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Time and Diurnal Motion

A. Geography: mapping the earth

B. Equatorial Coordinates

C. Local Horizon System

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A. Geography: mapping the earth

Geometry: measure the earth!

- 1) The earth is a ball
- 2) Geographic Coordinates
- 3) Definition of Time

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1a. The Earth Is Flat

Thales' Theorems
by Kathleen Norton

1) A circle is bisected by its diameter (d)

2) Base angles of an isosceles triangle are equal.

3) $\angle ZEC = \angle ABD$
 $\angle ZEA = \angle DBC$
Opposite angles are equal.

4) $BC = (BD + AD) \frac{AB}{AC}$

Corresponding angles in a triangle have proportional sides.

Thales of Miletos (624-545BC)

"the father of Greek science and astronomy"

the earth is a disk floating on the world ocean, surrounded by the rotating celestial sphere carrying the stars and other heavenly objects.

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1b. The Earth Is Round

- Pythagoras of Samos (569-475 BC)
- "the first pure mathematician"
- lunar eclipses imply shadow of earth is round
- Bottoms of ships vanish first as ship goes to sea

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1c. Aristotle (384-322 BC)

- notes change in altitude of north star as travel south
- Estimates circumference of earth is 74,000 km (true value is 40,000 km)

1d. Erastotenes (275-194 BC) 7

- 236 BC Director Library of Alexandria
- 240 BC Was told that on midsummer day (June 21) in the town of Syene in southern Egypt (today Aswan, near a huge dam on the river Nile) the noontime Sun was reflected in a deep well, meaning that it was right overhead, at zenith.
- But at Alexandria, on same day it was NOT overhead!

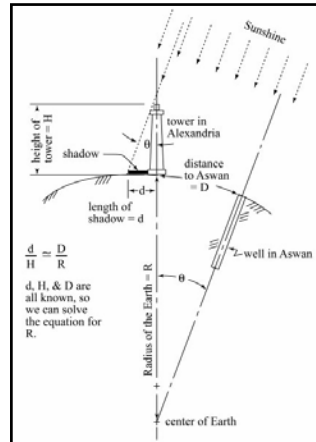


Erastotenes 8

In Alexandria sunlight was at angle of 7.2 degrees.

Hired soldier to march out distance between cities: 5040 Stadia

Calculates circumference is 250,000 stades [within 1% true value]



2. Cartography 9

in Greek *chartis* = map and *graphein* = write

a) Anaximander (611-547 BC)

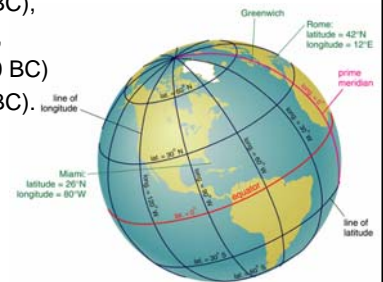
- Pupil of Thales
- First map maker
- Invented Longitude & Latitude
- Earth is a stone cylindrical column, we live on the top of it



2b. Geographic Coordinates 10

Spherical coordinates probably first done by

- Pythagoras (470 BC),
- Aristotle (330 BC),
- Eratosthenes (240 BC)
- Hipparchus (120 BC).



2c. Latitude 11

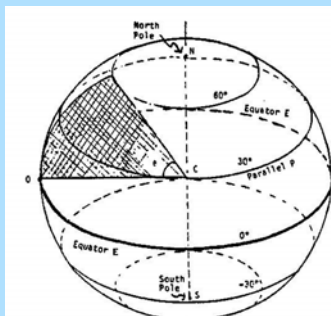
is measured up from equator

Hayward is 37°N40'08"

Equator is 0°

N pole is 90°

What is latitude of South Pole?



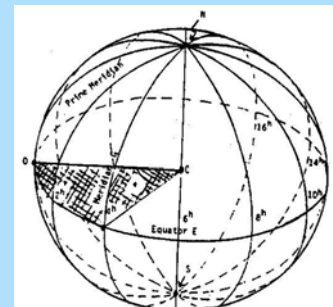
2d. Longitude 12

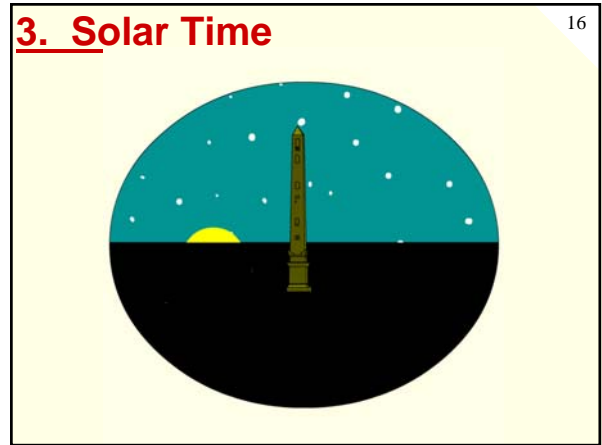
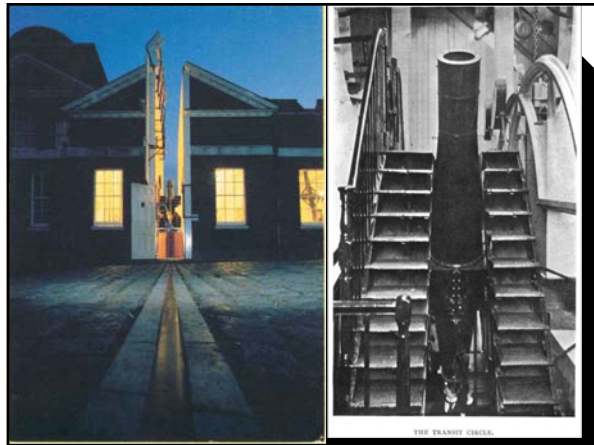
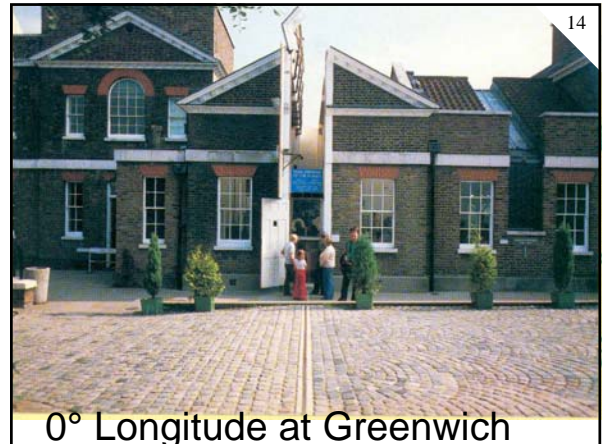
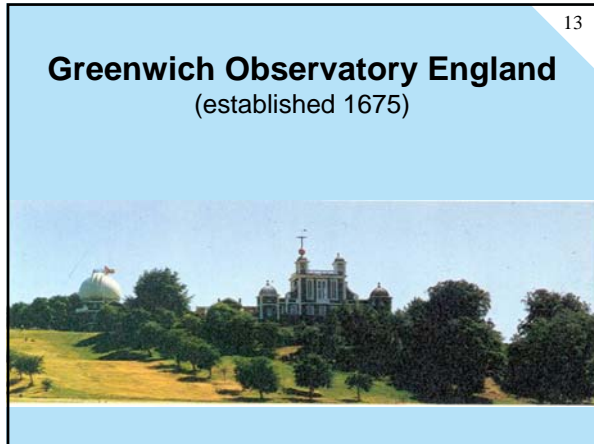
is measured westward from Greenwich, England

Hayward is 122°W04'51"

Greenwich is 0°

What is the longitude of the north pole?





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3. Sundial measures "Solar Time"

- **Gnomon** invented by Anaximander (611 – 547 BC). Shadow tells time of sun (first sundial)

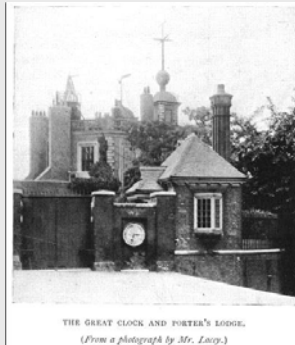
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3a) Sundials

- **340 BC Berosus Sundial**
Shaped like a bowl. The "longitude" line on the bowl tells how many hours before or after noon the sun is.
- **The transit of the sun** (when sun crosses local prime meridian) happens at **NOON** (12 hours) local time.

3b) UT: Universal Time 19

GMT: Greenwich mean time, the solar time at Greenwich Observatory

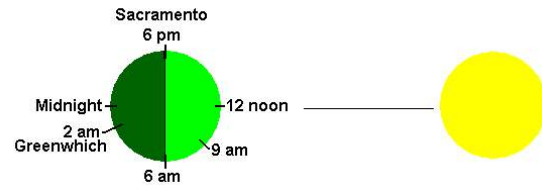


[nearly the same as UT: "Universal Time" and ET: "Ephemeris Time"]

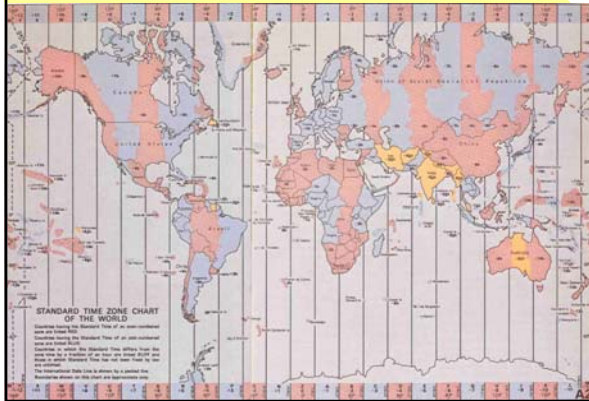
3b) Time is Relative 20

When its 6 pm here

- What time is it in Boston? **9 pm**
- What time is it in Greenwich? **2 am**
- Definition of time depends upon your longitude



3b) Time Zones are 15° (1 hr) wide 21



California is in Pacific Zone 22

- Pacific Standard time: PST
- PST = GMT - 8 hours
- Pacific Daylight Time: PDT = PST+1

Zone #	Name	Central Longitude	Symbol
5	Eastern	75	EST
6	Central	90	CST
7	Mountain	105	MST
8	Pacific	120	PST

3c. Local Mean Time 23

- At Lake Tahoe (120 degrees longitude) sun transits on average at 12:00 noon PST.
- At Hayward, sun transits on average 12:08 PST. This is because we are 2 degrees longitude west of Lake Tahoe. It takes 8 minutes for the sun to travel from Lake Tahoe to here.

- Local Mean Time: LMT = PST - 8 minutes
- Sun Transits (on average) at 12:00 LMT



3c. Longitude 24

Measure longitude in terms of time

- If 360° is equal to 24 hour
 - 15° is equal to 1 hour
 - 1° is equal to 4 minutes
 - 1' is equal to 4 seconds
- Hayward's Longitude is hence
 - 8 hours, 8 minutes, 19.4 seconds
 - West of Greenwich, England

3c) Relate Time & Longitude 25

- **LMT**: Local mean time or sundial time the time at our location
- **LMT=GMT - Longitude**
 - Hayward's Longitude is $8^{\text{h}}8^{\text{m}}$
 - $\text{LMT} = \text{GMT} - 8^{\text{h}}8^{\text{m}}$
 - $\text{LMT} = \text{PST} - 8^{\text{m}}$

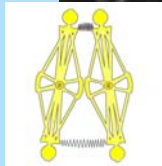
3d) Navigation 26

- To accurately determine longitude
- Measure Local Time (from stars)
- Compare with Greenwich time
- Need a good clock!
- Prize of 20,000 pounds!

John Harrison (1693-1776) 27



1730-1735 Maritime Clock



1737-1740 Maritime Clock #H2 28

In 1740 Harrison realized its design was wrong. The bar balances did not always counter the motion of a ship



1740-1759 Maritime Clock #H3 29

New design got rid of pendulums in favor torsion pendulum.

1. a bimetallic strip, to compensate the balance spring for the effects of changes in temperature
2. a caged roller bearing, the ultimate version of his anti-friction devices.



(1755-1759) Maritime Clock #H4 30

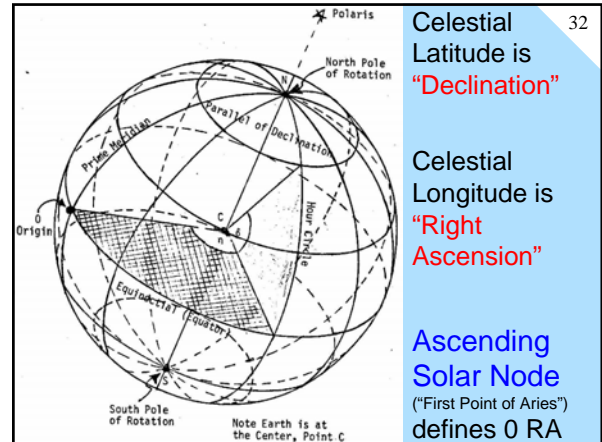
watch's error was computed to be 39.2 seconds over a voyage of 47 days, three times better than required to win the £20,000 longitude prize



5 inches in diameter

B1. Celestial Coordinates 31

- Anaximander (580 BC) invents idea of celestial sphere. (?)
- Eudoxus (360 BC) makes early map of constellations
- Hipparchus (130 BC) made a star catalog of 850 stars with some sort of coordinates
- Claudius **Ptolemy** (150 A.D.?): The first really accurate map, 48 constellations, 1025 stars with measured ecliptic longitude & latitude

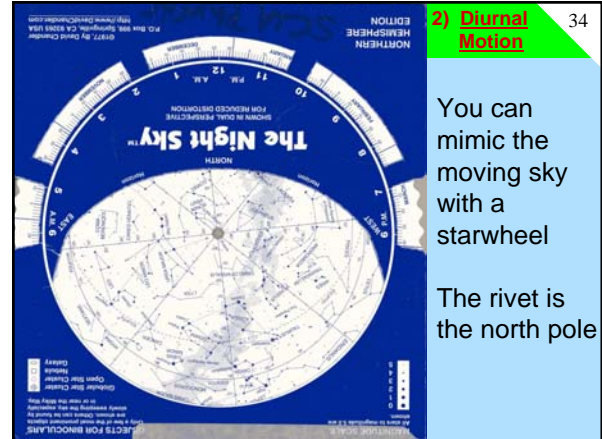
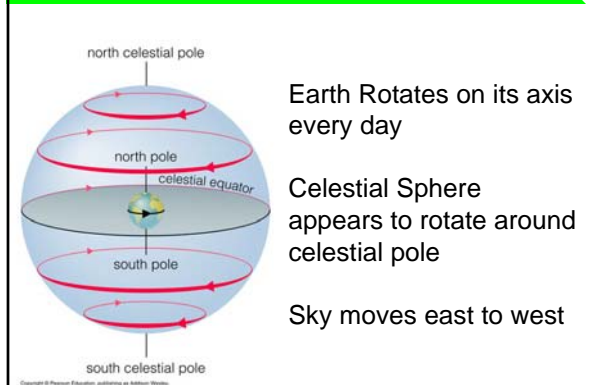


Celestial Latitude is "Declination"

Celestial Longitude is "Right Ascension"

Ascending Solar Node ("First Point of Aries") defines 0 RA

B2) Diurnal Motion 33



B3a). Star Time (sidereal) 35

The RA (right ascension) which is transiting (crossing the **prime meridian**) is "sidereal time"

Example:
At 6^h Sidereal time, Betelgeuse (Orion) is transiting



B3a). Sidereal vs Solar Time 36

- **Sundial (Solar) Time** tells the position of the sun (noon or 12:00 = transit)
- **Sidereal (Star) Time:** tells which stars are transiting
- Example, January 5th at 3 am, the sidereal time is 10 hours (i.e. the star Regulus with RA=10 hours, is transiting)

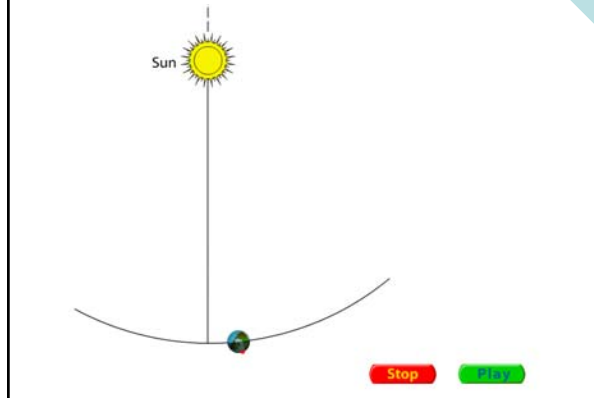
B3b. Sidereal Period

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- **Solar day: 24 hours**
time between transits of sun
- **Sidereal day: 23 hours 56 min**
time between transits of a star
- **This means a star will transit**
 - 4 minutes earlier each day
 - 2 hours earlier each month

B3b. Sidereal Period (animation)

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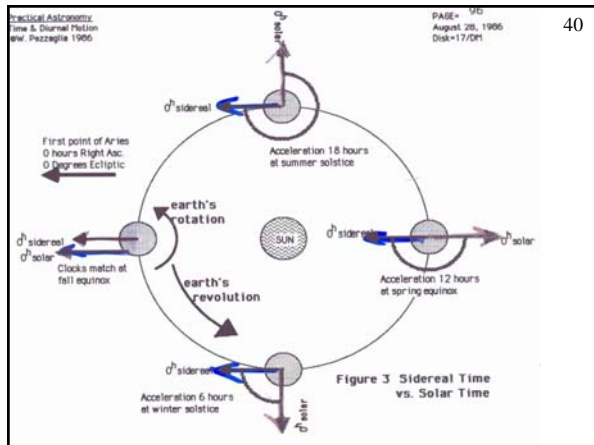
B3c). Acceleration of Sidereal Clock

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Your starwheel converts solar time
(plus date) to sidereal time!

LST: Local Sidereal Time

- **LST=LMT at fall equinox**
- **LST= LMT + “acceleration”**
 - Acceleration increases 4 min per day
 - Or 2 hours per month



Acceleration of Sidereal Clock

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- The amount by which the sidereal time is ahead of the solar time
- It's a function of date

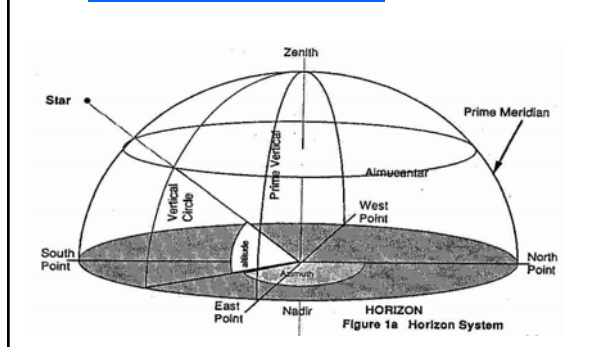
DATE	Ecliptic Long. of Sun	Acceleration* of Sidereal Clock	DATE	Ecliptic Long. of Sun	Acceleration* of Sidereal Clock
Sept 21	180°	0h	Mar. 22	0°	12h
Oct. 21	210°	2h	Apr. 19	30°	14h
Nov. 21	240°	4h	May 21	60°	16h
Dec. 21	270°	6h	June 21	90°	18h
Jan. 21	300°	8h	July 21	120°	20h
Feb. 21	330°	10h	Aug. 22	150°	22h

*Equivalently the hour angle of the First Point of Aries at 0 hours local mean time.

C. Horizon Coordinates

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1. Altitude and Azimuth



Your starwheel is a device for predicting the altitude and azimuth of stars, from the date and time

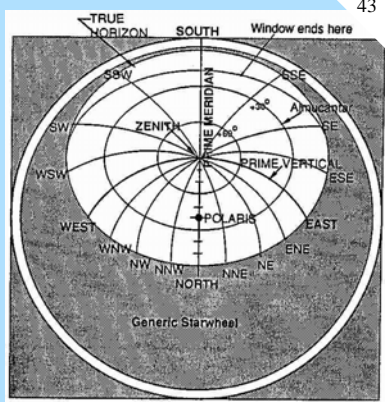


Figure 1b Horizon System on a Starwheel

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Arabian Astronomers perfected the "Astrolabe", the ancestor of your starwheel (more properly called a "planisphere")

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2. Relate Horizon to Equatorial 45

Declination of zenith is your latitude
 Declination at North point is $90 - \text{Latitude}$ (52°)
 Declination at South point is $\text{Latitude} - 90$ (-52°)

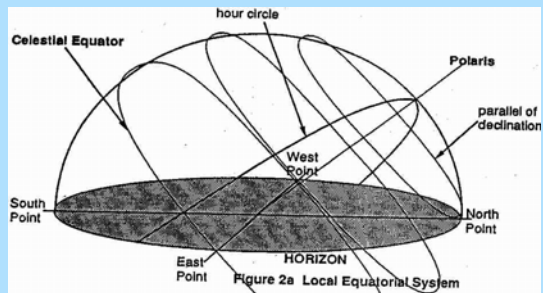
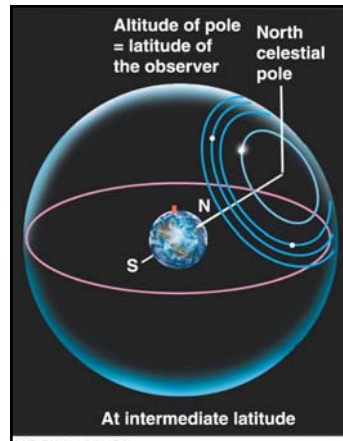


Figure 2a Local Equatorial System



At intermediate latitude

Stars inside the "circle of perpetual occultation" are called "circumpolar", and never set!

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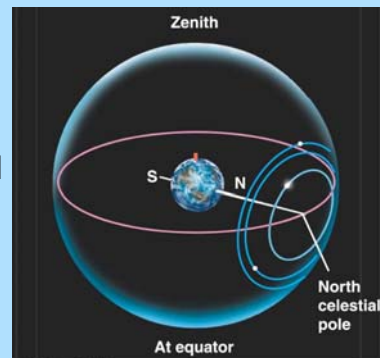


At North Pole

If you lived at the north pole, all stars would be circumpolar!
 Polaris would be at zenith

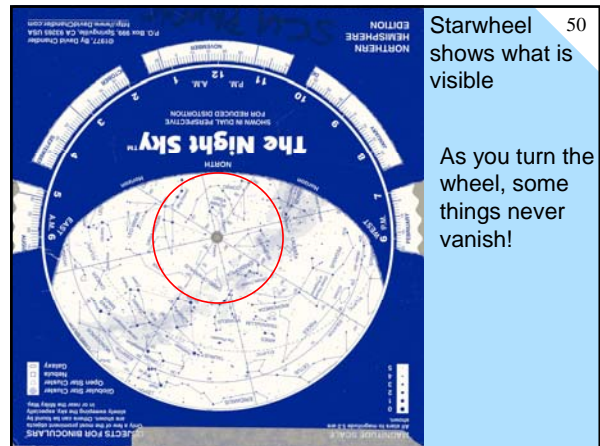
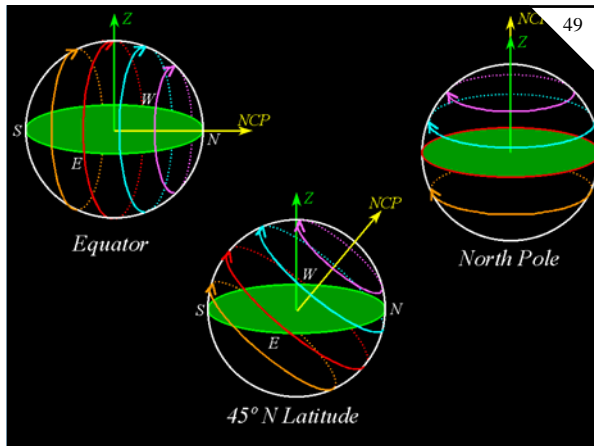
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If you lived at the equator, no stars would be circumpolar!
 Polaris would be at horizon



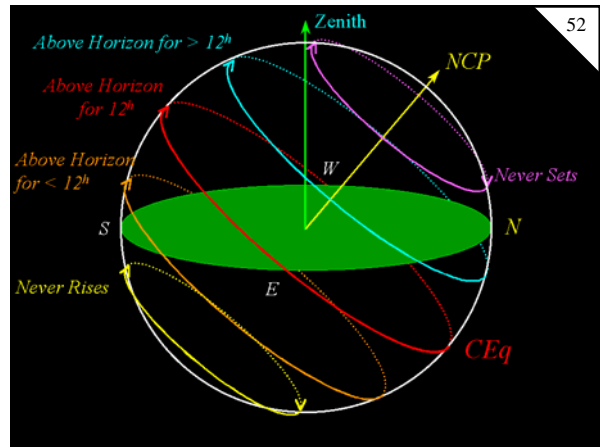
At equator

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3. Path of a star

Will follow a parallel of declination
 Transit altitude is $90^\circ - |\text{dec} - \text{lat}|$
 Rising/setting points will be same each day!



References

- <http://members.optusnet.com.au/~gtosiris/page11-9a.html>
- Note: the size of a "stadia" was different for Aristotle than Erastotenes. See
- <http://www.eso.org/public/outreach/eduoff/aol/market/collaboration/erastostenes/>

Things to Do

- Make a movie out of the Wolfram map of sky, and use it in title sequence?