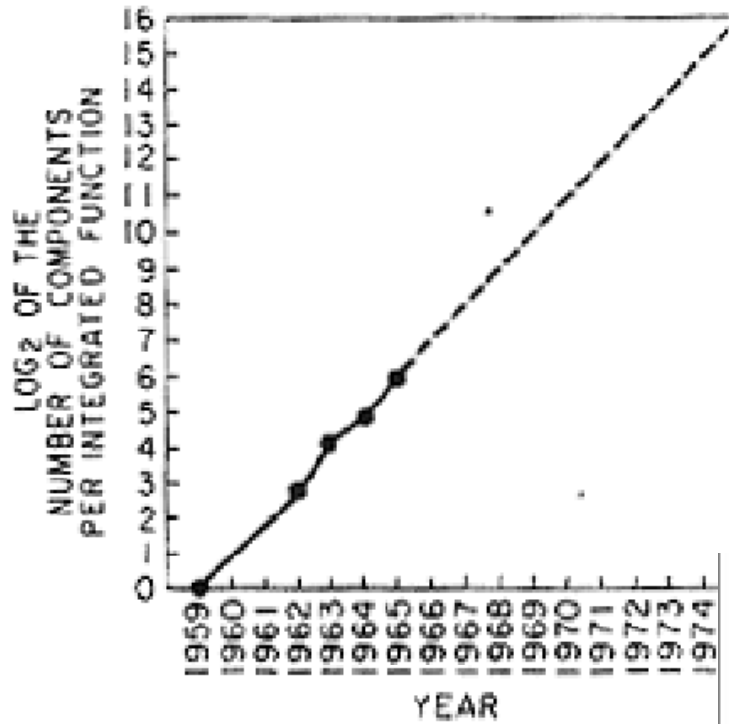


The first transistor!
Bell-Labs (1947)

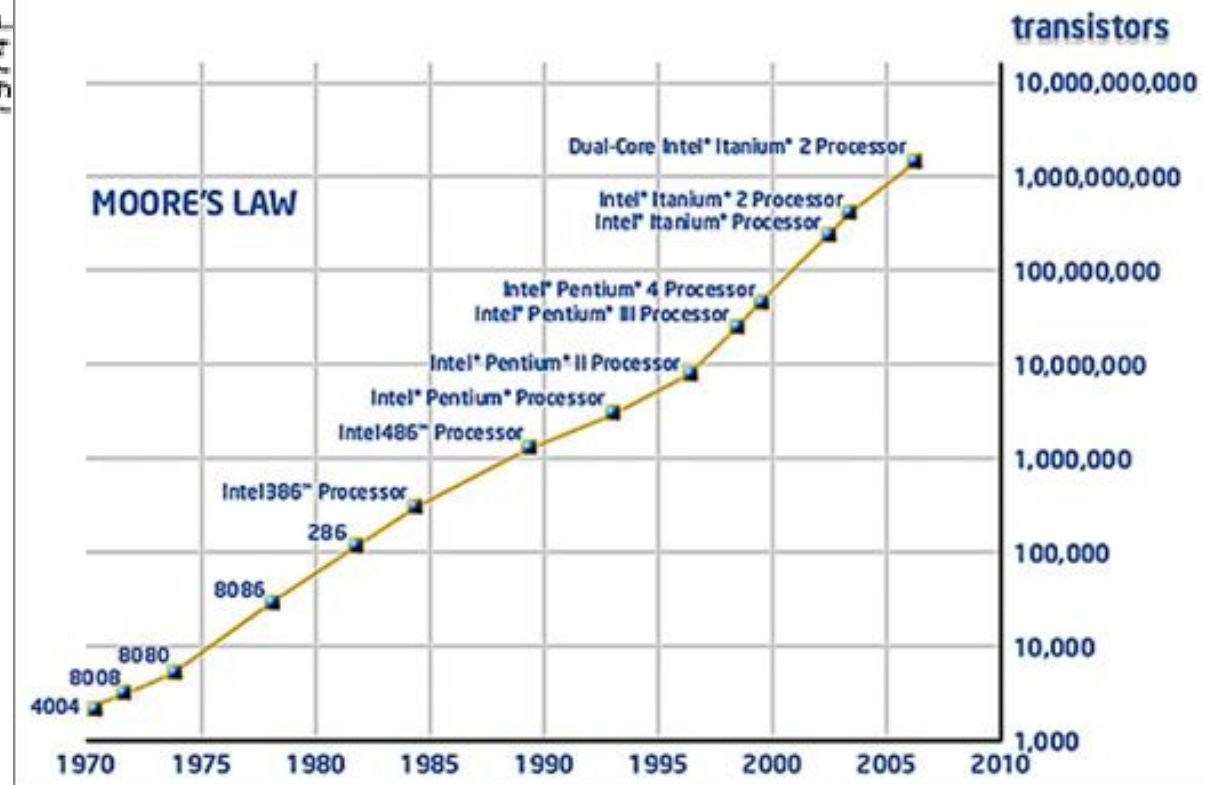


**John Bardeen, Walter Brattain,
and William Shockley working
on the first transistor**

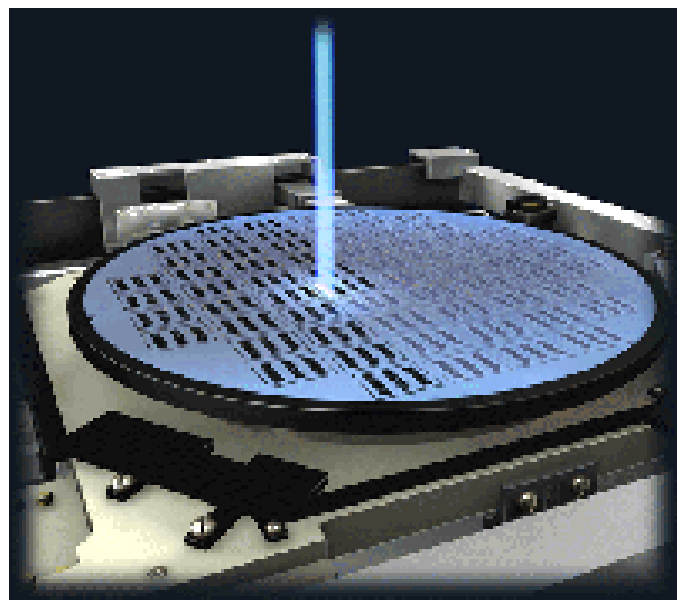
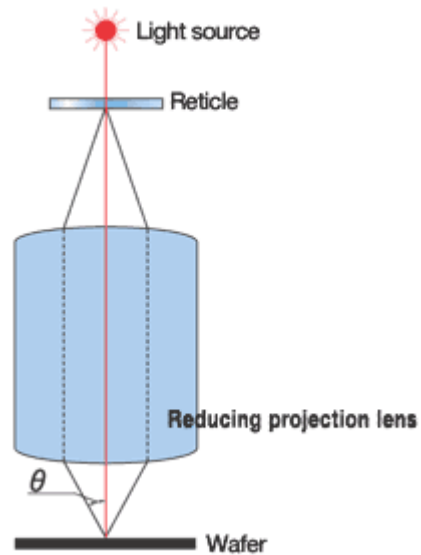
Nobelprize 1956



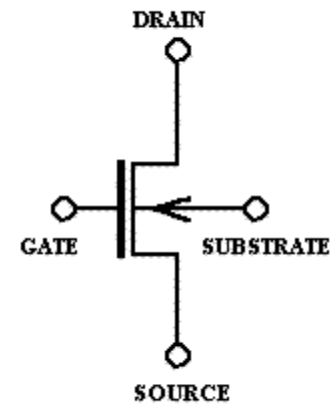
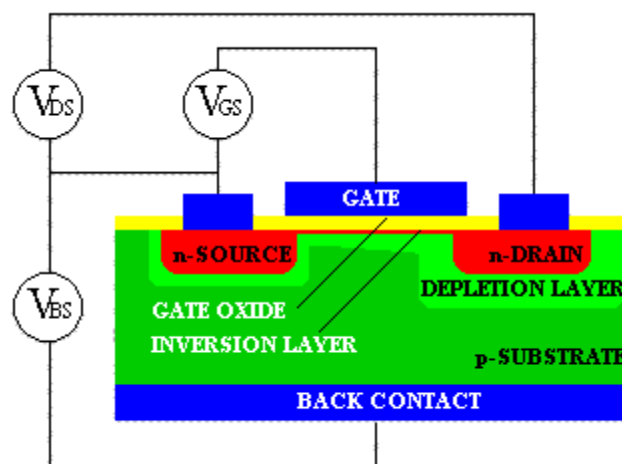
Moore's Law



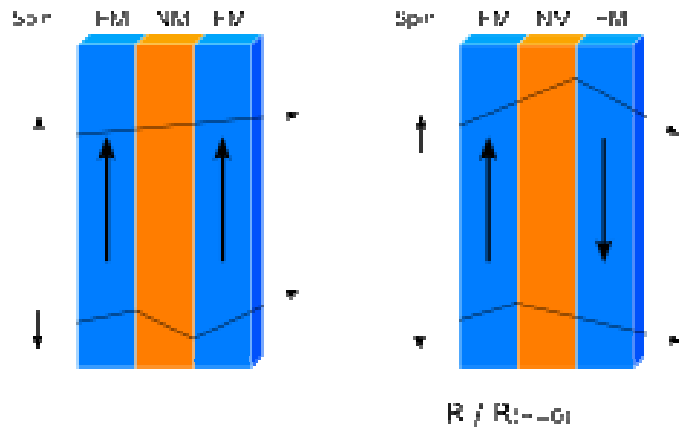
Optical lithography



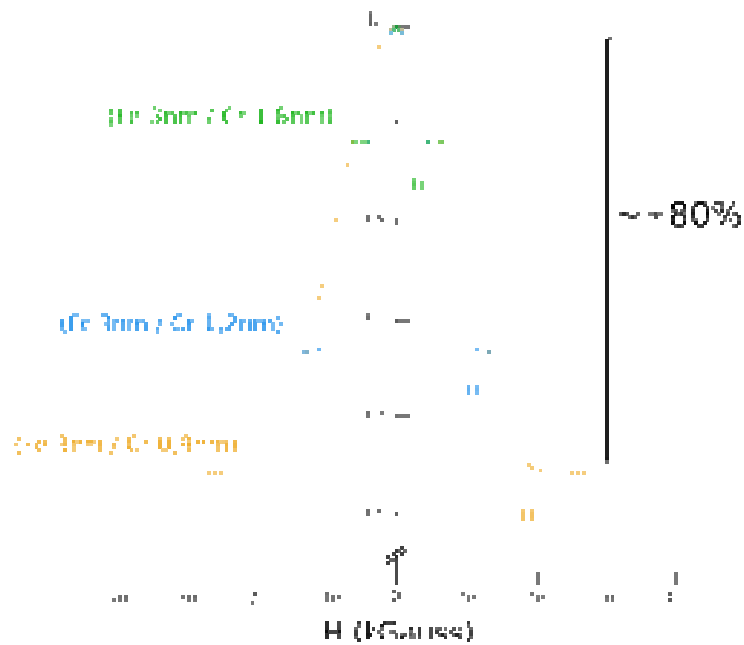
$$RES = k_1 \lambda / NA$$



Other device technology: giant magnetoresistance



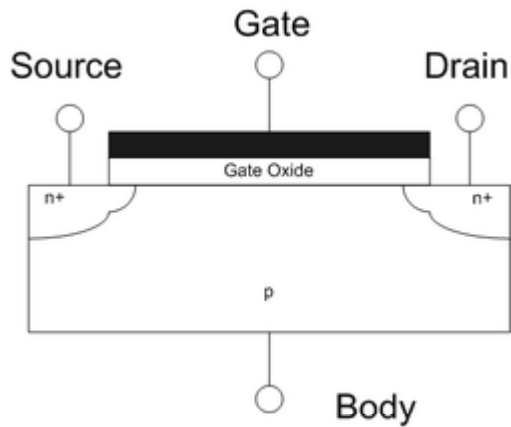
Albert Fert (1938) Peter Grünberg (1939)
 Nobel 2007



Read heads of hard drives

More exotic phenomena (only at low temperatures)

Quantized resistance, quantum Hall effect, fractional quantum Hall effect



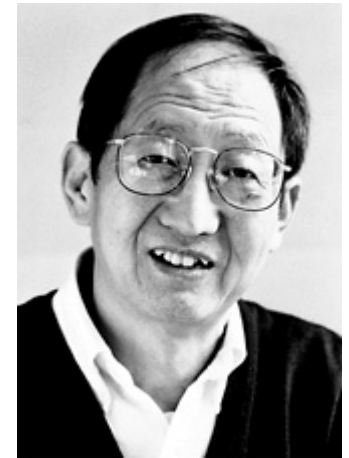
Klaus von Klitzing (1943)



Robert B. Laughlin (1950)

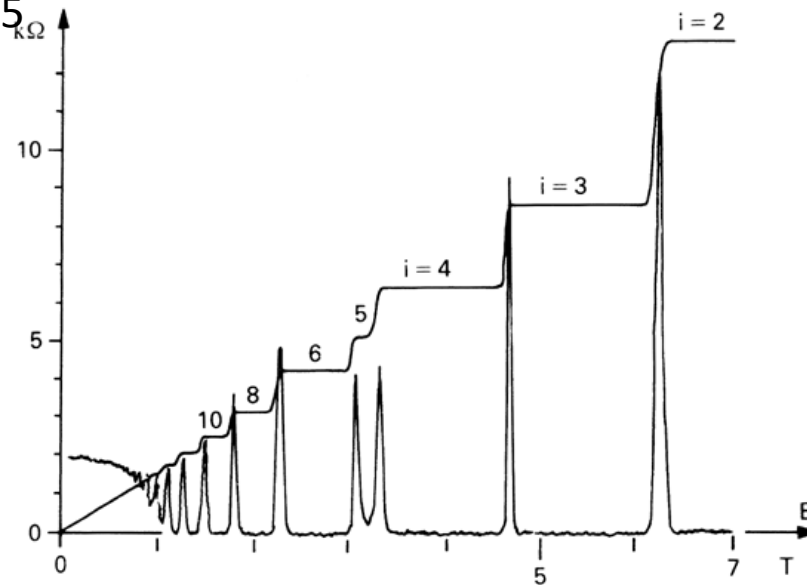
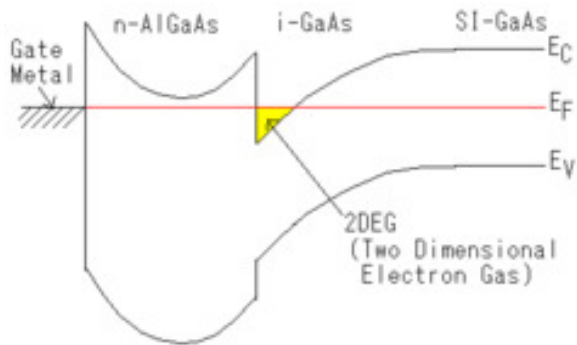


Horst L. Störmer (1949)



Daniel C. Tsui (1939)

Nobel 1985



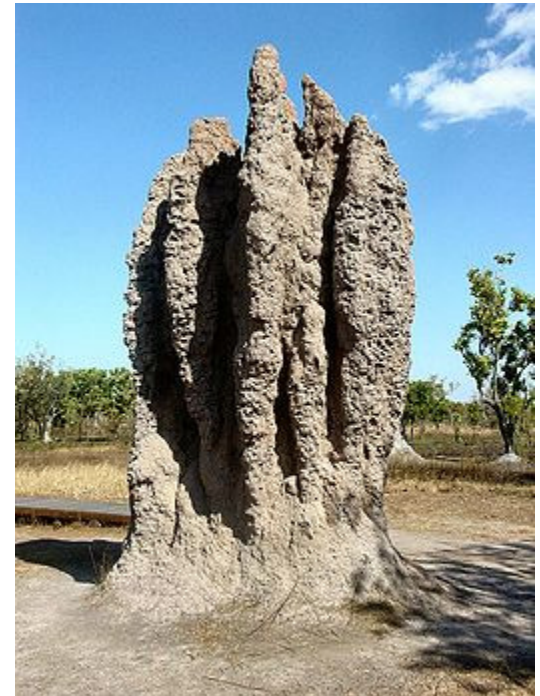
Nobel 1998

EMERGENT PHENOMENA

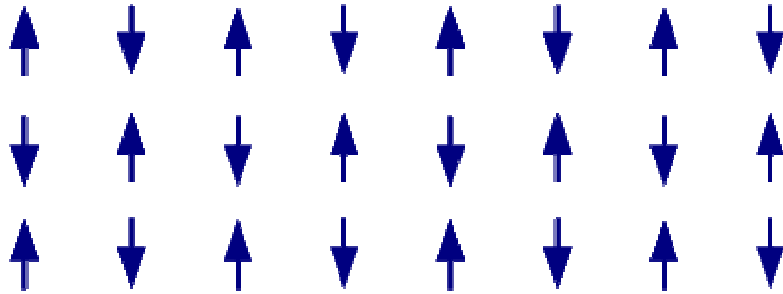
The whole is greater than the sum of its parts.

Sometimes when you put things together new order appears

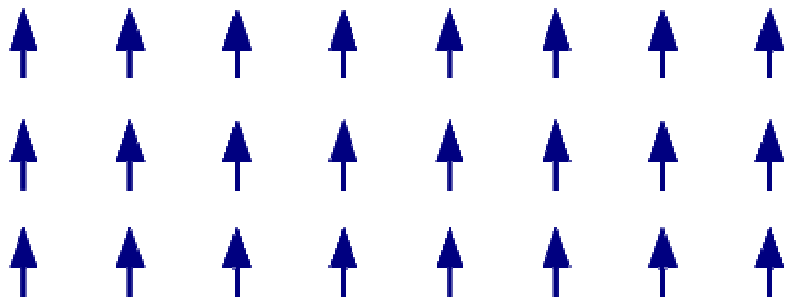
- Magnetism
- Superconductivity
- Superfluidity



An example: Antiferromagnetism



(as opposed to ferromagnetism



)



Louis Néel (1904 –2000)

Nobel 1970

Experimental developments:

Low-temperature physics:

Kamerlingh-Onnes

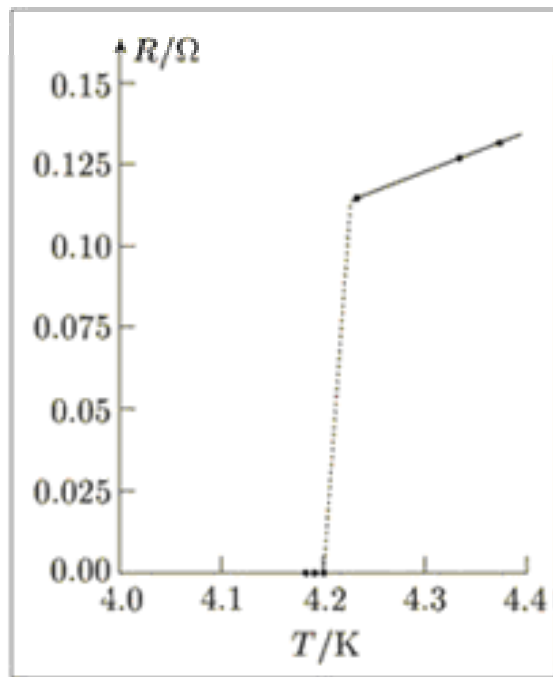
1908: Liquefaction of Helium

1911: Discovery of superconductivity

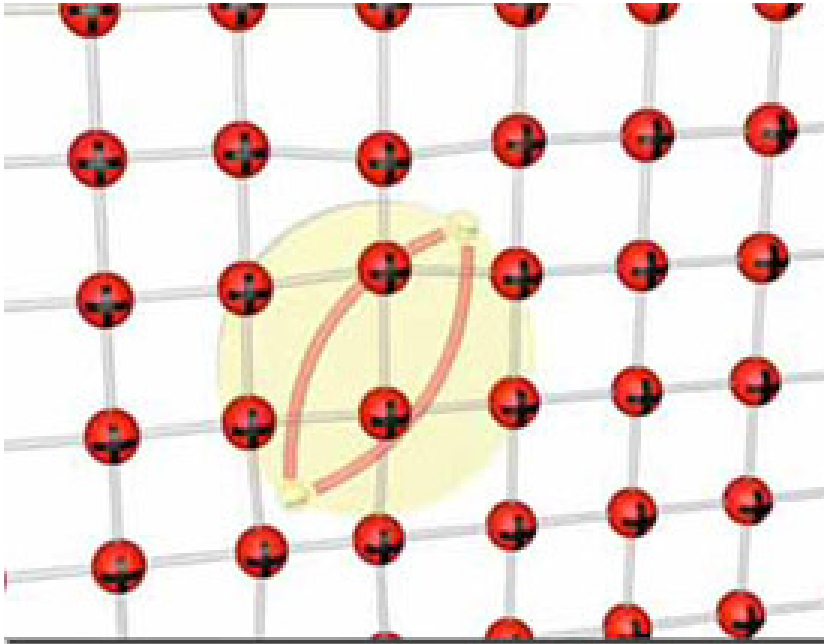


Heike Kamerlingh Onnes (1853 –1926)

Nobel 1913



Only explained in 1957:



Cooper pairs



John Bardeen (1908 –1991)

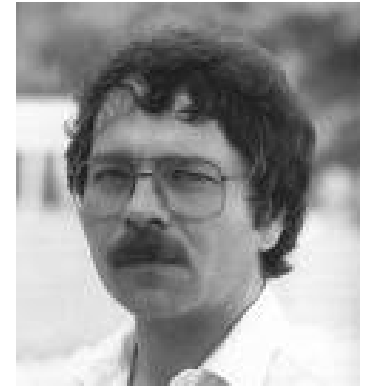
Leon Cooper (1930)

John Robert Schrieffer (1931)

Nobel 1972

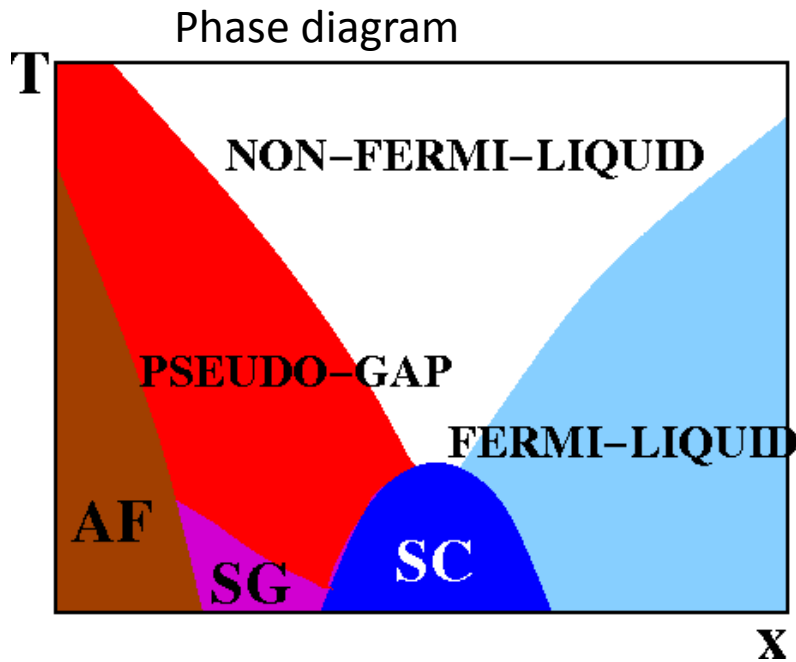
High temperature superconductivity

Johannes Georg Bednorz (1950)



Karl Alexander Müller (1927)

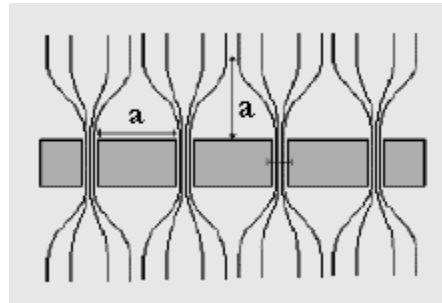
Nobel 1987



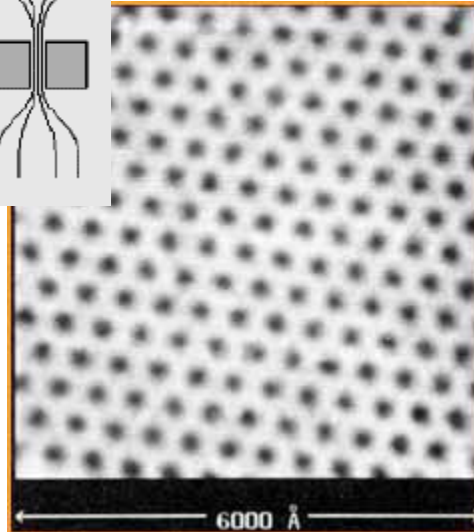
Still under debate....



Infinitely fascinating, apparently, superconductors

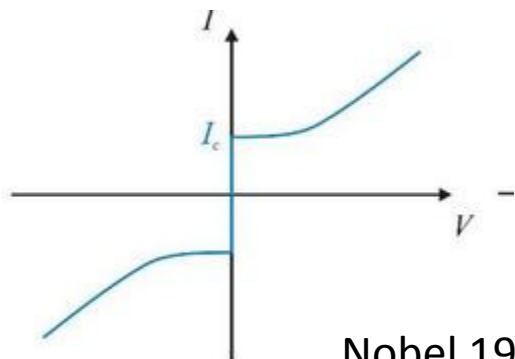


vortices



Alexei Abrikosov (1928-)
Nobel prize 2004

Plus all kinds of interesting junction effects



Nobel 1973



Brian Josephson (1940)

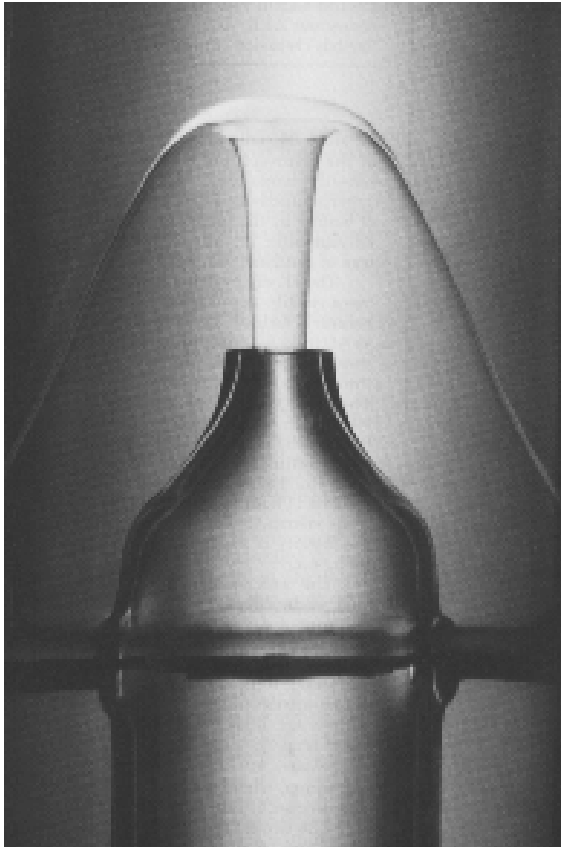


Ivar Giaever (1929)



Leona Esaki (1925)

A strongly related phenomena (especially theoretically) is superfluidity



Helium II



Lev Landau (1908 –1968) Nobel 1962



Sir Anthony Leggett, KBE, FRS (1938-)
Nobel 2004